

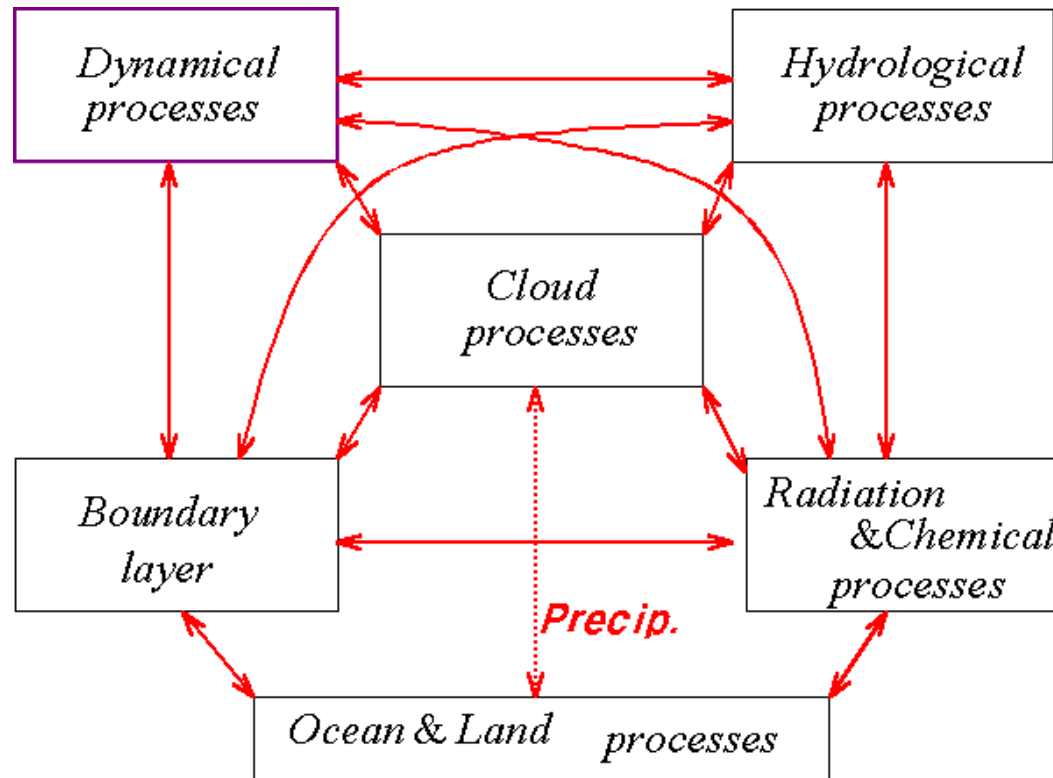
# **Overview of Physical Parameterizations**

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# 1) Concept



## \* Physical processes in the atmosphere

: Specification of **heating, moistening and frictional terms** in terms of dependent variables ( $U, V, T, q$  etc) of prediction model  
→ Each process is a specialized branch of atmospheric sciences.

## \* Parameterization

The formulation of physical process **in terms of the model variables** as parameters, i.e., constants or functional relations.

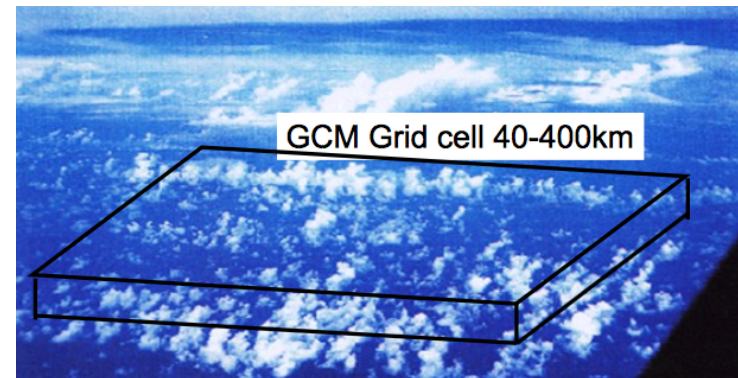
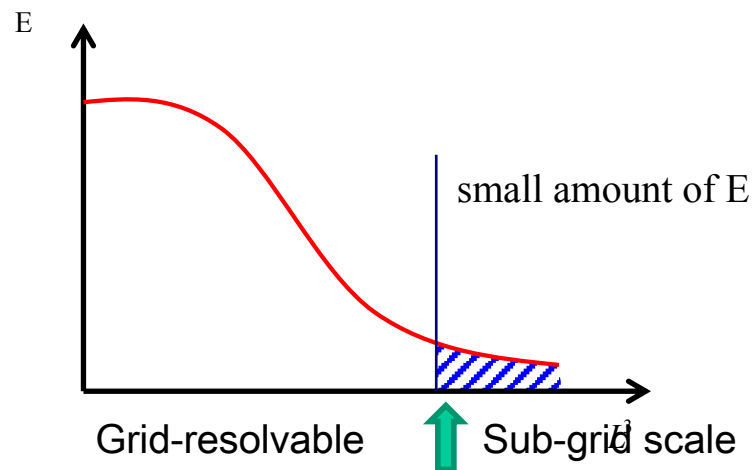


# 1) Concept...continued

## Subgrid scale process

Any numerical model of the atmosphere must use a finite resolution in representing continuum certain **physical & dynamical phenomena that are smaller than computational grid.**

- Subgrid process (Energy perspective)



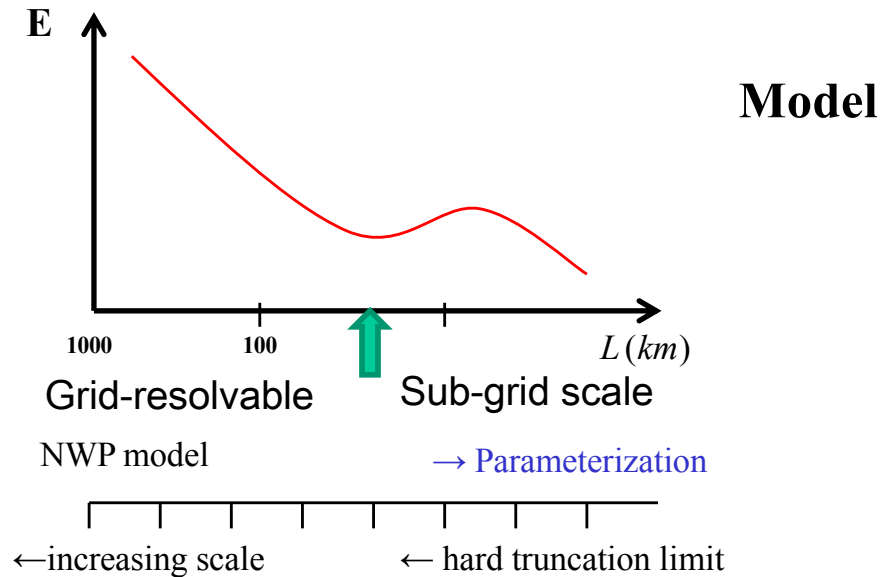
$\Delta x \rightarrow 0$ , the energy dissipation takes place by molecular viscosity

## Objective of subgrid-scale parameterization

“To design the physical formulation of energy sink, withdrawing the **equivalent amount of energy** comparable to cascading energy down **at the grid scale** in an ideal situation.”

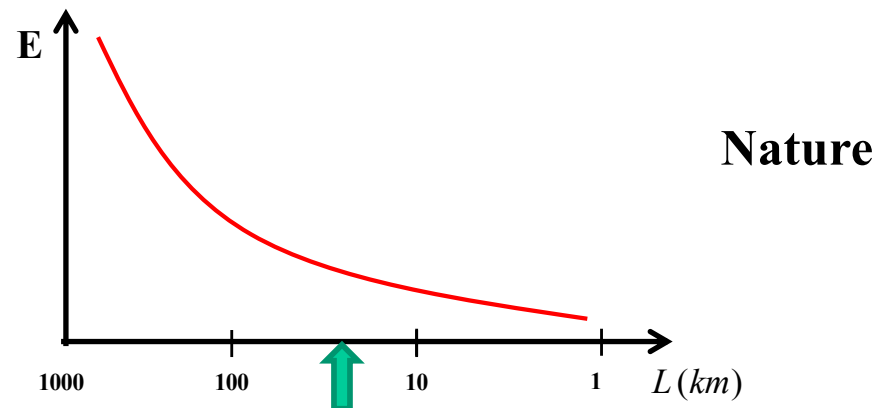
# 1) Concept...continued (difficulties in physical parameterizations)

⌘ Parameterization that are only somewhat smaller than the smallest resolved scales.



where truncation limit ; spectral gap

Unfortunately, there is **no** spectral gap



## 2) Subgrid scale process & Reynolds averaging

Consider prognostic water vapor equation

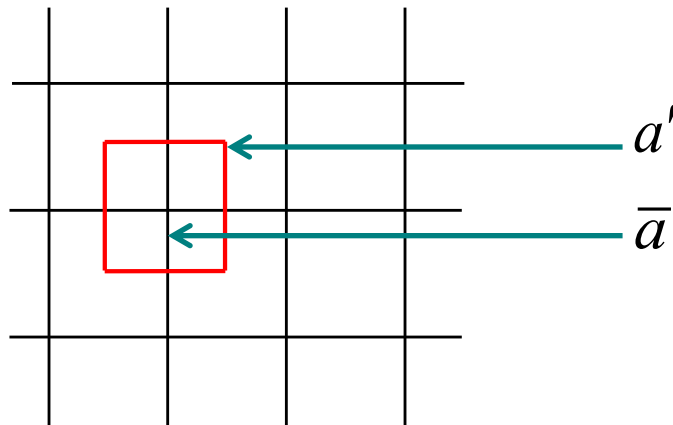
E: evaporation, C: condensation

$$\frac{\partial \rho q}{\partial t} = -\frac{\partial \rho u q}{\partial x} - \frac{\partial \rho v q}{\partial y} - \frac{\partial \rho w q}{\partial z} + \rho E - \rho C \quad \dots(1)$$

In the real atmosphere,

$$u = \bar{u} + u', \quad q = \bar{q} + q' \quad \left( \begin{array}{l} \bar{a} : \text{grid-resolvable} \\ a' : \text{subgrid scale perturbation} \end{array} \right)$$

$\rho'$  is neglected



Model predicts the domain-averaged quantity,  $\bar{a}$

Then, how to represent the subgrid-scale process ?

