# Observational- and modeling-based budget of lightning-produced $NO_x$ in a continental thunderstorm

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[1] NO<sub>x</sub> transport and production by lightning for the 10 July 1996 Stratosphere-Troposphere Experiment-Radiation, Aerosols, and Ozone convective storm is examined using radar, in situ observations and cloud model simulations. Observations and model simulations indicate that most of the NO<sub>x</sub> produced by the storm was transported out into the anvil. The analyzed NO<sub>x</sub> flux into the anvil, combined with results from a cloud model simulation, indicate that approximately 60 percent of the NO<sub>x</sub> transported into the anvil during the observational period is produced by lightning. Lightning flash rate and channel length measurements, obtained using the Office National d'Etudes et de Recherche Arospatiales lightning interferometer, are combined with the NO<sub>x</sub> budget to give estimates of average lightning NO<sub>x</sub> production per interferometer flash and per unit flash channel length. The analysis yields production rates of approximately 43.2 moles ( $2.6 \times 10^{25}$ molecules) NO<sub>x</sub> per interferometer flash and  $1.7 \times 10^{-3}$  moles (1.0 ×  $10^{21}$  molecules) NO<sub>x</sub> per meter of flash channel. These production rates fall within the bounds of rates derived in previous studies using completely different approaches. INDEX TERMS: 0320 Atmospheric Composition and Structure: Cloud physics and chemistry; 0368 Atmospheric Composition and Structure: Troposphere—constituent transport and chemistry; 3314 Meteorology and Atmospheric Dynamics: Convective processes; 3324 Meteorology and Atmospheric Dynamics: Lightning; KEYWORDS: lightning, NOx, convection, transport

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#### 1. Introduction

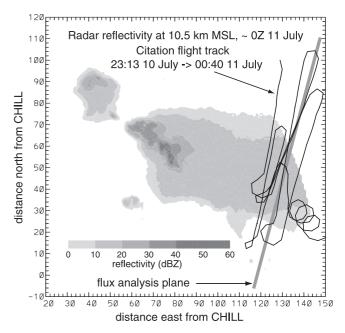
[2] Nitrogen oxides  $(NO_x = NO + NO_2)$  are important reactive species in the photochemistry of the troposphere [Bradshaw et al., 2000]. Currently, production of NO<sub>x</sub> by lightning represents a large uncertainty in the global NO<sub>x</sub> budget with estimates varying by up to two orders of magnitude, although the range is likely within 2-20 Tg(N)/yr [Liaw et al., 1990; World Meteorological Organization (WMO), 1995; Huntrieser et al., 1998]. The uncertainty is amplified for global models of tropospheric processes because lightning-produced NO<sub>x</sub> can be a direct source to the middle and upper troposphere where its lifetime is considerably longer than in the boundary layer. Estimates of the global budget of lightning-produced NO<sub>x</sub> (LNO<sub>x</sub>) rely on a number of assumptions that lead to this large uncertainty [e.g., Price et al., 1997; Lawrence et al., 1995]. For example, there is an incomplete understanding of the physics of NO<sub>x</sub> production by the different components of a lightning flash and uncertainty associated with quantitatively characterizing the lightning discharge energy and channel length in convective storms. Further uncertainty arises in the extrapolation of lightning flash characteristics and storm production rates to the global LNO<sub>x</sub> budget.

- [3] As a step toward quantifying and reducing the uncertainty in the LNO<sub>x</sub> budgets, the assumptions and parameters used in these models should be compared with direct observations where possible. Toward this end, we present an NO<sub>x</sub> budget for a continental thunderstorm observed on 10 July 1996 during the Stratospheric-Tropospheric Experiment: Radiation, Aerosols, and Ozone (STERAO)/Deep Convection project. In this study we use detailed measurements of lightning flash rates and channel lengths of lightning components to examine NO<sub>x</sub> production by lightning. The budget includes estimates of the time varying lightning production rate of NO<sub>x</sub>, total lightning NO<sub>x</sub> production over the lifetime of the storm, and estimates of the average NO<sub>x</sub> production per flash and per meter flash channel length during the storm. These budget products can be directly compared with parameters used in global LNO<sub>x</sub> budgets and with estimates from other observational, laboratory and theoretical studies.
- [4] The budget is produced by combining meteorological and chemical observations from aircraft in the anvil outflow (section 2), an analysis of NO<sub>x</sub> transport in a 3D cloud-

model simulation of the storm (also in section 2), and measured lightning flash rates and estimated lightning channel lengths using data from the ONERA lightning interferometer (section 3). Extrapolation of these results to the global scale is considered in the last part of section 3, followed by a summary and discussion in section 4.

#### 2. NO<sub>x</sub> Flux Analysis

- [5] The 10 July STERAO storm developed along the southern portion of the border between Wyoming and Nebraska over a locally elevated topographic feature known as the Cheyenne ridge. Through most of its lifetime (from approximately 2200 GMT 10 July to 0100 GMT 11 July, or 1600 to 1900 10 July local time) the convective system consisted of multiple cells aligned along a NW-SE axis. Storm inflow at low levels was from the south and west, and anvil outflow was to the SE in the lower and middle parts of the anvil (from 6.5 to 10.5 km MSL) and toward the east above approximately 12.5 km MSL. The orientation and growth of the anvil are dictated by the upper level winds, which were northwesterlies turning to westerlies above 12.5 km, and by the slow storm movement relative to these upper level winds (approximately 5 m/s toward the SSE [Skamarock et al., 2000]. After 0100 GMT 11 July, the storm consisted of a single strong cell (with supercell characteristics) that propagated at a somewhat higher speed to the south.
- [6] Extensive observations of the storm were obtained from a variety of platforms [see Dye et al., 2000]. Of particular interest for this study, NO, CO and ozone measurements were taken on the North Dakota Citation aircraft [Stith et al., 1999; Dye et al., 2000] while it flew through the anvil of the storm. Figure 1 shows the projection of the Citation flight track between 2313 10 July and 0040 GMT 11 July on the horizontal plot of radar reflectivity (at 00 GMT 11 July) at 10.5 km MSL. The Citation aircraft mapped out the anvil structure during this time period by traversing the anvil in horizontal passes, approximately perpendicular to the long axis of the anvil, at elevations starting at approximately 11.8 km MSL (close to the anvil top) and ending at approximately 6.8 km MSL. Both satellite photos and radar reflectivity [see Dye et al., 2000; Skamarock et al., 2000, Figure 5] suggest that the evolution of the storm from a line of convective cells to a single supercell occurred somewhat after the period in which the Citation observations in the anvil were taken. Based on the mean upper level winds, a parcel detraining from the core into the anvil would take around 45 minutes to reach the area in which the Citation was flying. Thus the aircraft was sampling anvil air from the multicellular period of the storm, detrained from the updrafts from between approximately 2230 GMT 10 July to 00 GMT 11 July.
- [7] Visible satellite images of the storm at 2315 10 July and 0015 11 July (Figure 2) suggest that the primary changes in the storm between these times are growth of the anvil to the southeast and east along with a propagation of the convective line to the south. Radar reflectivity plots (see Figure 1 and *Skamarock et al.* [2000]) show that between 2 to 4 cells were active in the convective line at any time during this period. Given that the area of active convection does not grow significantly relative to the



**Figure 1.** Horizontal cross section of radar reflectivity at 10.5 km MSL. The analysis plane is shown in the thick gray line and the black line is the aircraft flight track.

