

# Sharing Physics Between WRF and MPAS with CCPP

Dave Gill Laura Fowler

NCAR UCAR Ming Chen Jimy Dudhia Jihyeon Jang Wei Wang Kelly Werner Cheryl Craig Steve Goldhaber

June 9, 2020



Boulder, CO

# **Big Picture: Sharing**

- MPAS and WRF share
  - Support technology: <u>http://forum.mmm.ucar.edu</u>
  - Pre-processor input from WPS
  - Repo management (git) and hosting services (github)
  - Basic tenets of gitflow-esque workflow
  - A couple of physics suites (officially)
  - Hybrid vertical coordinate and moist theta in WRF came from MPAS
- Want to more easily transfer physics between MPAS and WRF
- A goal is to allow physics transfer between MPAS and other models
- Several groups within NCAR (ACOM, CGD, MMM) and NOAA are working on (and are in various stages of implementing) a strategy for more widely sharing physics

# CCPP = Common Community Physics Packages

## **Big Picture: Why are we doing this**



Original interface to physics and host model.

The CCPP interface permits the same physics to be used with both the original and a new host model.

## **Initially Selected Schemes to Port**

Surface Layer YSU PBL New Tiedtke CU WSM6 MP Xu-Randall Cloud Fraction GWDO RRTMG LW RRTMG SW

- Choosing to wait for the Land Model.
  - Does not pose troubles for eventual CESM port.
  - Does not pose troubles for MPAS or WRF port.
- CTSM effort is underway at NCAR. LILAC may provide a preferred way to access a supported land model for both MPAS and WRF.





- Modifications to scheme for CCPP
  - Push as much pre- and postprocessing into scheme as makes sense
  - Array assignments in scheme top-level call due to horizontal index reduction are OK and expected
  - Basically: allow removal of toplevel of scheme

xt24 = mod(xtime+radt\*0.5,1440.)
tloctm = gmt + xt24/60. + xlong(i,j)/15.
hrang = 15. \* (tloctm-12.) \* degrad
xxlat = xlat(i,j) \* degrad

qv2d(its:ite,:) = 0.0

ZAP: Computation of constants ZAP: Initializations to zero ZAP Diagnostics

```
do k = kts,kte
   do i = its,ite
    rthblten(i,k,j) = rthblten(i,k,j)/pi3d(i,k,j)
```

- Modifications to scheme for CCPP
  - Interstitial code for helper routines called in top-level (non CCPP-ized) routines

NEW:

Assignments (3d -> 2d) call ra\_rrtmg\_sw\_mpas\_timestep\_init call ra\_rrtmg\_sw\_run call ra\_rrtmg\_sw\_mpas\_timestep\_final Assignments (2d -> 3d)

ORIGINAL: Starting on line 10000 – 400 lines of diagnostics surrounding the call to SWRAD

- General Modifications
  - Clean up INTENT missing, particularize, wrong
  - Reduce argument list to required minimal set
  - Fix internal comments
  - Single argument *mostly* = single field
    - Unless multiple fields are not distinguishable within scheme
    - Derived types are allowable, but break portability
    - So, interstitial code would be OK for MPAS DDTs
  - Most of the OPTIONAL arguments were not
- A cookbook for developers to follow
  - Automated code depends on an accurate metadata file
  - Constructing the metadata file takes time

## The Metadata File

- For each of the CCPP-ized files, a metadata file is required
- For each CCPP-ized subroutine, a metadata section is required in that file
- Each argument in a CCPP-ized routine has a metadata entry
- The mapping of the host model's variables (also with a metadata file) and the physics scheme's variables are through the unique identifier of the field's name
- The ordering of the arguments in the subroutine matches the ordering of the variables in the metadata file

### **Physics Scheme and Metadata**

```
subroutine bl gwdo run(
                        &
  sina, cosa,
                        &
  rublten,rvblten,
                        &
  dtaux3d,dtauy3d,
                        &
  dusfcq,dvsfcq,
                        &
  uproj, vproj,
                        &
  t1, q1,
                        &
  prsi, prsl, prslk, zl, &
  kpblmax,
                        &
  var, ocl,
                        δ
  oa2d1, oa2d2,
                        &
  oa2d3, oa2d4,
                        &
  ol2d1, ol2d2,
                        &
  ol2d3, ol2d4,
                        &
  g, cp, rd,
                        &
  rv_, fv_, pi_,
                        &
  dxmeter, deltim, kpbl, &
  its, ite,
                        &
  kts, kte, kme,
                        &
  errmsq, errflq
                        )
```