## 2) Subgrid scale process & Reynolds averaging...continued

\* **Rule of Reynolds average :**  $\overline{q'} = 0$ ,  $\overline{u'q} = 0$ ,  $\overline{\bar{u} \bar{q}} = \bar{u} \bar{q}$

then eq.(1) becomes

$$\frac{\partial \rho \bar{q}}{\partial t} = \underbrace{-\frac{\partial \rho \bar{u} \bar{q}}{\partial x} - \frac{\partial \rho \bar{v} \bar{q}}{\partial y} - \frac{\partial \rho \bar{w} \bar{q}}{\partial z}}_{\textcircled{1}} - \underbrace{\frac{\partial \rho \overline{u'q'}}{\partial x} - \frac{\partial \rho \overline{v'q'}}{\partial y} - \frac{\partial \rho \overline{w'q'}}{\partial z}}_{\textcircled{2}} + \rho E - \rho C \dots (2)$$

① grid-resolvable advection (dynamical process)

② turbulent transport (parameterized process)

### \* How to parameterize the effect of turbulent transport

a)  $-\rho \overline{w'q'} = 0$  : 0th order closure

b)  $-\rho \overline{w'q'} = K \frac{\partial \bar{q}}{\partial z}$  : 1st order closure (K-theory)

c) obtain a prognostic equation for  $\overline{w'q'}$  from (1), (2)

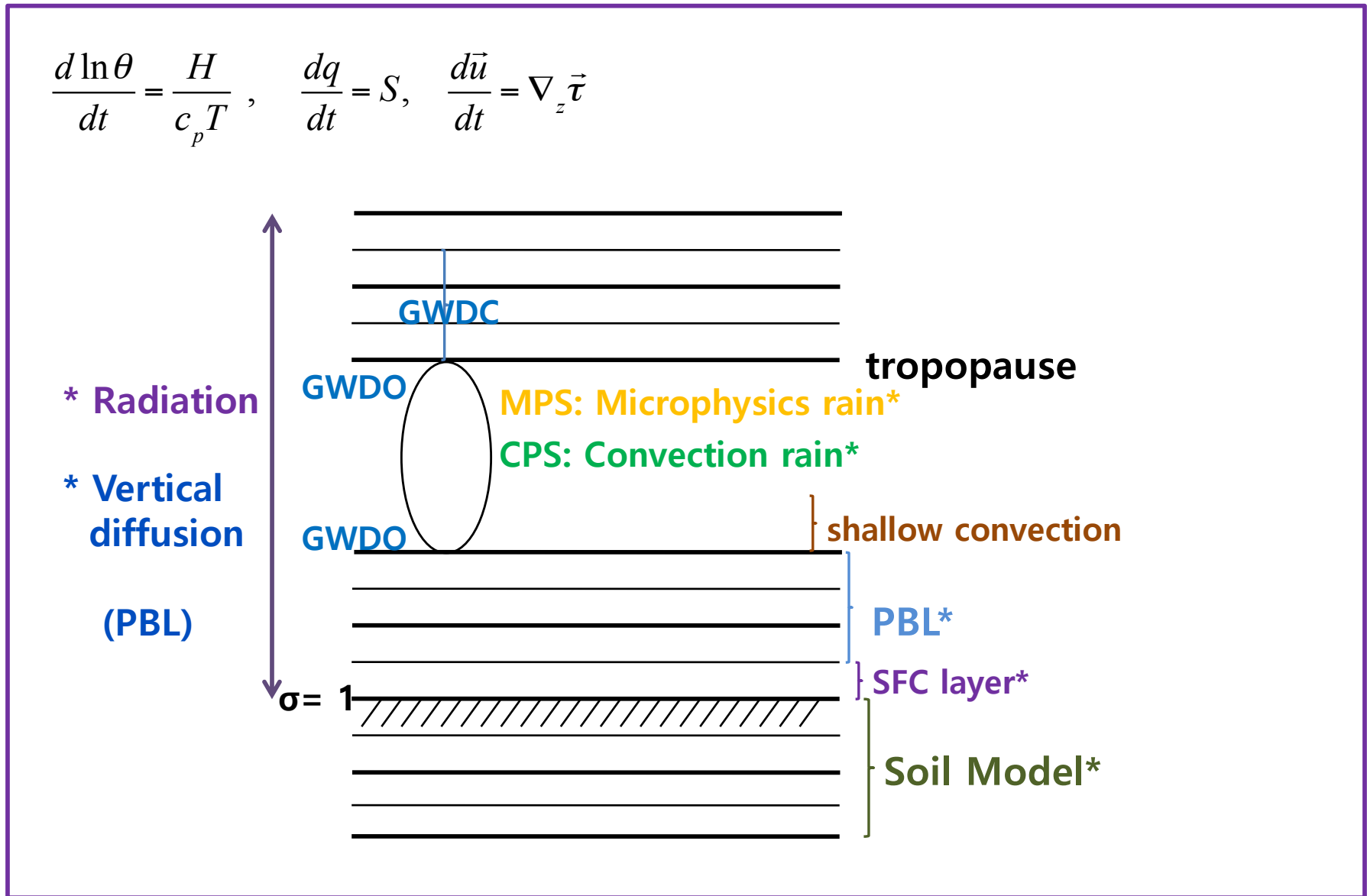
$$\frac{\partial \rho w q}{\partial t} = -\frac{\partial \rho u w q}{\partial x} + \dots$$

taking Reynolds averaging,

$$\frac{\partial \overline{\rho w'q'}}{\partial t} = \frac{\partial \overline{\rho w'w'q'}}{\partial z}$$

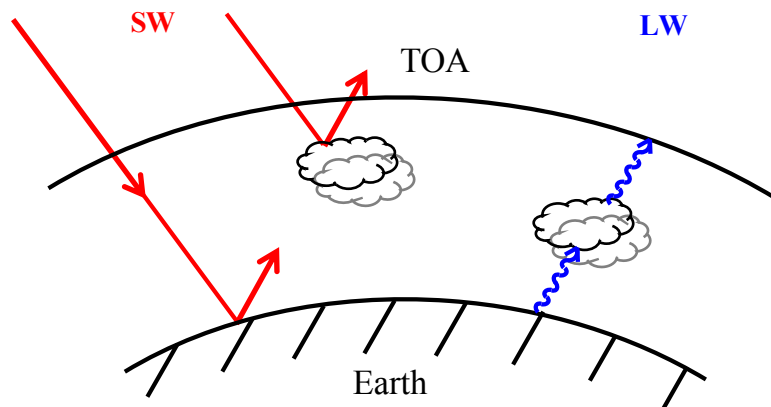
$$-\rho \overline{w'w'q'} = K' \frac{\partial \overline{\rho w'q'}}{\partial z} : \text{2nd order closure}$$

### 3) Schematic of physics algorithm : In modeled atmosphere : 7\* (essential) ~10



# 1. Radiation

## 1.1 Concept



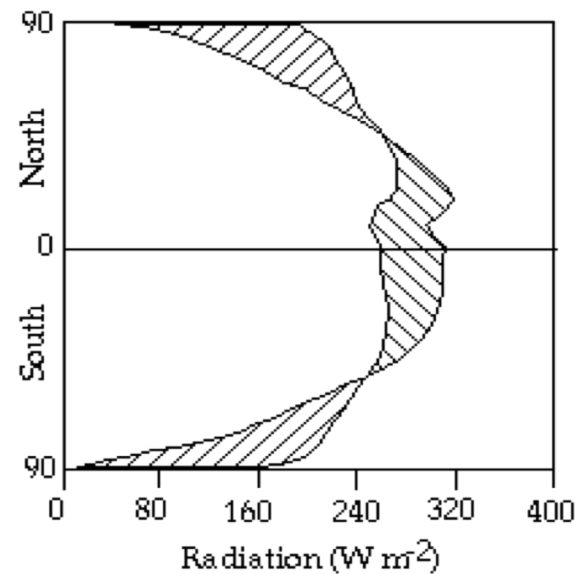
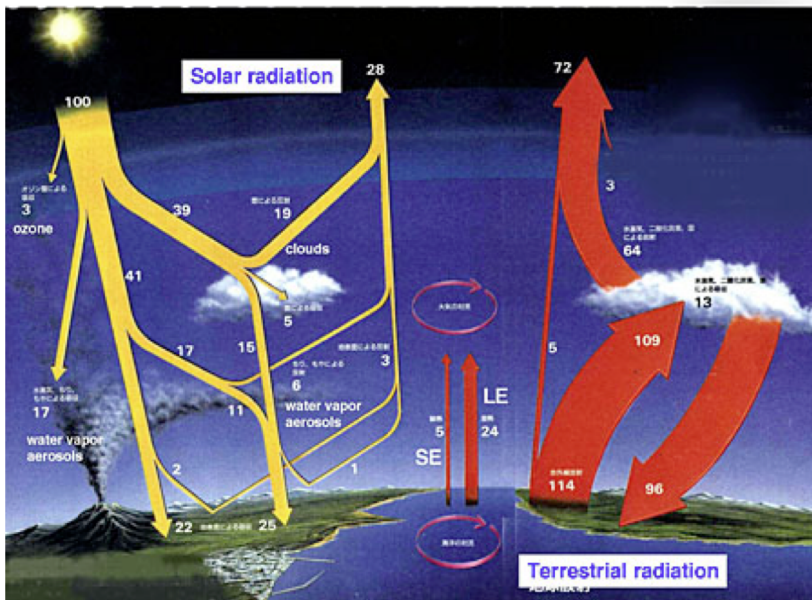
TOA (top of the atmos) :  $S = 1360 \text{ Wm}^{-2}$

Mean Flux :  $\frac{S}{4} = 340 \text{ Wm}^{-2} \rightarrow$  Energy source for Earth

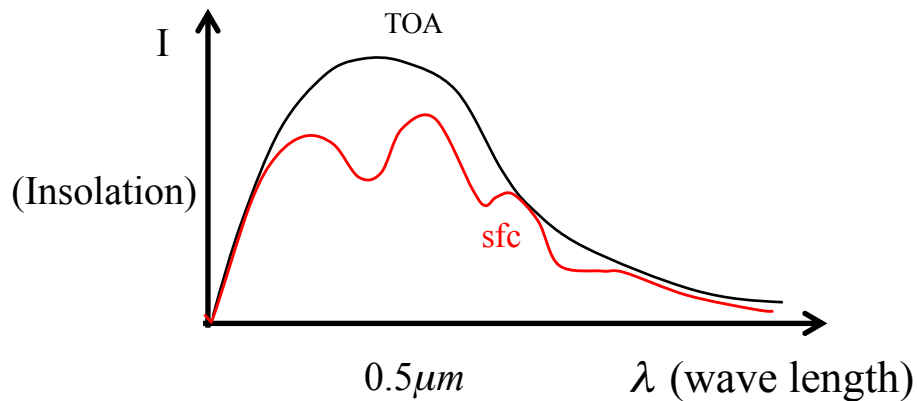
30% : reflected from the atmosphere clouds

$\rightarrow$  Back to space by terrestrial infrared radiation

$\left\{ \begin{array}{l} 25\% : \text{absorbed in the atmosphere} \\ 45\% : \text{absorbed at the earth surface} \end{array} \right\}$



## 1.2 Solar radiative transfer



- At TOA,

$$F = S \left( \frac{dm}{d} \right)^2 \cos \theta_0 \quad (\theta_0 : \text{Zenith angle})$$