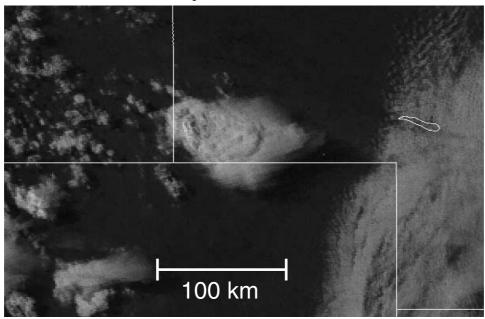
growth of the anvil area, it is likely that the growth of the anvil represents the net convective detrainment during this time period. Observations of NO, CO and O<sub>3</sub> taken on the Citation aircraft allow us to compute these fluxes (the convective detrainment) into the growing anvil.

[8] Because the Citation obtained measurements of NO and not NO<sub>x</sub>, we perform the flux analysis on measured NO mixing ratios and then determine the flux of NO<sub>x</sub> using the photochemical steady state assumption. For this assumption we consider only the NO, NO<sub>2</sub>, and O<sub>3</sub> chemistry such that the NO<sub>2</sub> to NO ratio is  $[NO_2]/[NO] = k_2 [O_3]/j_1$ , where  $j_1$  is the photolysis frequency of NO2 which produces NO and  $O(^{3}P)$ , and  $k_{2}$  is the rate constant of the NO +  $O_{3}$  reaction. We calculate the NO<sub>2</sub> to NO ratio to be 0.3 for clear sky conditions in the upper troposphere. In the upper part of the anvil where photolysis frequencies are enhanced by cloud scattering by a factor of two or so, the NO<sub>2</sub> to NO ratio could be half the clear sky value, while in the optically thick lower portion of the anvil the NO<sub>2</sub> to NO ratio would be larger. In addition, departure from the photochemical steady state due to peroxy radicals would also increase the NO<sub>2</sub> to NO ratio. However, lightning-produced NO in the anvil will decrease the NO2 to NO ratio. Thus, we estimate that the NO<sub>x</sub> to NO ratio could range from 1.1 to 1.6 [Dye et al., 2000], and we use a constant value of 1.3 in the remainder of the analysis.

#### 2.1. Control Volume Analysis

[9] We perform a flux analysis for the 10 July storm in order to quantify the net detrainment of air mass,  $NO_x$  and CO, with the goal of relating the  $NO_x$  flux to lightning  $NO_x$  production. The flux analysis can be understood by considering the 2D depiction of a generic storm (Figure 3) in which a control volume is drawn around the storm. The time rate of change of  $NO_x$  mass in the control volume will be the difference between the inflow and outflow through the

## 2315 UTC, 10 July 1996



## 0015 UTC, 11 July 1996

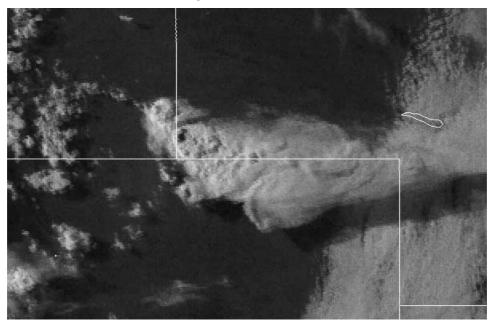


Figure 2. Visible satellite images of the storm at 2315 GMT 10 July 1996 and 0015 GMT 11 July 1996.

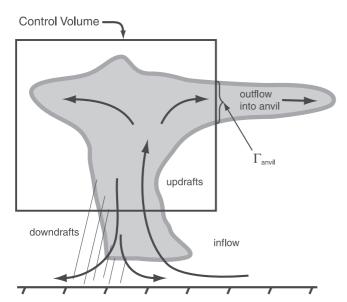
control volume boundary plus any sources within the volume. This is quantitatively stated by the conservation equation

$$\int_{V} \frac{\partial \rho \phi_{NO_{x}}}{\partial t} = \int_{\Gamma} \left( \rho \mathbf{v} \cdot \mathbf{n} \, \phi_{NO_{x}} \right) + S_{NO_{x}}, \tag{1}$$

where  $\rho$  is the dry air density,  $\phi_{NOx}$  is the  $NO_x$  mixing ratio with respect to the dry air, V is the control volume,  $\Gamma$  is the control volume boundary, V is the velocity, V is the unit

normal to the boundary  $\Gamma$ , and  $S_{NOx}$  is an  $NO_x$  source. All the terms are time-dependent.

[10] Generally, all the terms in equation (1) can be significant. The Citation aircraft measured only part of the first term on the RHS of (1), that being the flux of  $NO_x$  out into the growing anvil. We have no measurements of the flux of  $NO_x$  from the boundary layer into the storm or downdraft fluxes of  $NO_x$  out of the storm (i.e., other major contributions to the second term in (1)). We also have no direct measurements of  $NO_x$  production by lightning in the



**Figure 3.** Two-dimensional conceptualization of a storm along with the control volume used for the flux analysis.

storm (the source term), and we have no direct measurements of the change of  $NO_x$  storage inside the control volume (within the storm, the leftmost term).

