Downloading and compiling MPAS-Atmosphere

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Before getting in to the details of how to download and compile the code, let's first answer the question:

On which systems can MPAS-Atmosphere be run?

There are essentially two considerations:

- Hardware
- Software





Generally speaking, no requirement regarding number of CPUs or type of CPU

 But you must have relatively modern Fortran and C/C++ compilers that target your CPU architecture

However, the amount of memory (RAM) dictates the maximum domain size

 Roughly 0.175 MB/memory per grid column (assuming 55 vertical levels and 'mesoscale_reference' physics suite)

As long as you're patient enough, and as long as your simulation fits in memory, a single processor is all you need!

• On the upper end of CPU count, model generally scales well to around 100 grid columns per MPI task







Example systems that run MPAS-A





Working under the assumption that we have a system that meets the basic requirements, we can then ask:

How do I get a copy of the source code?

There are essentially two options:

- 1) The "traditional", but not necessarily encouraged, method
 - Download a .tar.gz or .zip file
- 2) The preferred method
 - Make a *clone* of the MPAS-Model repository



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One can navigate to a download link from the MPAS homepage at https://mpas-dev.github.io/

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MPAS Overview

The Model for Prediction Across Scales (MPAS) is a collaborative project for developing atmosphere, ocean and other earth-system simulation components for use in climate, regional climate and weather studies. The primary development partners are the climate modeling group at Los Alamos National Laboratory (COSIM) and the <u>National Center for</u> <u>Atmospheric Research</u>. Both primary partners are responsible for the MPAS framework, operators and tools common to the applications; LANL has primary responsibility for the ocean and land ice models, and NCAR has primary responsibility for the atmospheric model.



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The defining features of MPAS are the unstructured <u>Voronoi meshes</u> and <u>C-grid</u> discretization used as the basis for many of the model components. The unstructured Voronoi meshes, formally Spherical Centriodal Voronoi Tesselations (SCVTs), allow for both quasi-uniform discretization of the sphere and local refinement. The C-grid discretization, rmal component of velocity on cell edges is prognosed, is especially well-suited to migner resolution, mesoscale <u>atmosphere</u> and <u>ocean</u> simulations. The land ice model takes advantage of the SCVT-dual mesh, which is a triangular Delaunay tessellation appropriate for use with Finite-Element-based discretizations.

The current MPAS release is version 8.1.0. Please refer to each core for changes, and to the GitHub repository for source.



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MPAS Atmosphere Public Releases

MPAS Atmosphere 8.1.0 was released on 18 April 2024.

For information on the GPU-enabled MPAS-Atmosphere model, please refer to this documentation

MPAS Atmosphere 8.1.0 release notes

Source code downloads:

- MPAS v8.1.0
- GPU-enabled MPA -Aunosphere v6.x

MPAS-Atmosphere Users' Guide

MPAS-Atmosphere tutorial

MPAS-Atmosphere meshes

Static geographical datasets

Configurations for idealized test cases

Sample input files for real-data simulations

Visualization and analysis tools



A variable resolution MPAS Voronoi mesh



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License Information

New Users: We request that users of MPAS software register. The information allows us to better determine how to support and develop the model. Please register using this <u>form</u>.

After you register: Please considering subscribing for the relevant mailing lists.

Registered Users: If you have already registered, please continue <u>here</u>.





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MPAS-Atmosphere download MPAS-Albany Land Ice download MPAS-Ocean download MPAS-Seaice download PLEASE READ THIS ENTIRE PAGE BEFORE DOWNLOADING ANY OF THE MPAS SOURCE CODE.

NOTE: If you intend to collaborate on a project, please read the Developer's guide before beginning work. The access method for the MPAS source code you choose will impact development efforts.

The easiest way to acquire the MPAS source code is through the archive file below. Although this is the easiest method, it is also the least flexible. The method used will largely be determined by your use case. You should choose the method that aligns best with your use case.

Archive files are provided both in zip and tar.gz formats. Each release provides an archive file, and users should download the archive file for the most relevant released version.

Archive file	Here	
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Our recommended method is to use git. This will allow users to easily update the version they are working with, and it facilitates the most direct method for contributing development back into the MPAS code base. This comes with the caveat of having to know how to use git. There are a few links at the bottom of this page that can help you get started with git. The most immediately useful tutorial is the GitHub fork tutorial, but the other tutorials will help you with learning how to make use of git.

GitHub Page Here

Git related resources:

One can navigate to a download link from the MPAS homepage at https://mpas-dev.github.io/

2 days ago ∰ mgduda ♥ v8.1.0 -> f084b36 Compare ▼	This release of MPAS introduces several upd • The MYNN Surface Layer and Planetary includes the options needed to run the E	Boundary Layer (PBL) schemes have	
	 A separate driver for the parameterizatio The build system now tests for the availather mpi_f08 module interface to Fortrant The atmosphere core now supports the offering significant disk space savings, e being run. See Section 8.3 of the User's The top-level Makefile provides a new i suite. To support real-data simulations on mes core provides an option, config_30s_su fields (terrain, land use, soil category, and support set of the top support real-data simulations on the suite. 	In MPI routines will be used in favor o storage and retrieval of time-invarian especially in cases where large ensen <u>Guide</u> for more details. intel build target for the Intel oneAP shes with grid distances approaching upersample_factor, for super-sampli	points has been added. le in the MPI library. If detected, f the older mpi module interface t fields in a separate I/O stream, nbles or cycled simulations are I Fortran, C, and C++ compiler 1 km or less, the init_atmosphere ng of the 30 arc-second terrestria
	 When compiling, PRECISION=single is r PRECISION=double . Assets 2 	now the default, and double-precision	n builds must be specified with
	Source code (zip)		2 days ago