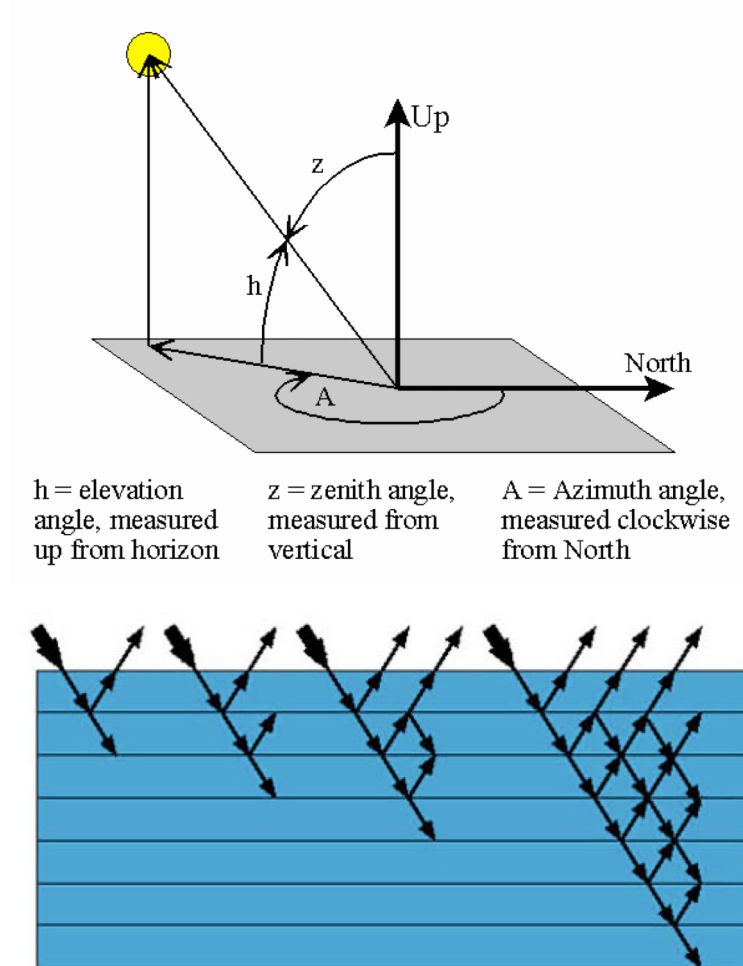
- Basic equations  $\mu = \cos \theta$

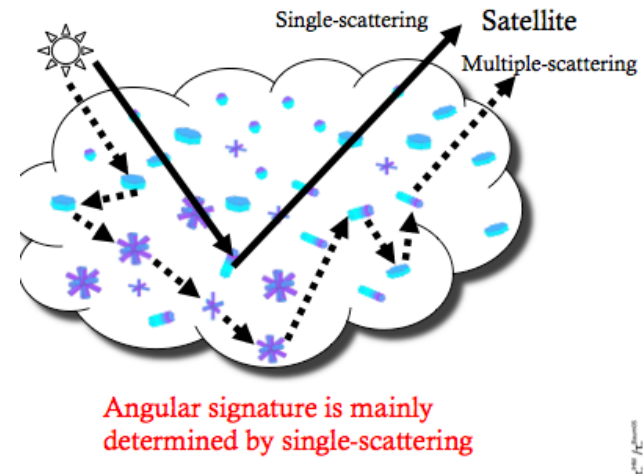
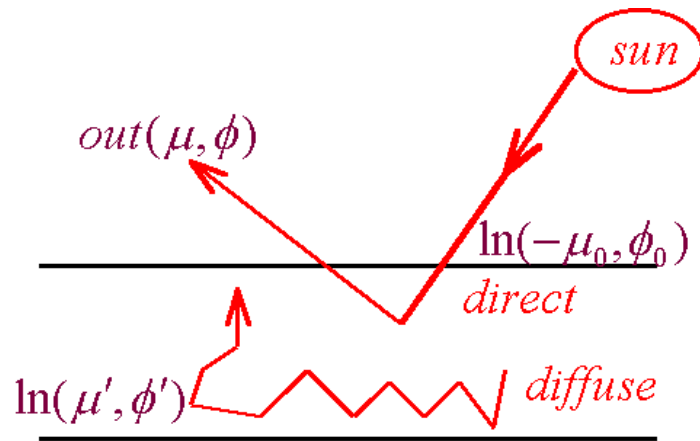
$$\mu \frac{dI(\tau, \mu, \phi)}{d\tau} = I(\tau, \mu, \phi) - J(\tau, \mu, \phi)$$

specific intensity      source function  
(absorption + scattering)

$$d\tau = -k_v \rho_a dz \quad \tau(\text{optical depth}) = \int_z^{z_\infty} k_v(z') \rho_a(z') dz' = \int_0^p k_v(p') q(p') \frac{dp'}{g}$$

Absorption coeff.





$$J = J(\tau, \mu, \phi) = \frac{\tilde{\omega}}{4\pi} \int_0^{2\pi} \int_{-1}^1 I(\tau, \mu', \phi') P(\mu, \phi; \mu', \phi') d\mu' d\phi [\text{diffuse (multiple) scattering}]$$

$$+ \frac{\tilde{\omega}}{4\pi} F_0 P(\mu, \phi; -\mu_0, \phi_0) e^{-\frac{\tau}{\mu_0}} [\text{single(direct) scattering}]$$

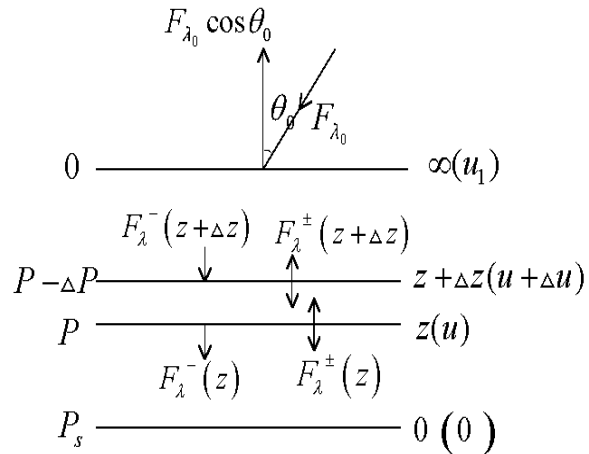
$$\left( \begin{array}{l} P : \text{Scattering phase function : redirects } (\mu', \phi') \rightarrow (\mu, \phi) \\ \tilde{\omega} = \frac{\sigma_s}{\sigma_e} : \text{Scatteing albedo} \\ \text{scattering cross section/extinction(scattering + absorption) cross section} \end{array} \right.$$

\* remove  $\phi$  dependency using  $P(\cos \theta)$  function

\*  $P$ ,  $\tilde{\omega}$ , Albedo depend on  $\lambda$ , particle size & shape.

$$P(\cos \phi) = \sum_{l=0}^N \tilde{\omega}_l P_l(\cos \phi) \quad : \text{Legendre Polynomial}$$

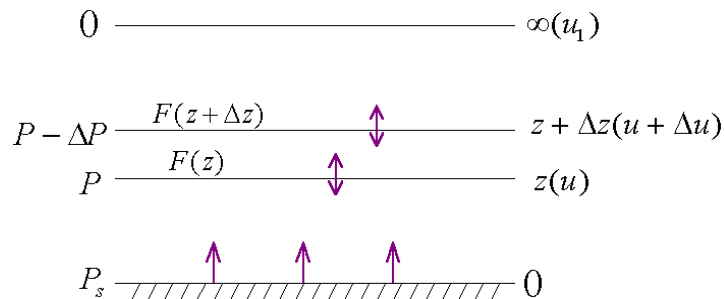




Radiative transfer equation solver. *simplification is needed*

- { Discrete - ordinates method  
*Two - Stream* and Eddington's approximation  
 Delta - function adjustment and similarity principle  
 Delta - Four stream approximation

## 1.3 Terrestrial radiation

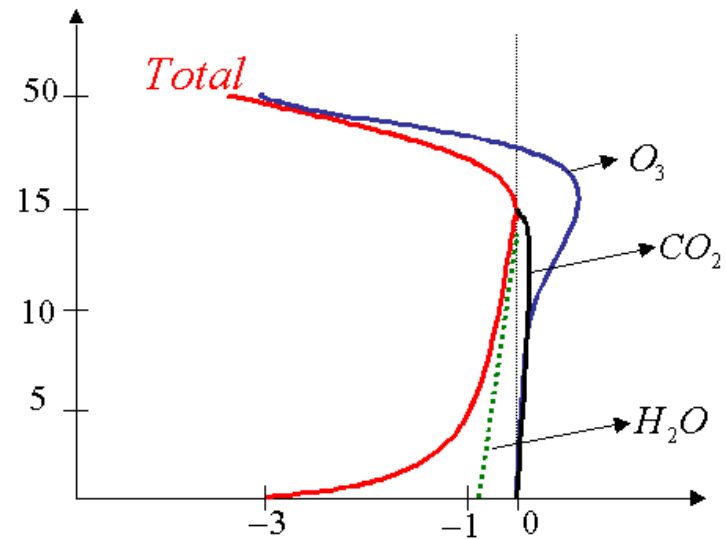
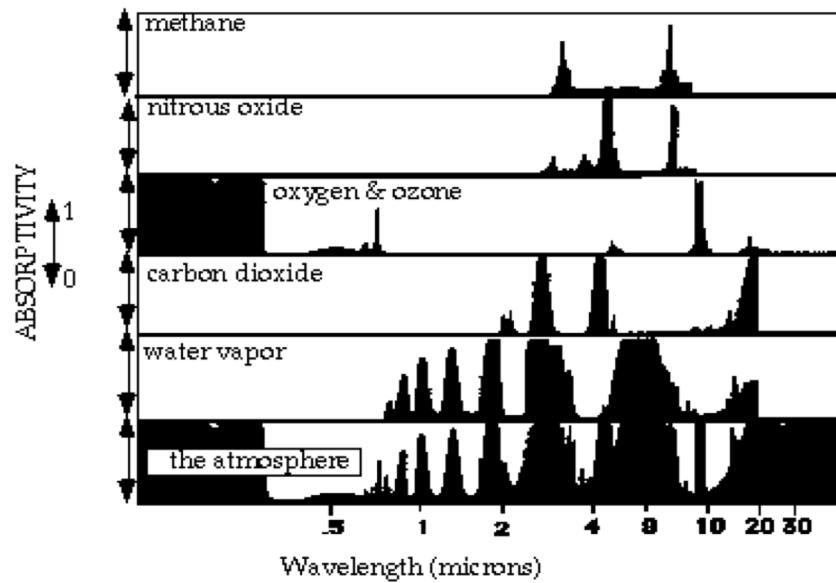


$0, \mu$	_____	$\tau_{\infty}, z_{\infty}, 0$
$\tau', \mu'$	.....	$\tau', z', P'$
$\tau, \mu$	_____	$\tau, z, P$
$\tau', \mu'$	.....	$\tau', z', P'$
$\tau, 0$	_____	$\tau_s, 0, P_*$

$$F(z) = F^{\uparrow}(z) - F^{\downarrow}(z)$$

$$\Delta F = F(z + \Delta z) - F(z)$$

$$\left. \frac{\partial T}{\partial t} \right|_{IR} = -\frac{1}{c_p \rho} \frac{\Delta F}{\Delta P} = -\frac{g}{c_p} \frac{\Delta F}{\Delta u}$$