- [11] For the 10 July storm, indirect evidence strongly suggests that the leftmost term in (1) is small relative to the other terms in the equation and that most of the  $NO_x$  produced by lightning is ultimately detrained into the anvil. The arguments and evidence that support these assumptions are as follows:
- [12] 1. As noted previously, the area of active convection did not change significantly in size relative to the change in size of the anvil over the period of the Citation anvil traverses. This is supported by visible imagery (Figure 2), and the radar reflectivity [see *Dye et al.*, 2000; *Skamarock et al.*, 2000].
- [13] 2. There is little evidence suggesting that convective downdrafts, which were generally weak, were transporting significant quantities of NO<sub>x</sub> (either lightning produced NO<sub>x</sub> or NO<sub>x</sub> previously transported upward from the PBL by convection). Specifically, the NOAA P3 aircraft saw little enhancement in NO below the anvil or at lower levels in the vicinity of the storm, and lightning was observed to occur mostly within moderate updrafts, and not within strong downdrafts [*Dye et al.*, 2000]. Additionally, modeling results suggest that downdrafts reaching the boundary layer are likely to have their origins below approximately 5.5 km MSL [*Skamarock et al.*, 2000]; this is below the level where most of the lightning was observed in this storm
- [14] 3. Comparison of lightning locations with Doppler radar derived fields of vertical velocity showed lightning occurring mostly in regions of moderate updraft [*Dye et al.*, 2000]
- [15] 4. The lightning was almost all intercloud (IC) [Dye et al., 2000; Defer et al., 2001] and the strong storm-relative upper level winds should advect most of the lightning-produced  $NO_x$  through the anvil.
- [16] These arguments and evidence suggest that it is reasonable to dispense with the storage term in (1) and to

assume that all the  $NO_x$  produced by lightning is ultimately detrained into the anvil. These two assumptions allow us to focus the flux analysis on the flux into the anvil through the boundary subregion  $\Gamma_{anvil}$  in Figure 3. The assumption that all lightning-produced  $NO_x$  fluxes out into the anvil can be stated as

$$S_{NO_x} = -\int_{\Gamma_{amvil}} \left( \rho V \cdot \mathbf{n} \phi_{NO_x}^S \right), \tag{2}$$

where  $\phi_{NOx}^{S}$  is the lightning-generated NO<sub>x</sub>. The Citation aircraft observations, however, measure the total flux of NO out into the growing anvil, allowing us to estimate the NO<sub>x</sub> flux, i.e.,

$$-\int_{\Gamma_{amil}} \rho V \cdot \mathbf{n} \left( \phi_{NO_x}^E + \phi_{NO_x}^S \right), \tag{3}$$

where  $\phi_{NOx}^{E}$  is the pre-existing (environmental) NO<sub>x</sub>. We cannot distinguish between lightning produced NOx and, say, that transported upward from the boundary layer by updrafts or entrained directly into the anvil flow from the environment. We must disentangle the two fluxes in (3) in order to compute the NO<sub>x</sub> production rate from lightning,  $S_{NOx}$  in (2). We will use results from a cloud model simulation of the 10 July storm to estimate the amount of NO<sub>x</sub> transported into the anvil that accounts for all the flux except that of lightning-produced NO<sub>x</sub>. The remaining NO<sub>x</sub> flux through the anvil can be attributed to the lightning produced NO<sub>x</sub> and thus is an estimate of the lightning NO<sub>x</sub> production  $S_{NOx}$  in (2). In the remainder of this section we analyze the total flux using the observations from the Citation, estimate the flux of pre-existing  $NO_x$  ( $\phi_{NOx}^E$ ) from the cloud model simulation, and combine the two fluxes to extract the lightning  $NO_x$  production rate  $S_{NOx}$ .

#### 2.2. Objective Analysis of Aircraft Observations

- [17] Figure 4a depicts the Citation aircraft track projected perpendicularly onto the vertical flux analysis plane (whose location is shown in Figure 1). The analysis plane moves with the storm at the system velocity (u, v) = (1.5, -5.5) m/s where u is positive toward the east and v is positive toward the north [see Skamarock et al., 2000]. Observations from the Citation aircraft are averaged over 30 sec intervals and likewise projected onto the analysis plane, and include the horizontal winds (u,v), NO, CO, O3 and cloud particle concentration from the Particle Measuring Systems 2D probe [see Dye et al., 2000]. The location of the midpoints of the 30 sec averages are shown as circles along the flight track in Figure 4a. The CO data is necessarily time averaged because of the response characteristics of the CO instrument [see Dye et al., 2000], the other observations are averaged in the same way for consistency. In the analysis that follows, the time-averaging does not significantly affect on the results.
- [18] For calculating fluxes, the projected observations are analyzed onto a regular grid ( $\Delta x = 2000 \text{ m}$ ,  $\Delta z = 200 \text{ m}$ ) using an objective analysis procedure similar to that developed by *Cressman* [1959]. In this application we represent a variable as the sum of a reference state and a perturbation,  $\phi = \phi_o(z) + \phi'$ , where the reference state represents the undisturbed (pre-convective) environment. On the regular grid at

### Citation flight 10 July 1996, 23:16:00 10 July to 00:36:40 11 July

vertical cross section, (ud,vd) = (1.5,-5.5) m/s, (xm,ym) = (130,65) km from CHILL, rotated  $15^{\circ}$ 

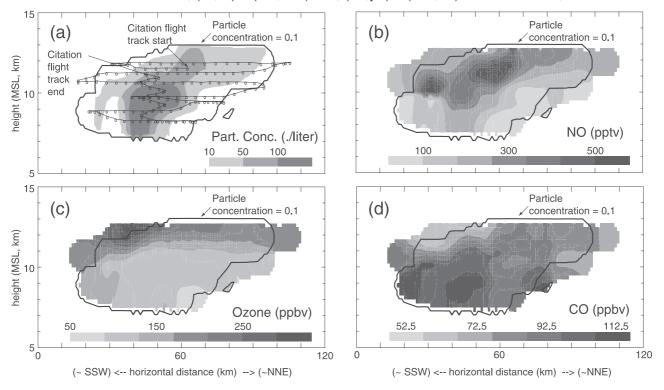


Figure 4. Objective analysis of (a) cloud particle concentration, (b) NO, (c) O<sub>3</sub>, and (d) CO using the Citation flight data. The flight track is shown in the cloud particle concentration plot. See color version of this figure at back of this issue.

points  $(z_a, x_a)$  we calculate the objectively analyzed value  $\phi_a$ =  $\phi_o(z) + \phi'_a$  by computing the perturbation value  $\phi'_a$  using

$$\phi'_{a} = \frac{\sum_{i=1}^{K} w_{i}(\phi_{obs} - \phi_{o}(z))}{\sum_{i=1}^{K} w_{i}},$$

where

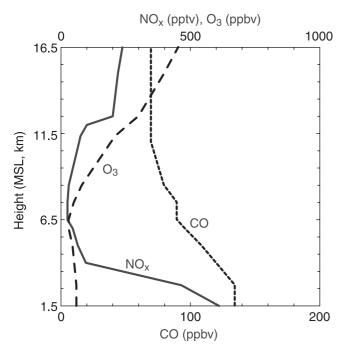
$$w_i = \frac{1 - d^2}{1 + d^2}, \text{ if } d^2 < 1,$$
  