MPAS

on Across Scales



Downloading a .tar.gz file of a particular release of the MPAS code certainly works, but it has several disadvantages:

- 1. You'll only obtain a specific release of the code
- 2. There's no direct way to see the history of changes to parts of the code
- 3. It's more difficult to see what local modifications have been made to the code
- 4. There's no easy route to updating to a newer release while preserving your local code modifications



A much better option is to *clone* the MPAS-Model repository

The repository URL can be found from the MPAS GitHub page at https://github.com/MPAS-Dev/MPAS-Model

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The Model for Prediction Across Scales	(MPAS) is a collab	porative project for	r developing atr	nosphere	🕐 🗂 🕰 🚳 🖉	



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From the command-line, the following should be sufficient:

git clone https://github.com/MPAS-Dev/MPAS-Model.git

Cloning the repository should take about 10 seconds or less...

• (I timed this at home, and it took 5.16 s)

```
$ git clone https://github.com/MPAS-Dev/MPAS-Model.git
Cloning into 'MPAS-Model'...
remote: Enumerating objects: 71, done.
remote: Counting objects: 100% (71/71), done.
remote: Compressing objects: 100% (47/47), done.
remote: Total 46608 (delta 38), reused 42 (delta 23), pack-reused
46537
Receiving objects: 100% (46608/46608), 19.65 MiB | 2.57 MiB/s,
done.
Resolving deltas: 100% (35848/35848), done.
```



You may also like to register as an MPAS user and join the MPAS-Atmosphere Users mailing list

MPAS Model for Prediction Across Scales

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After you register: Please considering subscribing for the relevant <u>mailing lists</u>.

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In order to compile MPAS and its required libraries, relatively modern Fortran, C, and C++ compilers are necessary

- The Fortran compiler should be recent enough to support the ISO_C_BINDING module from the Fortran 2003 standard, procedure pointer components of derived types, and a few bits of the Fortran 2008 standard
- Most versions of common compilers from the last couple of years should be fine

Building MPAS requires at least the following libraries:

- Any implementation of MPI-2, e.g., MPICH, MVAPICH, OpenMPI
 - Ensure that mpif90 and mpicc commands are in your path
- Parallel-NetCDF (<u>http://trac.mcs.anl.gov/projects/parallel-netcdf</u>/)
 - Set PNETCDF environment variable to base installation directory



Installing the full set of I/O libraries required by MPAS can be tedious.

To make this easier, we've prepared a shell script that may be helpful:

https://www2.mmm.ucar.edu/projects/mpas/scripts/mpas_lib_install.sh

Notes:

- Before running or following the above script, you will need to have downloaded the library sources from https://www2.mmm.ucar.edu/projects/mpas/scripts
- If you already have an MPI library, skip the MPICH installation
- After editing paths and compiler names at the top, you may be able to run the script, but in general the script may be best used as a guide



Model Organization



Checking out the MPAS code provides all MPAS models, not just MPAS-Atmosphere

- All models share a common set of infrastructure modules
- Each MPAS model is implemented as a "core" that lives in its own directory
- User must select which "core" to compile
- Each "core" is associated with a source code subdirectory under src/ and has a Registry file (similar to WRF)



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Running MPAS-Atmosphere involves two "cores":

- The init_atmosphere core is responsible for
 - Interpolating static fields to the mesh
 - Generating a vertical grid
 - Horizontally and vertically interpolating meteorological data to the 3-d grid
- The **atmosphere** core is the model itself; integrates forward in time from initial state



There is no "configuration" step for MPAS, unlike, e.g., for the WRF model

• All build flags are either set in the top-level Makefile or on the command-line

General MPAS build command:

\$ make target CORE=core <options>

```
target can be either
clean
or
gnu
intel
nvhpc
llvm
... plus a few others...
```

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For MPAS-Atmosphere, core may be

atmosphere init_atmosphere

<options> can be any of

DEBUG=true AUTOCLEAN=true PRECISION=double OPENMP=true



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Typical build of both the init_atmosphere and atmosphere cores involves:

make gnu CORE=init_atmosphere (build init_atmosphere_model)

make gnu CORE=atmosphere (build atmosphere_model)

By default, MPAS cores are built with single-precision floating-point variables

MPAS-Atmosphere can be built in double precision

- Add PRECISION=double to build commands for double-precisoin executables
- Execution time is slower compared with single-precision
- Output files approximately twice as large



- Ensure that you have a system running a UNIX-like operating system with the usual commands: make, awk, sed, grep, *git*, etc.
 - Available memory dictates maximum number of grid cells
- Ensure that you have modern Fortran, C, and C++ compilers
- Install supporting libraries (MPI, Parallel-NetCDF; or if you need PIO: HDF5, NetCDF4, Parallel-NetCDF, PIO)
 - Can use mpas_lib_install.sh script as a guide
- Obtain the MPAS-Model source code with *git clone*
- Compile the *init_atmosphere* and *atmosphere* cores