In spectral bands (monochromatic)

$$\uparrow \mu \frac{dI_v(\tau, \mu)}{d\tau} = I_v(\tau, \mu) - B_v(T)$$

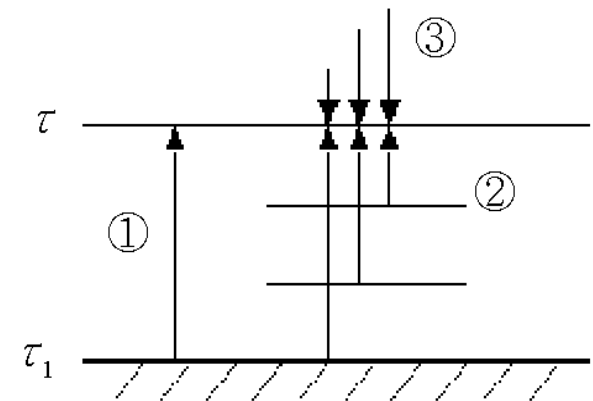
$$\downarrow -\mu \frac{dI_v(\tau, -\mu)}{d\tau} = I_v(\tau, -\mu) - B_v(T)$$

$$\text{B.C.} \quad \begin{cases} \text{SFC } (\tau = \tau_1), I_v(\tau, \mu) = B_v(T_s) \\ \text{TOP } (\tau = 0), I_v(0, -\mu) = 0 \end{cases}$$

$$F^\uparrow(\tau) = 2\pi B_v(T_s) \int_0^1 e^{-\frac{(\tau_1 - \tau)}{\mu}} \mu d\mu + 2 \int_0^1 \int_\tau^{\tau_1} \pi B_v[T(\tau')] e^{-\frac{(\tau' - \tau)}{\mu}} d\tau' d\mu$$

$$F^\downarrow(\tau) = 2 \int_0^1 \int_0^\tau \pi B_v[T(\tau')] e^{-\frac{(\tau - \tau')}{\mu}} d\tau' d\mu$$

$$d\tau = -k_v \rho dz \quad \tau_1 = \int_0^{u_1} k_v du, \quad u_1 = \int_0^\infty \rho dz \quad (\text{path length})$$



**LW is time consuming !**

## 1.4 Cloud fraction

**i) Earlier method (Slingo) :** when hydrometeor information is unavailable

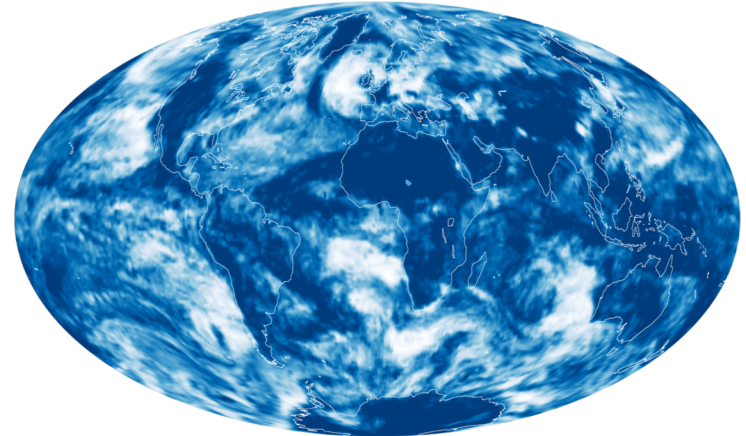
$$f = f_c(\text{convective cloud, CPS}) + f_l(\text{large-scale cloud, MPS})$$

$f_c$  : depends on precipitation,  $p_{\text{top}}$ ,  $p_{\text{bottom}}$

$$f_l : \text{depends on RH} = 1 - \left[ \frac{1 - RH}{1 - RH_0} \right]^{0.5}$$

where  $RH_0$  is the critical value of RH,

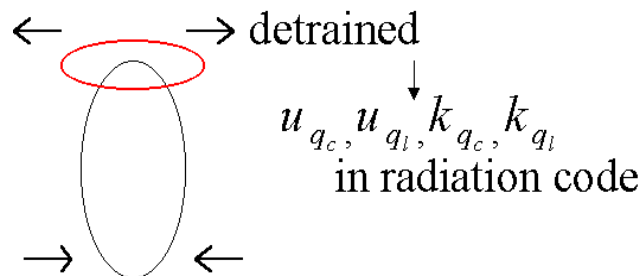
which is optimized based on observations



**ii) Advanced method (Chou)**

- inclusion of ice, liquid hydrometeors as prognostic water substances
- consistent treatment of water substance for both precipitation & radiative properties.

$f_c$  : uses information of detrained water substances from sub-grid scale clouds in convective parameterization scheme (CPS)



$f_l$  use the hydrometeor information from microphysics routine (MPS),  $q_c, q_s, q_i, \dots$

### iii) Radiation properties

$$\tau_i^c = cwp \left[ a_i + \frac{b_i}{r_{ei}} \right] f_{ice} \quad (\text{optical thickness})$$

$$w_i^c = 1 - c_i - d_i r_{ei} \quad (\text{co-albedo})$$

$$g_i^c = e_i - f_i r_{ei} \quad (\text{asymmetry factor})$$

$$f_i^c = (g_i)^2 \quad (\text{forward peak fraction})$$

a-f : coeff : depends upon band and k-

$$\overline{\tau_c} = \sum_i \tau_i \quad i : \text{each gas}$$

(The effective optical thickness for each spectral band)

- The long wave cloud emissivity (  $E_{cld}$  )

$$c'_f = E_{cld} c_f$$

$$E_{cld} = 1 - e^{-Dk_{abc}cwp}$$

where  $D = 1.66$  : diffusivity factor

$k_{abc} = k_l(1 - f_{ice}) + k_i f_{ice}$  : absorptivity coefficient

- **For diagnostic microphysics scheme** (hydrometeor is unavailable)

- cloud water scale height

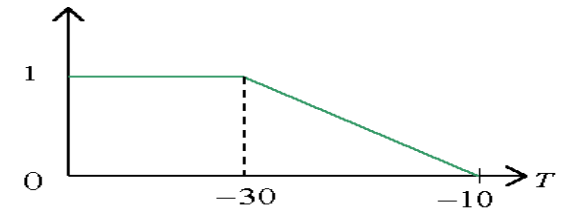
- cloud droplet size,

- ice fraction,  $f_{ice}$

$$h_1 = a \ln(1.0 + \frac{b}{g} \int_{P_T}^{P_S} q dp)$$

$$r_{ee} \begin{cases} = 10 \mu m & \text{over ocean} \\ < 10 \mu m & \text{over land} \end{cases} : \text{warm clouds}$$

$r_{ei}$ :  $10 \mu m$  (low)  $\sim 30 \mu m$  : (high) ice clouds



- **For prognostic microphysics scheme,**

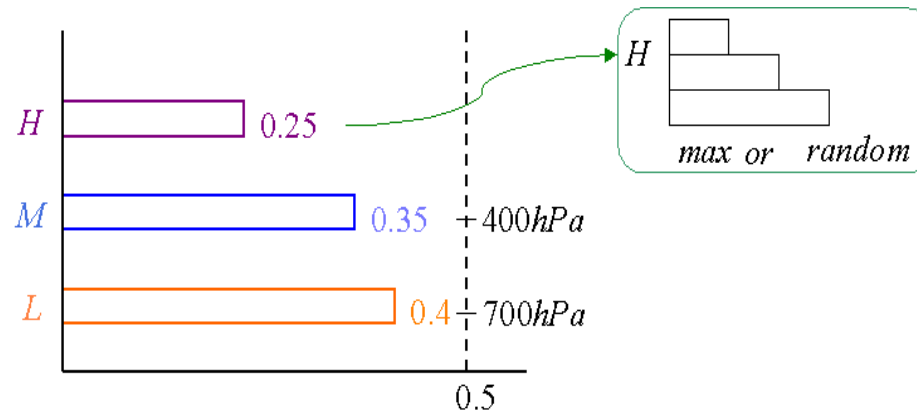
: Broadband radiation : Radiative properties are explicitly computed from prognostic water substances

: Simplified radiation : Dudhia (1989)

$$\alpha_p (\text{absorption coefficient}) = \frac{1.66}{2000} \left( \frac{\pi N_0}{\rho_{rs}^3} \right)^{\frac{1}{4}} m^2 g^{-1} = \begin{cases} 2.34 \times 10^{-3} m^2 g^{-1} & \text{for snow} \\ 0.33 \times 10^{-3} m^2 g^{-1} & \text{for rain} \end{cases}$$

$$u_p (\text{effective water path length}) = (\rho q_{rs})^{\frac{3}{4}} \Delta z \times 1000 \quad g m^{-2} \rightarrow \tau_p (\text{transmission}) = \exp(-\alpha_p u_p)$$

#### iv) Cloud overlapping



Maximum overlapping : 0.4

Minimum overlapping : 1.0

Random overlapping :  $H + (1 - H)M + \{1 - H - (1 - H)M\}L = 0.6$

#### - Computation :

$\tau$  is scaled by  $A_c$  (cloud cover) at a given layer.



- Flux for each of  $A_c, (1 - A_c) \rightarrow$  summation

- **Issues :**  $A_c = 0$  or  $1$   $\leftarrow$  in WRF versus partial cloudiness in GFS : WRF has partial cloudiness recently

#### - Interaction :

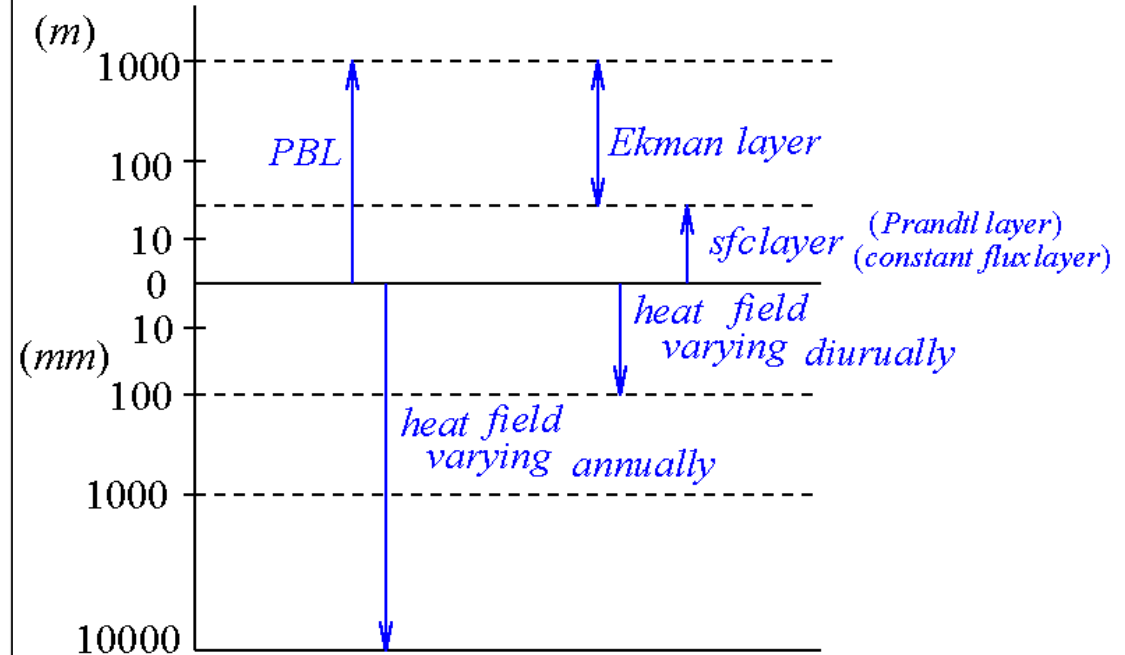
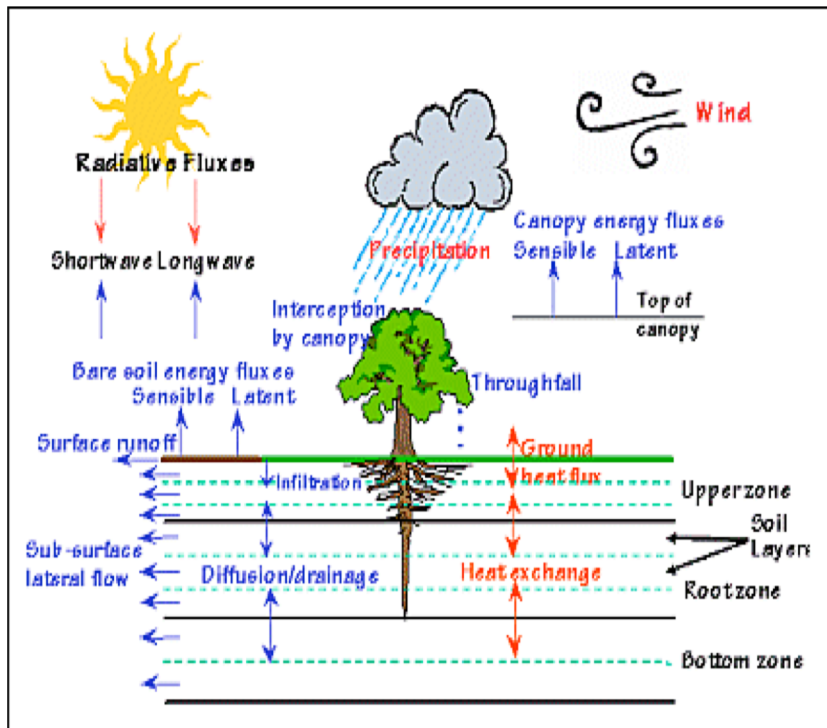
Radiation