 $w_i = \frac{1 - d^2}{1 + d^2}, \text{ if } d^2 \ge 1,$ 

and

$$d^2 = \left(\frac{x_a - x_{obs}}{x_r}\right)^2 + \left(\frac{z_a - z_{obs}}{z_r}\right)^2.$$

[19] The observed values are denoted as  $\phi_{obs}(z_{obs}, x_{obs})$ , and  $x_r$  and  $z_r$  define the radius of influence of the observation on the analyzed values, which in our analyses are  $x_r = 5$  km and  $z_r = 1.5$  km. The results of the following flux analysis are not sensitive to values of  $x_r$  and  $z_r$ , and the primary constraint on the values are that  $z_r$  be significantly greater than half the height difference between flight levels (>1/2 km), and  $x_r$  be significantly greater than half the horizontal distance between time-averaging locations. Neither are the results sensitive to the analysis grid lengths  $\Delta x$  and  $\Delta z$ . The results are also relatively insensitive to the choice of reference states; the reference states from Skamarock et al. [2000] are employed here for the horizontal winds (u,v), CO and O<sub>3</sub>. The reference state for NO (and CO and O<sub>3</sub>) can be found in Figure 5, and the background cloud particle concentration is zero.

- [20] Figure 4 presents the results from the objective analyses for NO, CO and O3, along with cloud particle concentrations from the 2D probe. The data are contoured only where observations have contributed to the values. The analyzed fields extend somewhat more than a km above and below the highest and lowest flight track observations. We have no data in these regions and the values represent an extrapolation.
- [21] The fields possess much structure. First, the tropopause is clearly depicted in the O<sub>3</sub> field, where a significant dip in the tropopause is evident at the top southwest edge of the anvil. This signature is also evident in the CO analysis. The structure can differ substantially between fields. For example, the NO analysis shows isolated maxima located above 10 km MSL while the CO analysis shows somewhat less-isolated maxima located below 10 km MSL. Two factors likely contribute to the different structures; the preconvective vertical distributions of NO and CO differ (higher CO levels extend well beyond the PBL top compared to NO, see Figure 5), and NO is also produced by lightning. The two maxima of NO >500 pptv are likely major out-flux areas of LNO<sub>x</sub>.
- [22] To compute fluxes of mass and other species through the anvil on the analysis plane, we use the plane-perpendicular velocities analyzed from the aircraft data (not shown) along with the analyzed NO and CO and a reference density profile. We define the anvil area as the area within the 0.1



**Figure 5.** Environmental profile of NO<sub>x</sub>, CO and O<sub>3</sub> used in the cloud model simulation.

particles per liter contour in the particle concentration analysis in Figure 4a. The resulting fluxes are given in Table 1. Next, we use a cloud model simulation that includes passive tracer transport to partition the  $NO_x$  fluxes in (3).

#### 2.3. Cloud Model Simulation

[23] A three-dimensional cloud model simulation of the 10 July storm is described by *Skamarock et al.* [2000]. The simulation, performed using a horizontally homogeneous environment and thermodynamic and wind soundings composited from the STERAO observations, generally reproduced the observed structure and evolution of the storm. A comparison of the simulated convective system with observations and a detailed analysis of CO and  $O_3$  transport in the storm was performed by *Skamarock et al.* [2000]. To produce the flux analysis, we have repeated the simulation using a larger horizontal domain to better capture the growing anvil (160  $\times$  160 km), and we have included

another passive tracer representing NO<sub>x</sub> in addition to CO and O<sub>3</sub>. There is no lightning source of NO<sub>x</sub> in the model, hence the NO<sub>x</sub> in the model represents NO<sub>x</sub> existing in the pre-storm environment. The environmental CO profile is given by Skamarock et al. [2000] and is shown in Figure 5. The environmental NO<sub>x</sub> profile, also shown in Figure 5, is derived from aircraft observations and climatology. NO measurements obtained from the WP3D aircraft were averaged at model levels up to 6.5 km MSL. The standard deviation of these averages was approximately 40% of the mean value. The NO<sub>x</sub> mixing ratios were computed by assuming a photochemical steady state for O<sub>3</sub>, NO, and NO<sub>2</sub> chemistry (again peroxy radicals were not included in the photochemical steady state calculations). The NO<sub>2</sub> photolysis rate used in the calculation of the photochemical steady state was determined from radiometric measurements made on the WP3D [Frost et al., 1998]. The NO<sub>x</sub> mixing ratios for the 6.5-14 km MSL altitudes were obtained from the UND Citation out-of-cloud NO measurements averaged at model levels and adjusted to NO<sub>x</sub> values using the photochemical steady-state assumption. Above 14 km MSL, NO<sub>x</sub> mixing ratios were determined by using NO<sub>x</sub> data reported in the literature [Salawitch et al., 1994] and mean July NO<sub>x</sub> mixing ratios as calculated by the global chemistry transport model MOZART [Brasseur et al., 1998] for the grid point (104.1°W, 40.5°N) nearest to northeastern Colorado.

[24] Figure 6 shows a horizontal cross section of radar reflectivity derived from the model output at 6000 sec in the simulation, and can be compared with Figure 1. The position of a vertically oriented analysis plane, comparable to that used in the analysis of the aircraft observations, is also shown in Figure 6. The location of this plane is chosen such that it is a similar distance downwind of the convective cells compared with the plane used for the observational analysis.