[ccpp-arg-table]
name = <mark>bl_gwdo_run</mark>
type = scheme
[sina]
standard_name = sine_of_map_rotation
long_name = sine of map rotation
units = dimensionless
dimensions = (horizontal_loop_begin:horizontal_loop_end)
type = real   kind = kind_phys
intent = in
optional = F
[cosa]
standard_name = cosine_of_map_rotation
long_name = cosine of map rotation
units = dimensionless
dimensions = (horizontal_loop_begin:horizontal_loop_end)
type = real   kind = kind_phys
intent = in
optional = F

# The Metadata File

 A github repository holds the growing list of known fields (such as dimensions) and conventionally accepted names:

https://github.com/ESCOMP/CCPPStandardNames/blob/master/Metadata-standard-names.md

- Fields with vertical indexing may have "layers" (default) and "levels" (explicitly stated in name)
- Horizontal indexing allows full size (without halo) and threaded size
- The list of known constants is increasing
- Construction of names is often required

### The Suite Definition File (SDF mostly TBD)

```
<suite name="MPAS meso ref rad" version="1.0">
                                                              <suite name="mesoscale reference" version="1.0">
  <proup name="radiation">
                                                                 <proup name="physics pre LSM">
     <scheme>cld xurandall</scheme>
                                                                    <suite name="MPAS meso ref rad" group="radiation" />
                                                                    <suite name="MPAS meso ref sfc" group="surface layer" />
     <scheme>ra rrtmg sw mpas</scheme>
     <scheme>ra rrtmg sw</scheme>
                                                                 </group>
     <scheme>ra rrtmg lw mpas</scheme>
                                                                 <proup name="physics post LSM">
                                                                    <suite name="MPAS meso ref pbl" group="boundary layer" />
     <scheme>ra rrtmg lw</scheme>
                                                                    <suite name="MPAS meso ref cu" group="cumulus" />
  </group>
</suite>
                                                                 </group>
                                                                  <proup name="physics post dynamics">
                                                                    <suite name="MPAS meso ref mp" group="microphysics"</pre>
<suite name="MPAS meso ref sfc" version="1.0">
  <proup name="surface layer">
                                                                 </group>
     <scheme>sf sfclay</scheme>
                                                               </suite>
  </group>
</suite>
                                                             The "group" element allows MPAS to
<suite name="MPAS meso ref pbl" version="1.0">
  <proup name="boundary layer">
                                                             schedule physics in a requested order
     <scheme>bl ysu</scheme>
     <scheme>bl gwdo</scheme>
  </aroup>
</suite>
<suite name="MPAS meso ref cu" version="1.0">
  <proup name="cumulus">
     <scheme>cu ntiedtke</scheme>
                                                             The "group" attribute allows other host
  </group>
</suite>
                                                             models to locate where the call to a
<suite name="MPAS meso ref mp" version="1.0">
                                                             particular scheme should go in the time
  <proup name="microphysics">
     <scheme>mp wsm6</scheme>
                                                             step
  </group>
</suite>
```

# **Changes MPAS**

- The WRF model will not be converted to a CCPP-ized host model
- MPAS will have a transition period, allowing access to physics by the original driver and the new CCPP option
- The existing top-level drivers of each scheme remain in place
  - These are not required when CCPP is used
- Rolling the calls to the \_to\_MPAS and \_from\_MPAS routines into interstitial routines
  - Interstitial routines are by definition associated with the host model
  - Maintains bit-for-bit results

### Easy Changes to MPAS to "interface" and "drivers"

mpas\_atmphys\_interface

•  $(k,i) \rightarrow (i,k,j) \rightarrow (i,k) \rightarrow (i,k,j) \rightarrow (k,i)$ 

mpas\_atmphys\_interface\_ccpp

if(.not.allocated(psfc\_p) ) allocate(psfc\_p(ims:ime,jms:jme) )

 $rho_p(i,k,j) = zz(k,i) * rho_zz(k,i)$ 

if(.not.allocated(psfc p) ) allocate(psfc p(ims:ime)

 $rho_p(i,k) = zz(k,i) * rho_zz(k,i)$ 

## **Big Picture: What we are doing**

- Scheme sharing between MPAS and WRF
- Purpose is to get MPAS and CAM physics available for each other
- Using a jointly developed utility to provide Common Community Physics Packages (CCPP)
- Conventionalize the top-level physics schemes
- Describe the arguments in the top-level physics schemes in a separate metadata file
- Describe the available fields from the host model is a metadata file
- Map the available fields from the host model and those required by the physics schemes

## **Status: Importance of Verb Tense**

- Started and stopped a few times to get our bearings and scope
- All first suite physics is ported, including non-conventionalized metadata files
- Getting all metadata and CCPP-ized routines to be uniform
- Merging all separate git repositories into a single location



Figuring out the data flow for the physics diagnostic data, leaning towards including / wrapping calls inside of interstitial routines