Cloud

## 2. Land-surface processes

### 2.1 Concept : Surface layer + soil model



**Atmospheric surface layer** : the lowest part of the atmospheric boundary layer (typically about a tenth of the height of the BL ) where mechanical (shear) generation of turbulence exceeds buoyant generation or consumption.

Turbulent fluxes and stress are nearly constant with height.

➔ In atmospheric models, it is defined the height of the lowest model level.

## 2.2 Surface layer parameterization

Surface layer schemes calculate **friction velocities and exchange coefficients** that enable the calculation of surface heat and moisture fluxes by the land-surface models. These fluxes provide a lower boundary condition for the vertical transport done in the PBL Schemes.

Over water surfaces, the surface fluxes and surface diagnostic fields are computed in the surface layer scheme itself. Sea surface temperature can be predicted by the surface energy budget and mixed layer mixing

### 1) Bulk method (before 1990, MM4)

$$H_0 = \rho C_p C_H |\vec{V}_a| \Delta T$$

$$E_0 = \rho L C_H |\vec{V}_a| \Delta q M_a$$

$$\vec{\tau}_0 = \rho C_D |\vec{V}_a| \vec{V}_a$$

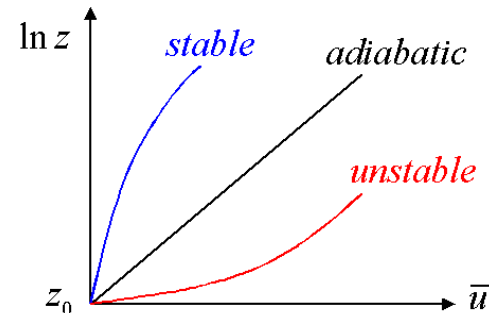
$C_D, C_H$  : *prescribed*

= 0.01 over land, = 0.001 over water

### 2) Monin-Obukov similarity

$$\frac{k_z}{u_*} \frac{\partial u}{\partial z} = \phi_m(z/L), \quad \frac{k_z}{u_*} \frac{\partial \theta}{\partial z} = \phi_t(z/L)$$

$$\text{Integrate,} \quad F_m = \int_{z_0}^{h_s} \frac{dz}{z} \phi_m = \ln\left(\frac{h_s}{z_0}\right) - \psi_m(h_s, z_0, L)$$



$$\vec{u} = C_u \left[ \ln \frac{z}{z_0} + \Phi \right]$$

$\Phi$  : curving factor

※ **Profile function** :  $\phi_m$  and  $\phi_t$

Dyer and Hicks formula for similarity

- unstable ( $L < 0$ )

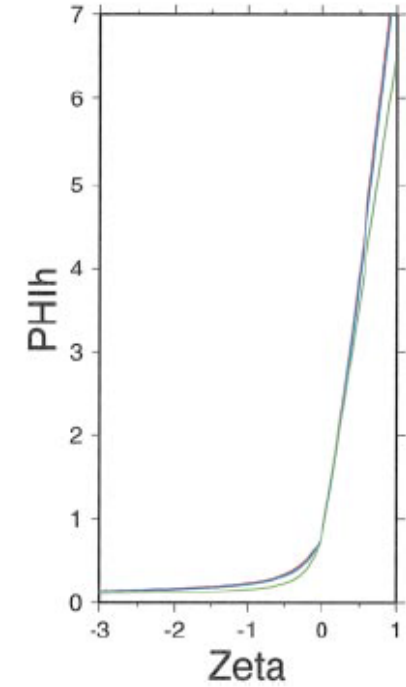
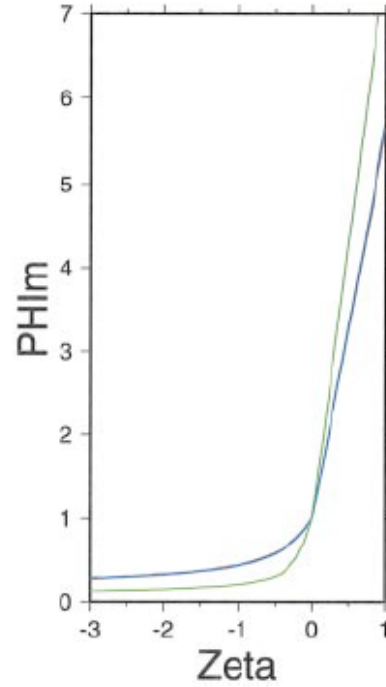
$$\phi_m = \left(1 - 16 \frac{0.1h}{L}\right)^{-\frac{1}{4}} \quad \text{for } u, v$$

$$\phi_t = \left(1 - 16 \frac{0.1h}{L}\right)^{-\frac{1}{2}} \quad \text{for } \theta, q$$

- stable ( $L > 0$ )

$$\phi_m = \phi_t = \left(1 + 5 \frac{0.1h}{L}\right)$$

$$\text{where } L = u_*^2 \bar{\theta} / (kg\theta_*) = -\frac{\rho C_p \theta_0 u_*^3}{kgH_0}$$



**Kantha  
(2003)**

$$\frac{h_s}{L} = \frac{\phi_m^2(hs/L)}{\phi_t(hs/L)} Ri = \zeta \quad (\text{Zeta} = hs/L)$$

$$\text{where } Ri = \frac{g}{\theta} \frac{\partial \theta}{\partial z} \left(\frac{\partial u}{\partial z}\right)^{-2}$$

※ **Useful relation** :

$$\text{Given the } F_m, F_H \quad C_D = k^2 / F_m^2, C_Q = C_H = k^2 / (F_m F_t) \quad u_* = kU / F_m$$

$$\tau_0 = \rho k_m \frac{du}{dz} = -\overline{u'w'} = \rho C_d U^2$$

$$H_0 = -\rho C_p k_h \frac{d\theta}{dz} = \rho C_p \overline{\theta'w'} = -\rho C_p C_H U \Delta \theta$$

$$E_0 = -\rho L \overline{q'w'} = -\rho L C_q U \Delta q$$

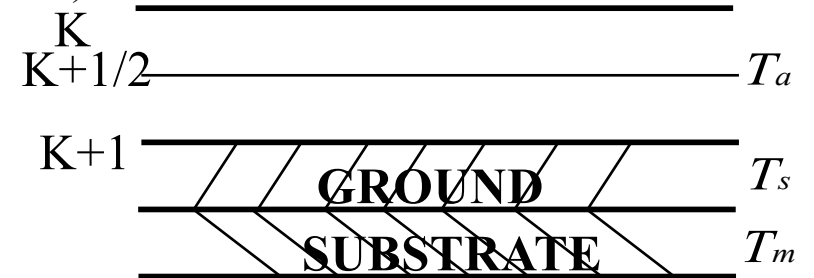


## 2.3 Soil model

1) **Slab model** : force-restore method (before 1990 : MM4)

$$\frac{\partial T_s}{\partial t} = \lambda_T (R_n - LE - H) - \frac{2\pi}{\tau} (T_s - T_a)$$

$$\frac{\partial T_s}{\partial t} = \frac{1}{\tau} (T_s - T_m) \quad : T_m, \text{ daily mean}$$



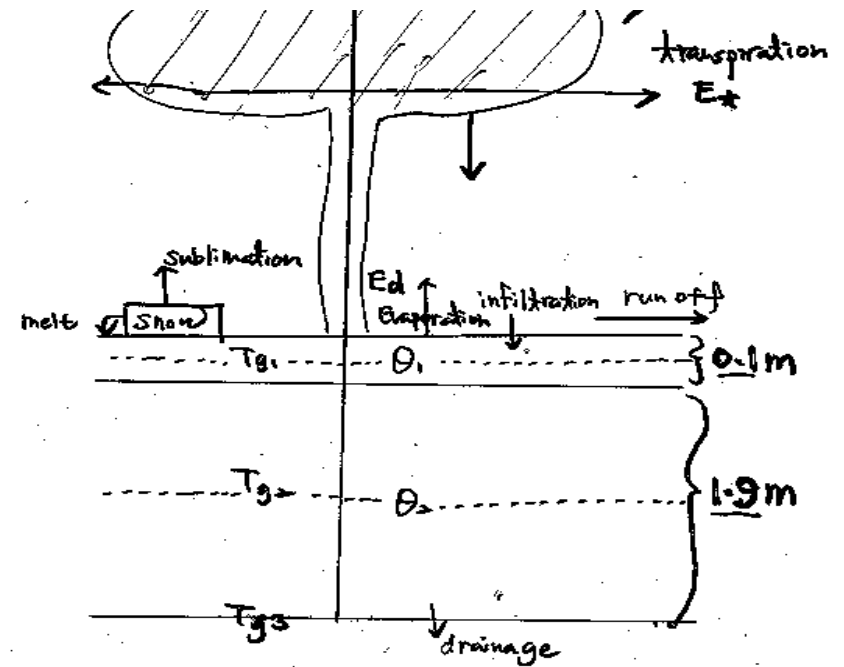
2) **Multi-layer model** : OSU method

$$\frac{\partial T_s}{\partial t} = \lambda_T (R_n - LE - H) \quad : \text{surface T}$$

$$(\rho C)_i \frac{\partial T_g}{\partial t} = \frac{\partial}{\partial z} \left( \lambda T_g \frac{\partial T_g}{\partial z} \right) \quad : \text{soil T}$$

$$\frac{\partial \Theta}{\partial t} = \frac{\partial}{\partial z} \left( D \frac{\partial \Theta}{\partial z} \right) + K \frac{\partial \Theta}{\partial z} + F_\Theta \quad : \text{soil moisture}$$