[25] Figure 7 shows an outline of the model cloud field and  $NO_x$ , and it can be compared with the observational analysis of NO in Figure 4. Several differences are obvious in the comparison. First, the simulated anvil cross-sectional area is significantly smaller than that observed, averaging less than two thirds the observed area through most of the simulation. We attribute this to the fact that the simulated convection produced a somewhat smaller system than that observed; convective cells were triggered or regenerated somewhat more readily in the observed storm than in the simulation. Second, the simulation does not exhibit the pronounced

Table 1. Flux Analysis for the Observations and Simulation

	Cross Section Area, 10 <sup>6</sup> m <sup>2</sup>	Dry Air Mass Flux Density, kg/s/m <sup>2</sup>	CO Flux Density, 10 <sup>-5</sup> moles/s/m <sup>2</sup>	$NO_x$ Flux Density, $10^{-8}$ moles/s/m <sup>2</sup>
Observations (Citation data)	315	5.9	1.9	$4.5 \text{ (NO) } 5.8 \text{ (NO}_{x})$ $(\text{NO}_{x} = 1.3*\text{NO})$
Simulation				
Time (s)				
3600	109	6.7	1.9	2.4
4200	177	6.6	1.9	2.2
4800	193	6.5	1.9	2.2
5400	232	6.6	1.9	2.3
6000	238	6.5	1.9	2.2
6600	187	6.6	2.0	1.9
7200	119	6.5	2.0	2.0
Average				
of the 7 times	180	6.6	1.9	2.2

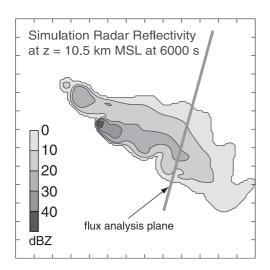


Figure 6. Horizontal cross section of radar reflectivity from the cloud model simulation at 6000 sec. The location of the analysis plane used for the simulation is also included.

peaks of NO<sub>x</sub> in the upper part of the anvil found in the observational analysis of NO. We attribute this to the presence of lightning-produced NO<sub>x</sub> in the observations.

[26] Fluxes of air mass, CO and NO<sub>x</sub> integrated over the anvil through the analysis plane have been computed. For quantitative comparison we have computed the flux densities (fluxes divided by the anvil cross-sectional areas) and these flux densities are presented in Table 1. The simulated fluxes are given at ten minute intervals between 1 hour and 2 hours.

This period corresponds to a time somewhat before the simulated storm transitions to a supercell. The observations used to compute the fluxes span a little over an hour starting at approximately 2315 GMT, somewhat before the observed storm transitioned to a supercell. We have chosen to sample the simulation so that the storm character during the period used for the simulation analysis corresponds to the observed storm's character for the time during which the aircraft data used in the observational flux analysis were taken.

[27] The mass and CO flux densities compare well, although the air mass flux density is consistently approximately 10% higher than the observed mass flux density. This difference could be attributed to errors in the initial wind profile or other errors in the convective simulation. The average normalized NO<sub>x</sub> fluxes differ dramatically. The observed flux density is approximately 2.5 times larger than the simulated flux density. From this we conclude that somewhat over 60 percent of the observed NO<sub>x</sub> flux out the anvil can be attributed to lightning-generated NO<sub>x</sub>. This results in an average of approximately 11.5 moles/s (6.9  $\times$  $10^{24}$  molecules/s) NO<sub>x</sub> produced by lightning (anvil area  $\times$  $NO_x$  flux per unit area  $\times$  0.62) transported through the anvil cross section during the observation period.

[28] There are a number of possible sources of error in this flux analysis. In the cloud model simulation, the distribution of NO<sub>x</sub> in the simulated anvil is moderately sensitive to the environmental profile (see Figure 5). Decreasing the NO<sub>x</sub> in the upper troposphere in the profile leads to a small increase in LNO<sub>x</sub> production (70% of the flux is attributed to LNO<sub>x</sub> if the upper level environmental NO<sub>x</sub> is specified as 50 pptv). Likewise decreasing the

### Vertical cross section from simulated 10 July storm at 6600 s

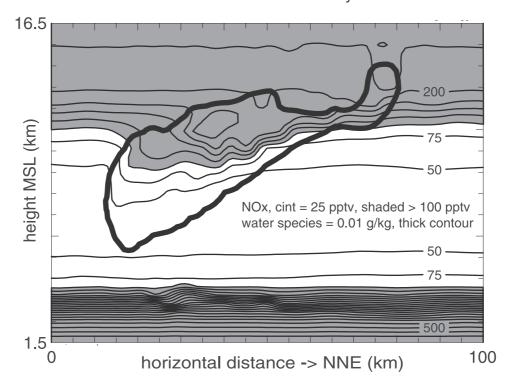


Figure 7. Vertical cross section of NO<sub>x</sub> and water species in a position similar to the analysis cross section.

boundary layer NO<sub>x</sub> leads to an increase in LNO<sub>x</sub> production. With respect to the observational analysis and apart from the assumptions discussed previously concerning the control volume and storm evolution, there is the issue of how to characterize the objective analysis of the aircraft data given that the storm was evolving over the time during which the anvil was mapped. Specifically, while the average of the fluxes from the simulation result in a definable time average, the observations from the anvil do not constitute a similar time average. The assumption we implicitly make in using the observations in this analysis is that this snapshot of the anvil structure is representative of the actual timeaveraged anvil structure. Radar reflectivity plots (not shown) suggest that the anvil cross-sectional area did not vary dramatically over the observational period-much less than the factor of 2 found in the simulation. Furthermore the values of NO, CO and O<sub>3</sub> measured during the spiral ascent at the end of the observation period are comparable to those from the previous horizontal pass at the different altitudes. The simulation also suggests that the flux densities do not vary much during the sampling period; the simulated flux densities of  $NO_x$ , CO and mass vary by less that  $\pm 10\%$  even though the anvil cross-sectional area varies by a factor of 2. These results suggest that our assumption of the representativeness of the anvil objective analysis may be reasonable.

#### 3. Lightning and NO<sub>x</sub> Production

[29] In the following section we produce estimates of the average LNO<sub>x</sub> production per unit channel length and per flash, and total storm LNO<sub>x</sub> production based on total lightning channel length produced by the storm (i.e., the sum of all lightning-component lengths) and total flashes, under the assumption that all lightning-produced NO<sub>x</sub> appears in the anvil outflow. Support for this assumption was outlined in the previous sections: The storm produced mostly IC lightning as opposed to CG lightning during the analysis period (2230-2400 GMT), the P3 aircraft did not find evidence of enhanced NO<sub>x</sub> values at low levels, and the multiple Doppler analyses did not show lightning occurring in pronounced downdrafts. Also, simulations we have performed with NO<sub>x</sub> sources used to simulate lightning production (distributed over a broad altitude and horizontal area) show LNO<sub>x</sub> being primarily transported out into the anvil.