- NOAH, SIB, PLACE, VIC, CLM, etc

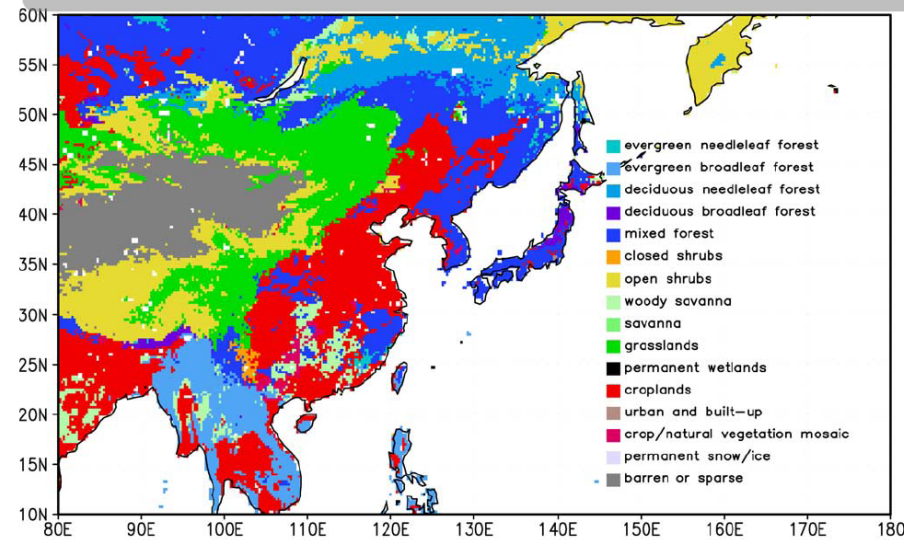


RA analysis: 2-layer soil model

## 2.4 Vegetation type $\rightarrow z_0$ , Albedo

### Vegetation types

#### MODIS (satellite)

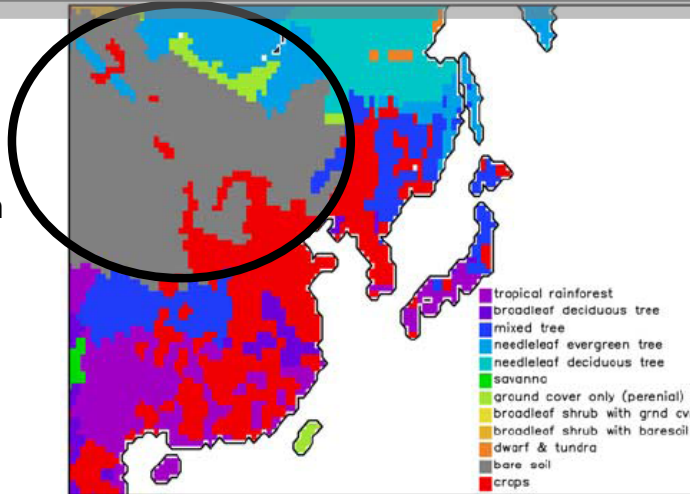


13 type data set  
1 degree

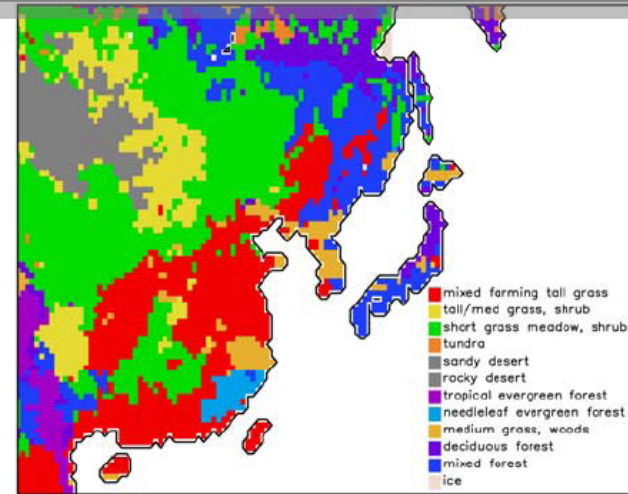
12 type data set  
20 min

#### The Simple Biosphere model (SiB)

quite broad  
desert region



#### United States Geological Survey's (USGS)



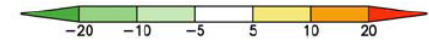
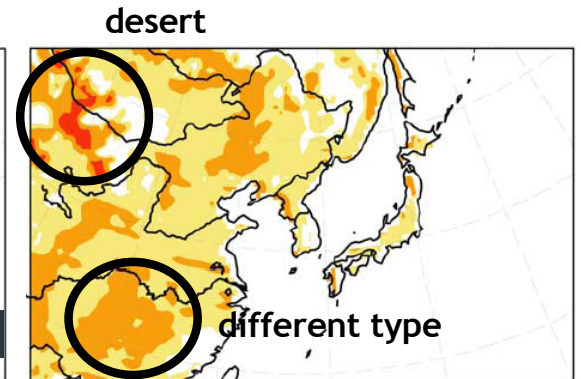
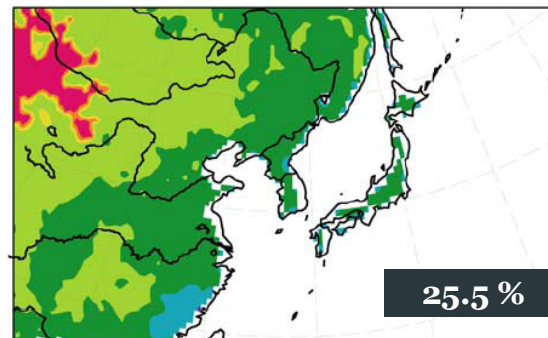
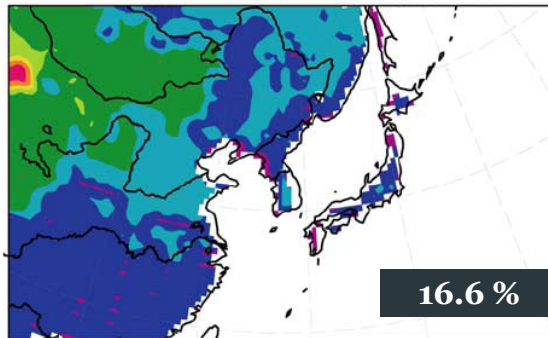
## Albedo and roughness length ( $z_0$ )

Sib

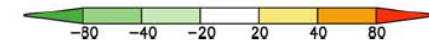
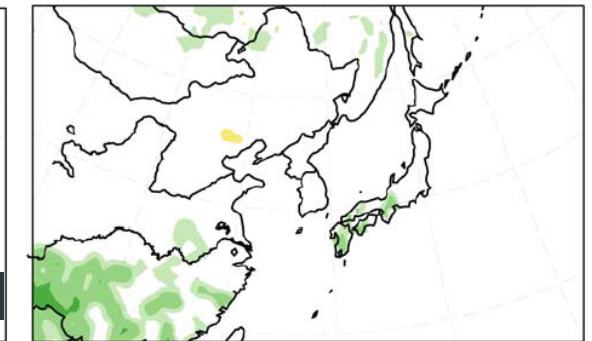
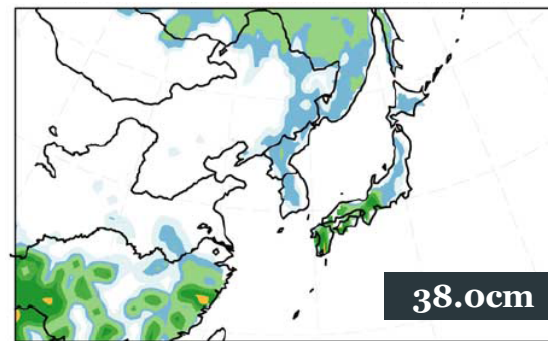
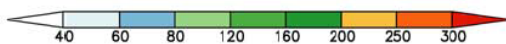
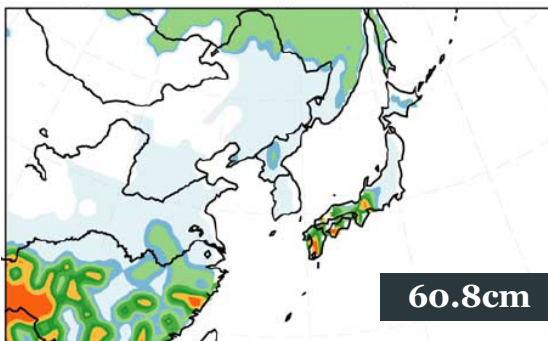
USGS

USGS-Sib

Albedo



Roughness length

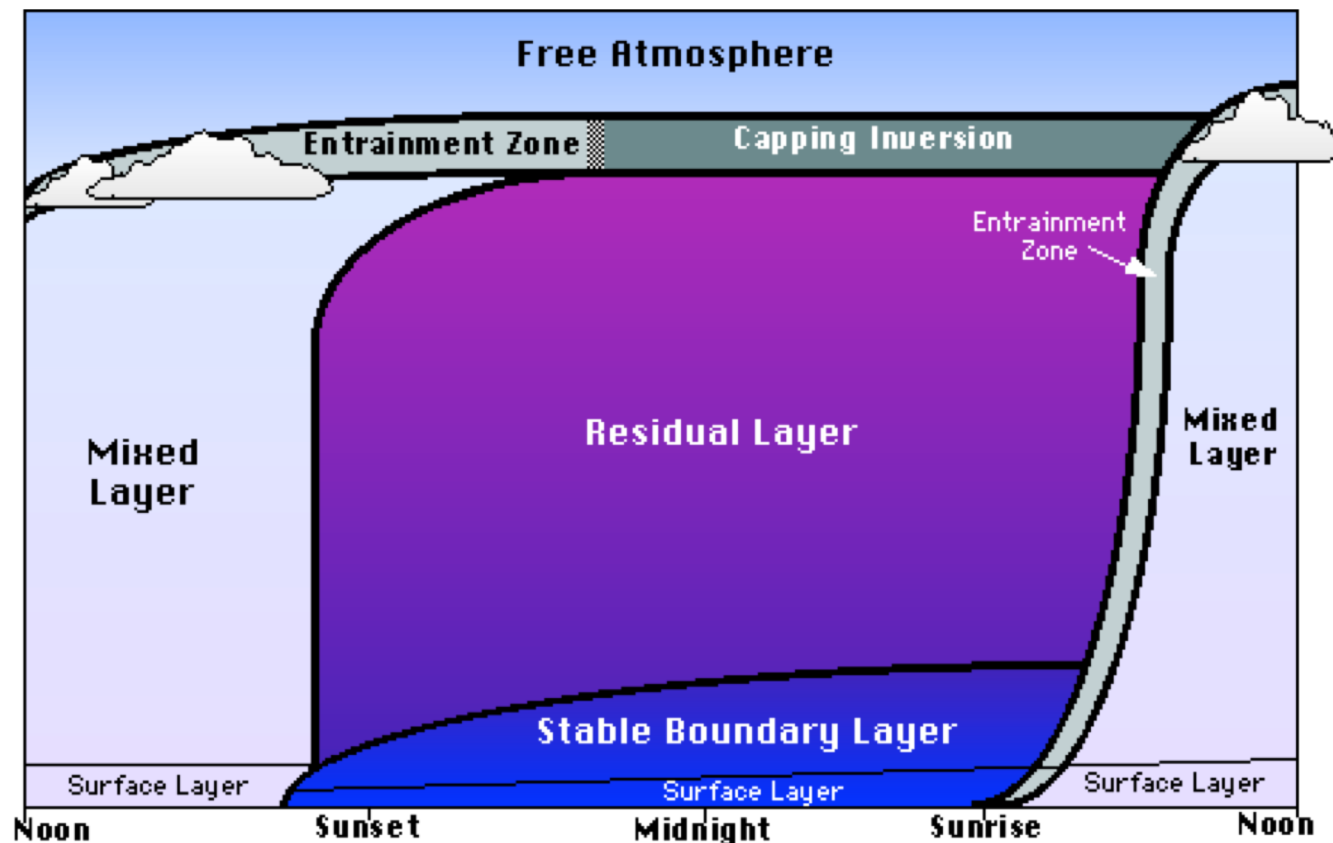


**Much uncertainties in calibration/optimization !!!**

### 3. Vertical diffusion (PBL)

#### 3.1 Concept

- computes the parameterized effects of vertical turbulent eddy diffusion of momentum, water vapor and sensible heat fluxes