[30] Estimates of the lightning channel lengths and flashes are derived from data obtained by the ONERA lightning interferometer. The interferometer was deployed to record the total lightning activity in the observed convective storms. The capabilities of the instrument and the analysis of the data for the 10 July storm are described by Defer et al. [2001], which includes a description of the methods used to determine lightning flashes. Flash reconstruction is accomplished by examining the time and spatial location of VHF pulses recorded by the interferometer. A VHF source is associated with a flash if the azimuth and elevation (as measured by the interferometer) of the VHF source is spatially and temporally close to those of the previous sources already considered part of the flash. For example, temporally, if the time between 2 sources was more than 500 ms, the second source is considered the start of a new flash. The determination of spatial closeness is somewhat more involved and we refer the reader to Defer et

al. [2001]. The methods used to estimate the length of the flash components are described by Defer et al. [2003]. We want to emphasize that our flash rates and channel lengths are different from those used in previous studies. The flash rate includes all flashes detected by the interferometer that had durations ranging from 23  $\mu$  sec (the shortest duration measurable by the interferometer) to over 1 sec. Likewise the total channel length is an estimate of the sum of lengths of the negative leaders, K changes and return strokes.

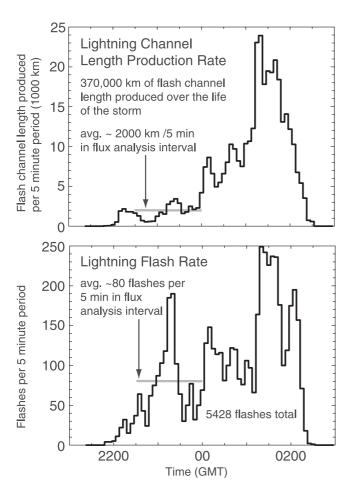
[31] In this study we assume that all the components radiating in the VHF domain take part in the LNO<sub>x</sub> production. Analysis of the VHF radiation during the period 2230–2400 indicated that most of the flash components were K change processes [*Defer et al.*, 2001], which are likely to have temperatures in excess of 20,000 K, similar to dart leaders [*Orville*, 1975] and well above the NO freeze-in temperature of 3500 K [*Chameides et al.*, 1977]. Very few negative leader processes were recorded during this period probably due to their weak signal at the lightning mapper stations.

#### 3.1. Lightning NO<sub>x</sub> Production Rate

[32] A flash-by-flash analysis has been performed in order to characterize the interferometer flash rate and flash durations for the entire storm [Defer et al., 2001]. Defer et al. [2003] applied simple relations to estimate the total length of all the flash components radiating in the VHF domain, relations deduced from the physics of lightning. The lightning component channel length production rate and the interferometer flash rate from this analysis are given in Figure 8 for the entire life of the storm. Both interferometer flash rate and component channel length production rate vary significantly during the storm. These two measures of lightning activity are not always well correlated, evidence the large peak at approximately 2320 UT in the interferometer flash rate that does not correspond to a particularly large peak in component channel length production. In fact, analysis of the flash population reveals that at about 2320 UT, half of the flashes had durations less than 1 ms [Defer et al., 2003]. Many of these short duration flashes emitted intense VHF radiation suggestive of large currents and temperatures. The flash lengths shown here are very different and much longer then those estimated or assumed in previous studies of lightning produced NO<sub>x</sub> [e.g., Price et al., 1997] where emphasis was on CG return strokes. Likewise the overall flash channel length production is orders of magnitude larger than estimates discussed in previous studies.

[33] The average lightning channel length production rate over the period 2230 GMT to 2400 GMT is approximately 6700 m/s (see Figure 8). Using the analyzed lightning production rate for  $NO_x$  given in the previous section, we estimate that the average amount of  $NO_x$  produced per meter path length over this period is approximately  $1.7\times10^{-3}$  moles/m  $NO_x$  (1.0  $\times$   $10^{21}$  molecules/m  $NO_x$ ).

[34] This estimate of  $NO_x$  production rate per meter channel length falls within other  $NO_x$  production rate estimates produced using entirely different approaches. *Stith et al.* [1999] analyzed  $NO_x$  peaks encountered along the flight track from the STERAO storms on 9, 10 and 12 July 1996. In that study, individual NO spikes (above a measured background) were fit to a simple model of lightning generated NO plumes. Using the model, they estimate that



**Figure 8.** Lightning channel length production and flash rate for the 10 July STERAO storm. Data from *Defer et al.* [2003].

between  $2 \times 10^{20}$  and  $10^{22}$  molecules of  $NO_x$  per meter were produced in the plumes that they sampled. The large range found by Stith et al. and also by Huntrieser et al. [2002] for a storm in Germany suggests that different channels of a lightning flash may produce varying amounts of NO<sub>x</sub>. Our estimate  $(1.0 \times 10^{21} \text{ molecules NO per meter})$ , which integrates the NO<sub>x</sub> over the storm outflow and uses estimates of total lightning channel length, is in the lower end of the Stith et al. [1999] range, suggesting that on average many channels are weak producers of  $NO_x$ . Wang et al. [1998] found between  $5 \times 10^{20}$  and  $3 \times 10^{21}$  molecules of  $NO_x$  per meter of flash channel were produced for lightning generated in the laboratory at sea level with currents between 10 to 30 kA, which are comparable to currents observed in the return strokes of negative CG lightning. Our estimate of LNO<sub>x</sub> production falls in the lower end of the Wang et al. [1998] range. While IC lightning is assumed to have low peak currents (a few kA) relative to CG flashes, our results and those of Stith et al. [1999] and Huntrieser et al. [2002] suggest that IC lightning is producing significantly more NO<sub>x</sub> than expected from simple scaling of the Wang et al. [1998] results using peak currents and pressure. This discrepancy suggests that peak currents in IC lightning may be higher, or that some other aspect of NO<sub>x</sub> production in IC lightning is not understood.

[35] Flashes are also determined from the interferometer data, and can be used to compute LNO<sub>x</sub> production per flash. Over the period from 2230 GMT to 2400 GMT, the average interferometer flash rate was approximately 0.27 flashes per second for flashes with durations from 23  $\mu$  sec to over 1 sec. Combining this flash rate with the LNO<sub>x</sub> production rate leads to an average LNO<sub>x</sub> production rate per flash of 43.2 moles/flash ( $2.6 \times 10^{25}$  molecules/flash). This estimate of NO<sub>x</sub> production per flash is consistent with existing estimates. Lawrence et al. [1995], in a review of previous theory, laboratory experiments and observational analyses of mostly CG flashes, found that between 0.5 to  $300 \times 10^{25}$  molecules of NO are produced per lightning flash. Our estimate of  $2.6 \times 10^{25}$  molecules per flash falls on the lower end of the Lawrence et al. range. Because the interferometer detects both shorter and weaker flashes compared to satellite observations or most previous ground-based observations of flashes, it is not surprising that our estimate is on the lower end. Additionally, the Lawrence et al. [1995] study was primarily concerned with CG lightning while the 10 July storm produced primarily IC lightning during the analysis period. IC flashes have lower peak currents (~ few kiloamps) and occur at lower air pressure and densities compared to CG lightning, and are expected to produce less NO<sub>x</sub> per flash.