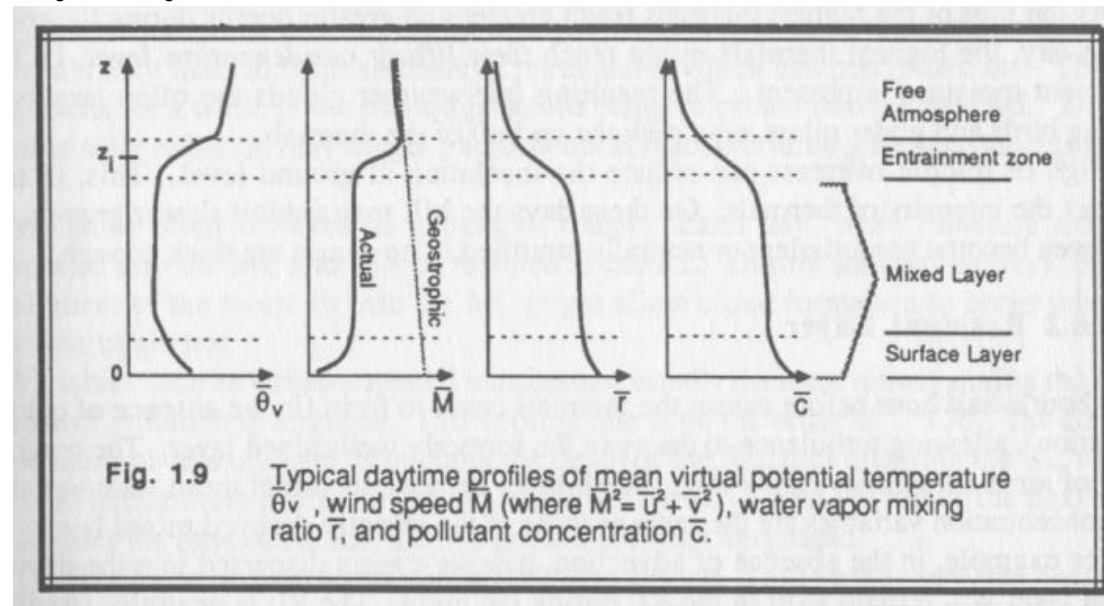


Structure of Planetary Boundary Layer

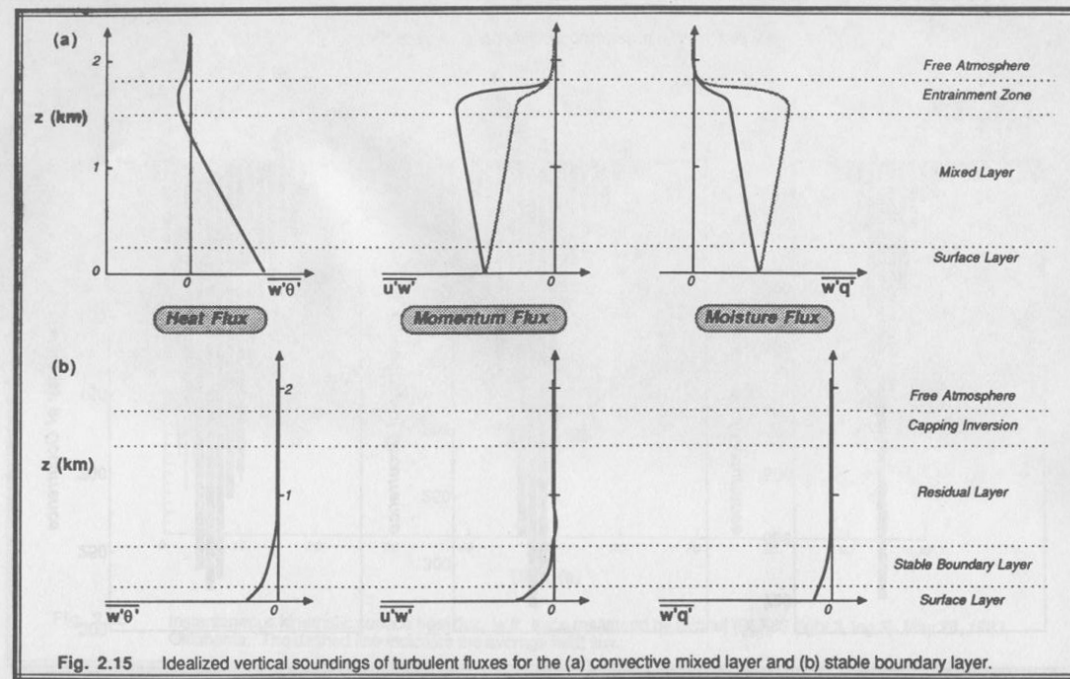
*adapted from Stull, 1988.*

### 3.2 Planetary Boundary Layer Structure : schematic

#### Daytime profiles



#### Daytime flux profiles



#### Nighttime flux profiles

### 3.3 Classifications : how to determine, $k_c$ , exchange coefficients

#### i) Local diffusion (Louis 1979)

$$\frac{\partial c}{\partial t} = \frac{\partial}{\partial z}(-\overline{w\overline{c}}) = \frac{\partial}{\partial z}(k_c \frac{\partial c}{\partial z}) \quad k_c : \text{ diffusivity, } k_m, k_t = l^2 f_{m,t}(Ri) \left| \frac{\partial U}{\partial z} \right|$$

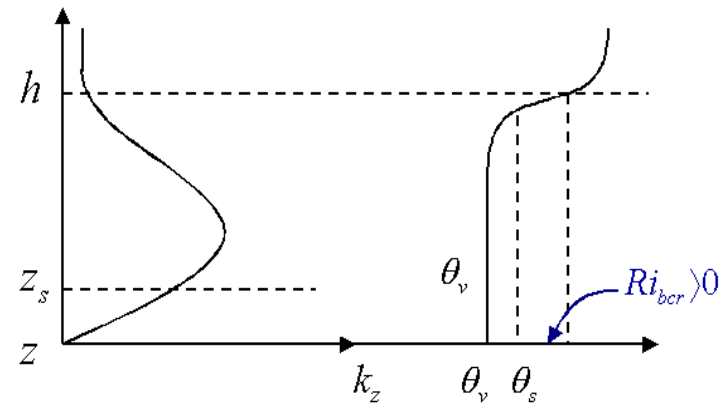
Local Richardson number

#### ii) Nonlocal diffusion (Troen and Mahrt 1986)

$$\frac{\partial c}{\partial t} = \frac{\partial}{\partial z}(-\overline{w\overline{c}}) = \frac{\partial}{\partial z}(k_c (\frac{\partial c}{\partial z} - \gamma_c))$$

$$k_{zm} = k w_s z (1 - \frac{z}{h})^p, \quad h = R_{ibcr} \frac{\theta_m}{g} \frac{U^2(h)}{(\theta_v(h) - \theta_s)}$$

$$\theta_s = \theta_{va} + \theta_T (= b \frac{(\overline{\theta_v' w'})_0}{w_s}), \quad w_s = u_* \phi_m^{-1}$$



#### iii) Eddy mass-flux diffusion (Siebesma et al. 2007)

$$\frac{\partial c}{\partial t} = \frac{\partial}{\partial z}(-\overline{w\overline{c}}) = -\frac{\partial}{\partial z} \left[ \underbrace{-k_c \frac{\partial \overline{c}}{\partial z}}_{\text{small eddies}} + \underbrace{M(c_u - \overline{c})}_{\text{strong updrafts}} \right]$$

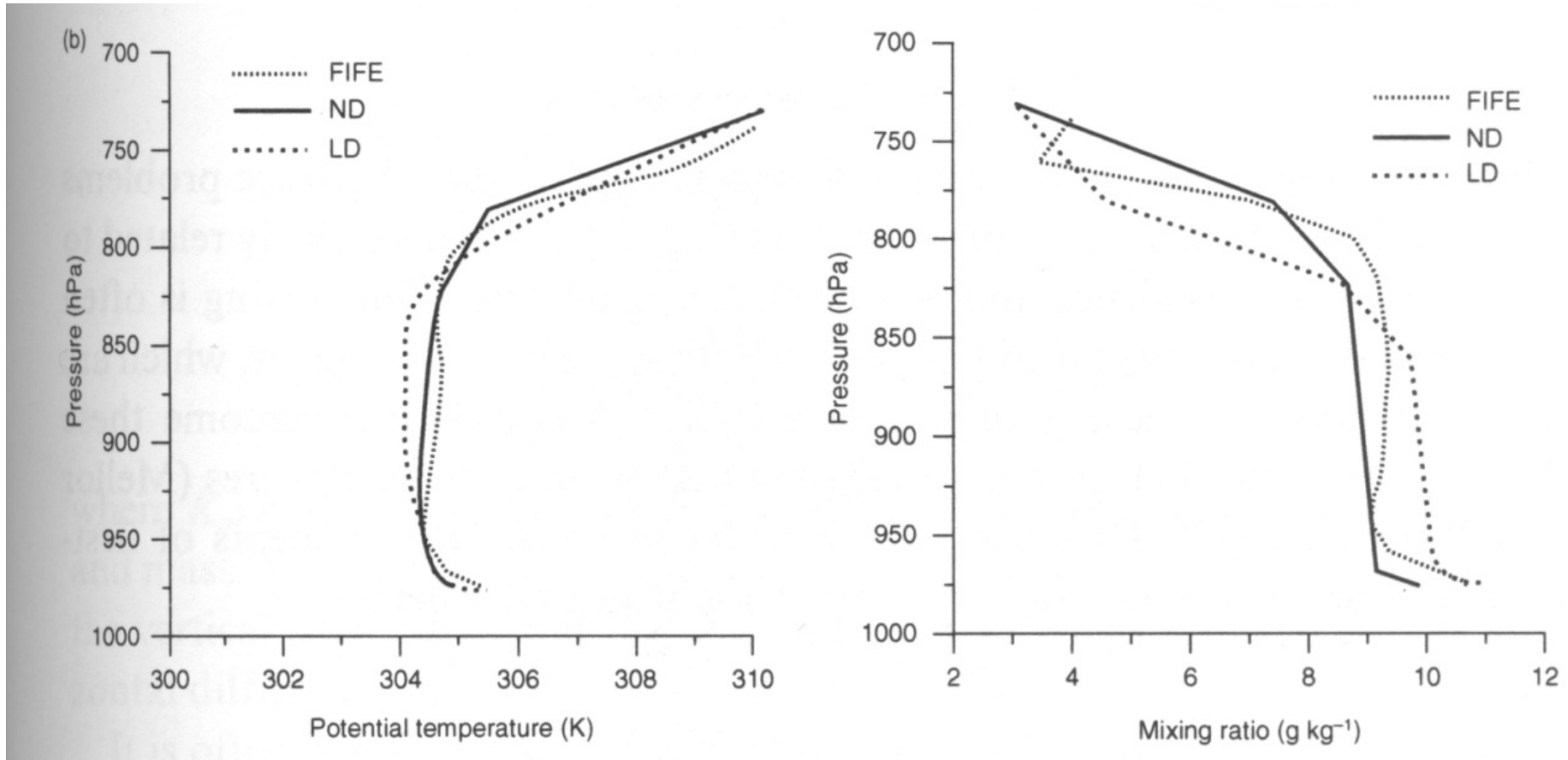
#### iv) TKE (Turbulent Kinetic Energy) diffusion (Mellor and Yamada 1982)

$$\text{TKE equation : } \frac{\partial \overline{u_i u_j}}{\partial t} + u_j \frac{\partial \overline{u_i u_j}}{\partial x_j} = -\frac{\partial}{\partial x_k} \left[ \overline{u_i u_j u_k} + \frac{1}{\rho} \dots \right]$$

$$\overline{u_i u_j} \implies k_z = \text{fn (TKE)}$$



### 3.4 Local versus nonlocal



Local scheme (LD) typically produces unstable mixed layer in order to transport heat upward. But, observation (FIFE) shows near neutral or slightly stable BL in the upper part of BL