[36] The measures of lightning activity for the 10 July storm were quite variable, and this variability raises the question of whether or not the objectively analyzed lightning-produced NO<sub>x</sub> flux is appropriately representative. Both the lightning channel-length production rate and the interferometer flash rate have minima in the early portion of the observational period and maxima in latter part of the period. The objectively analyzed NO has its peak values in the upper part of the anvil, presumably attributable to lightning-produced NO. This portion of the anvil was sampled early in the observational period, thus it may be that our analysis underestimates the lightning NO<sub>x</sub> production. Conversely, the final pass through the anvil (the spiral pass up through the middle of the anvil that can be seen in the aircraft track in Figure 4) also showed more NO<sub>x</sub> in the upper part of the anvil and did not detect any significant increase in NO<sub>x</sub> compared to the earlier horizontal traversals. This suggests that the analysis has produced a reasonable estimate of the lightning-produced NO<sub>x</sub> flux.

[37] Finally, we would like to note that the analyses that have been combined to produce the lightning production estimates are linear. Thus an error of  $\pm 20\%$  in the  $LNO_x$  production will produce a  $\pm 20\%$  error in the per channel length and per flash production estimates.

#### 3.2. NO<sub>x</sub> Storm Budget

[38] The ONERA lightning interferometer data span the entire lifetime of the 10 July storm. The total  $NO_x$  produced by lightning in the storm can be estimated by multiplying the  $NO_x$  production rate per unit flash by the total number of flashes, or by multiplying the production rate per meter lightning channel length by the total channel length produced over the lifetime of the storm. Using the total lightning channel length ( $\sim 370,000~\text{km}$ ) and the  $NO_x$  production rate per meter ( $\sim 1.7~\times~10^{-3}~\text{moles/m}$ ) we compute a net production of  $\sim 6.3~\times~10^{5}~\text{moles}$  (3.8  $\times~10^{29}~\text{molecules}$ ) of  $NO_x$ . Using the total number of flashes ( $\sim 5428$ ) and the

 $NO_x$  production rate per flash ( $\sim 2.6 \times 10^{25}$  molecules per interferometer flash) we compute a net production of  $\sim 2.3 \times 10^5$  moles ( $1.4 \times 10^{29}$  molecules) of  $NO_x$ .

- [39] The difference between the net storm  $LNO_x$  production estimates is the result of the change in the average flash channel length over the storm lifetime; at later times the flash channels were on average longer than those measured during the observation period used to compute anvil  $NO_x$  flux, thus using the flash rate produces a lower integrated  $NO_x$  production estimate.
- [40] As in the lightning analyses, the analyses that have been combined to produce the net storm  $LNO_x$  production estimates are linear. Thus an error of  $\pm 20\%$  in the  $LNO_x$  production will produce a  $\pm 20\%$  error in the net storm production estimate.

#### 3.3. Global LNO<sub>x</sub> Budgets

[41] An important goal for studies of NO<sub>x</sub> production by lightning is to determine the contribution which lightning makes to the global nitrogen budget. Generally, two basic approaches have been used to determine a global LNO<sub>x</sub> production rate; either an estimate of per-storm LNO<sub>x</sub> production is multiplied by an estimate of the number of storms occurring over the globe per year [e.g., Chameides et al., 1987], or an estimate of the global lightning flash count per year is multiplied by an estimate of the average LNO<sub>x</sub> production per flash (e.g., the Flash Extrapolation Approach (FEA) reviewed by Lawrence et al. [1995]). Estimating global LNO<sub>x</sub> production using the results from our study and using either of the extrapolation approaches is problematic. For example, our estimate of total storm LNO<sub>x</sub> production  $(32-88 \times 10^5 \text{ g(N)/storm})$  can be combined with the commonly cited storm frequency number from Brooks [1925] of  $16 \times 10^6$ , storms/year (44,000 storms per day) to arrive at an LNO<sub>x</sub> production rate of 52-141 Tg(N)/year. This estimate is on the high end of global LNO<sub>x</sub> production rate estimates [see *Huntrieser et al.*, 1998], perhaps suggesting that the 10 July storm produced much more LNO<sub>x</sub> than the average storm (it was quite long-lived and produced a large amount of lightning). Unfortunately, we have no quantitative way to normalize our LNO<sub>x</sub> production rate for the 10 July storm without independent measures of average-storm characteristics, hence this estimate is highly uncertain. Alternatively, consider multiplying our estimate of production per flash (600 g(N)/flash) by an estimate of the global flash rate (44  $\pm$  5 flashes/s; D. J. Boccippio, personal communication) [Mackerras et al., 1998; Williams et al., 2000]. This approach leads to a global production rate of 0.8 Tg(N)/year. The difficulty with this estimate is that we do not have sufficient information to relate interferometer flashes for the 10 July storm (determined in this study) with optical detector flashes (from which the global flash rate is derived). The interferometer likely sees many more flashes, and certainly many more of the weaker flashes, than the optical detectors, suggesting perhaps that this production rate estimate may be low. Thus this alternative estimate is also highly uncertain.

#### 4. Summary and Discussion

[42] Using in situ aircraft observations, cloud model simulations, lightning interferometer observations and radar,

we have produced an estimate of LNO<sub>x</sub> (lightning-produced NO<sub>x</sub>) production during the 10 July STERAO storm, along with estimates of average LNO<sub>x</sub> production rate per flash, per flash channel length, and net LNO<sub>x</sub> production over the lifetime of the storm. The production rates are computed using a control volume analysis around the storm that makes use of objectively analyzed in situ aircraft observations of NO<sub>x</sub> flux through a control volume surface into the anvil, and cloud-model simulations that allow partitioning of the NO<sub>x</sub> flux into an LNO<sub>x</sub> flux and an environmental NO<sub>x</sub> flux. The analysis reveals that approximately 60% of the NO<sub>x</sub> passing out into the anvil was produced by lightning, with an average LNO<sub>x</sub> flux of 11.5 moles s<sup>-1</sup> (6.9  $\times$  $10^{24}$  molecules s<sup>-1</sup>) during the analysis period. From the lightning interferometer data, we find that the average channel length and flash production during the analysis period was 6700 m s<sup>-1</sup> and 27 flashes per second, respectively. Combining these with the LNO<sub>x</sub> flash length production rate leads to an estimate of the average LNO<sub>x</sub> production per meter flash channel length of  $1.7 \times 10^{-3}$ moles/m (or  $1.0 \times 10^{21}$  molecules/m). This is within the range of field estimates computed by Stith et al. [1999] of  $2 \times 10^{20}$  and  $10^{22}$  molecules/m and the laboratory estimates computed by Wang et al. [1998] of between  $5 \times 10^{20}$  and  $3 \times 10^{21}$  molecules/m. The production rate per flash from our analysis is 43 moles  $NO_x$  per flash (2.6  $\times$  10<sup>25</sup> molecules per flash), and falls at the lower end of the estimate of Lawrence et al. [1995] of 0.5 to  $300 \times 10^{25}$ molecules per flash. The interferometer can detect both weaker flashes and flashes of shorter duration than detected by most previously used measurements systems, and the 10 July storm produced mostly IC lightning in the analysis period. Both of these considerations are consistent with the fact that our estimates of LNO<sub>x</sub> production per flash and per unit channel length are at the low end estimates derived in previous studies.

- [43] The interferometer data also provide measurements of total flashes and total flash channel length produced over the lifetime of the storm, both of which can be used to compute net storm LNO<sub>x</sub> production. The net storm LNO<sub>y</sub> production is estimated as  $\sim 6.3 \times 10^5$  moles (3.8  $\times$  10<sup>29</sup> molecules) of NO<sub>x</sub> using 3.7  $\times$  10<sup>8</sup> meters of interferometer flash channel produced over the lifetime of the storm [*Deferet al.*, 2003]. Alternatively, we have derived a net storm LNO<sub>x</sub> production of  $\sim$ 2.4  $\times$  10<sup>5</sup> moles (1.4  $\times$  10<sup>29</sup> molecules) of NO<sub>x</sub> using 5428 interferometer flashes [*Deferet al.*, 2003] and our estimate of 43 moles (2.6  $\times$  10<sup>25</sup> molecules) per flash.
- [44] The importance of LNO<sub>x</sub> production derives from the need to understand NO<sub>x</sub> and ozone chemistry on a global scale. The LNO<sub>x</sub> component of global NO<sub>x</sub> budgets have typically been derived by producing estimates of global flash rates and combining these estimates with estimates of LNO<sub>x</sub> production per flash [e.g. *Lawrence et al.*, 1995], or by estimating per-storm LNO<sub>x</sub> production and multiplying this by an estimate of the number of storms occurring globally in a year [*Chameides et al.*, 1987]. These global LNO<sub>x</sub> production rates vary by more than an order of magnitude [*Liaw et al.*, 1990], and represent the largest uncertainty in the global NO<sub>x</sub> budgets. Neither approach for producing a global budget can be directly applied using the results from this study. Problems arise from the fact that (1)

the interferometer flashes represent a new metric in  $LNO_x$  studies, and we don't have sufficient data to relate this new metric to the satellite optical flash detector counts needed for an extrapolation to the globe, and (2) we have little information that allows us to relate the 10 July storm to the average storm which is used in other estimates of global  $LNO_x$  production rates.

[45] Additional studies such as the one presented here may be able to reduce the uncertainty in LNO<sub>x</sub> production estimates to a limited extent, but they will only apply to the particular storms being studied and extrapolation will be difficult. To make further progress we need to be able to relate the new local measurements (interferometer flashes) to existing global measurements (OTD flashes), and we need to better understand the underlying mechanisms controlling flash rates, lengths, durations, energies, and associated LNO<sub>x</sub> production. More detailed studies of storm structure and evolution are needed, specifically with respect to the critical role storm microphysics plays in lightning generation, and these studies must encompass a variety of convective types known to be electrically active, both midlatitude and tropical in their various seasonal regimes. Our more detailed but limited observations show that the character of the lightning in individual storms can vary greatly. For example, it is well documented that continental thunderstorms produce much more lightning than maritime storms. Thus the significant uncertainty associated with studies of individual storms and the uncertainty associated with extrapolations to other storms is magnified.

[46] Another important aspect of the LNO<sub>x</sub> budget is quantifying the vertical distribution of LNO<sub>x</sub> sources and subsequent LNO<sub>x</sub> transport. The general consensus on where LNO<sub>x</sub> is deposited after its production has changed considerably over the last decade. While we have evolved from an assumption of a uniform vertical NOx distribution to a more kinematically based assumption of a C-shaped NO<sub>x</sub> profile, the differences in production and transport in convective storms suggest that no single generalization is applicable to all cases. For example, in the 10 July storm it appears that most of the LNO<sub>x</sub> was transported into the anvil with little deposited below cloud base. Storms with different kinematic structures are likely to transport LNO<sub>x</sub> to different altitudes, but until we have more observations in different type storms we cannot quantify this. In this respect it is very important to examine LNOx production and transport by tropical storms because globally much of the earth's lightning is produced in the tropics.

[47] **Acknowledgments.** We would like to thank all those who participated in the STERAO field campaign and subsequent data analysis. This work is partially supported by NOAA CGCP award NA76GP00400. The Citation data collection and analysis were supported by the National Science Foundation through grant ATM-9634125. The National Center for Atmospheric Research is operated by the University Corporation for Atmospheric Research under the sponsorship of the National Science Foundation.

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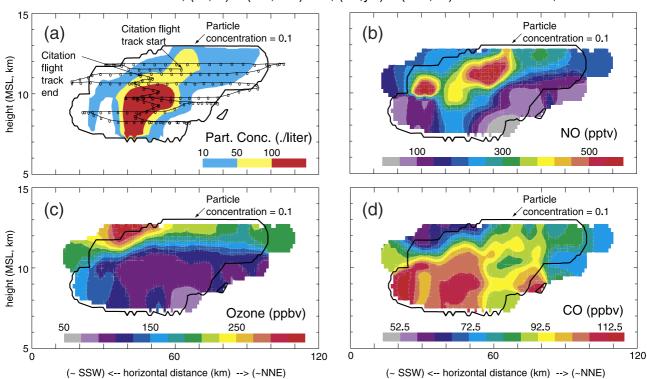
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# Citation flight 10 July 1996, 23:16:00 10 July to 00:36:40 11 July vertical cross section, $(ud, vd) = (1.5, -5.5) \text{ m/s}, (xm, ym) = (130, 65) \text{ km from CHILL, rotated } 15^{\circ}$



**Figure 4.** Objective analysis of (a) cloud particle concentration, (b) NO, (c) O<sub>3</sub>, and (d) CO using the Citation flight data. The flight track is shown in the cloud particle concentration plot.