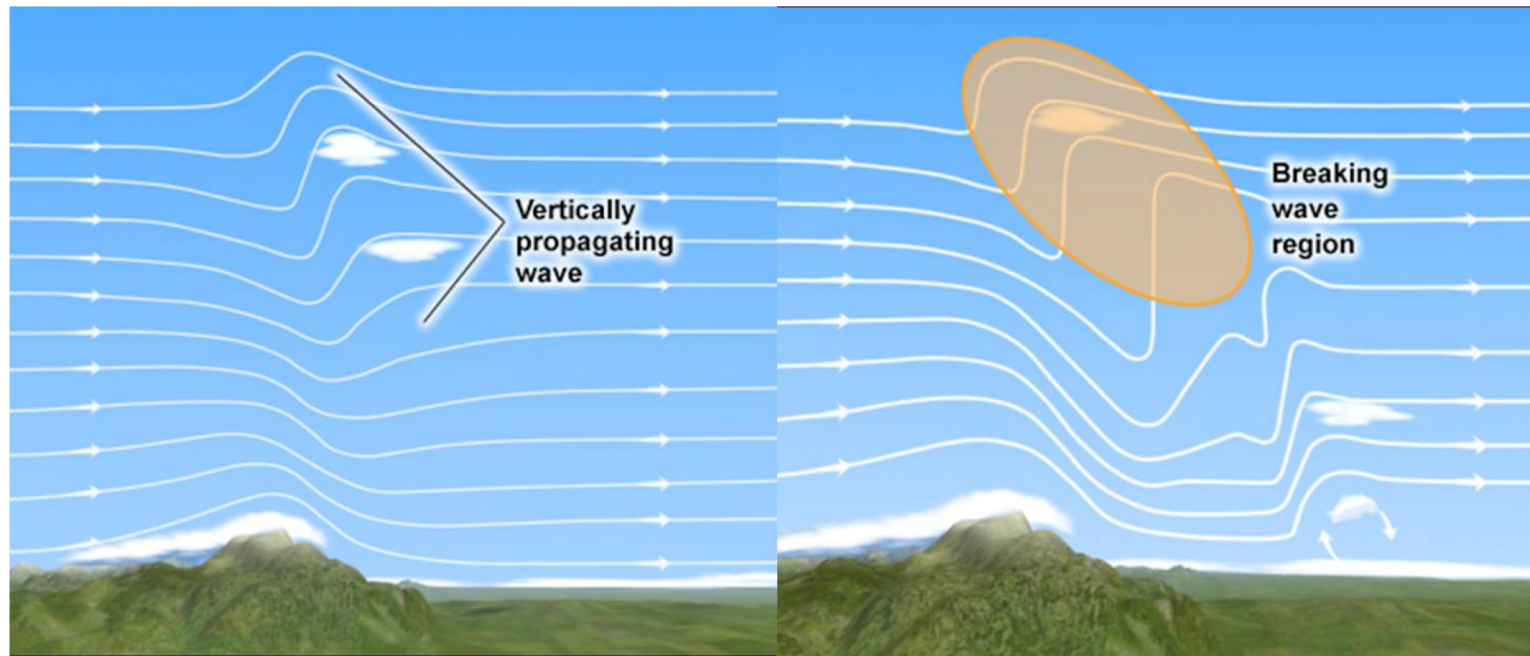
Hong and Pan (1996)

## 4. Gravity Wave Drag

- GWDO : GWD induced by sub-grid scale orography
- GWDC : GWD induced by precipitating deep convection

### 4.1 Concept

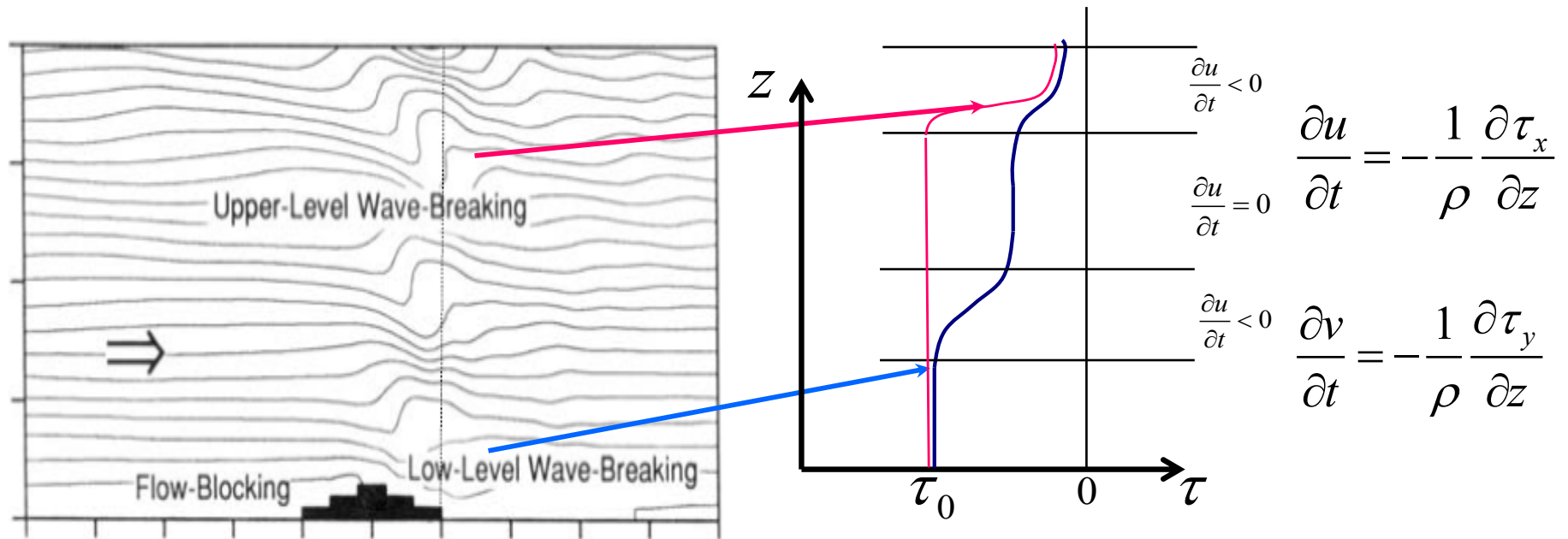
This scheme (GWDO) includes the effect of mountain induced gravity wave drag from sub-grid scale orography including convective breaking, shear breaking and the presence of critical levels. Effects are strong in the **presence of strong vertical wind shear and thermally stable layer**.



\* In smoothed model orography, momentum stress in mountain cannot be generated



## 4.2 Enhanced lower tropospheric gravity wave drag (Kim and Arakawa 1995)



Stress at reference level :

$$\tau_0 = -E \frac{1}{\Delta x} \frac{\rho_0 U_0^3}{N_0} \frac{Fr^2}{Fr^2 + 0.5/OC}, \quad U_0 = \frac{1}{h} \int_{k=1}^{k=k_{pbl}} U dz$$

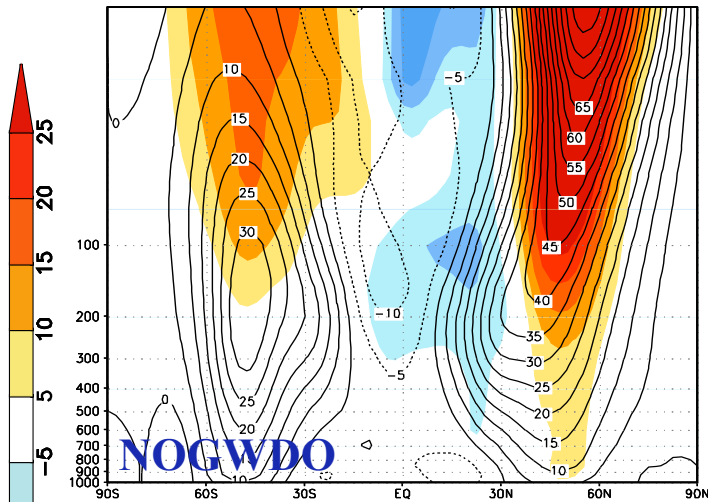
Reference level (KA95) : Max (2, KPBL)

**Earlier method** : the conventional Ri number-based wave- breaking mechanism using the saturation hypothesis, which works mainly in the upper atmosphere

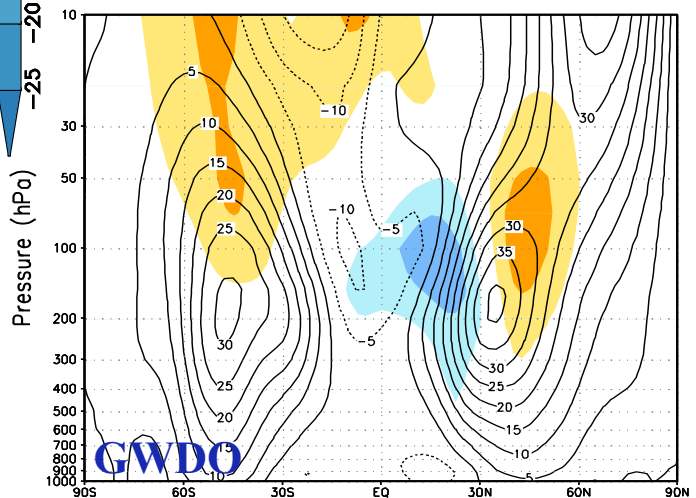
**Advanced**: the higher-moment orographic statistics-based wave- breaking mechanism using half-theory (Scorer parameter  $\sim BVF^{**2} / U^{**2}$ ) and half-empiricism obtained from mesoscale mountain wave simulations, which breaks in the lower atmosphere as well as upper layer, together with flow blocking

## 4.3 Impact of GWDO

Zonal-averaged zonal wind (96/97 DJF)

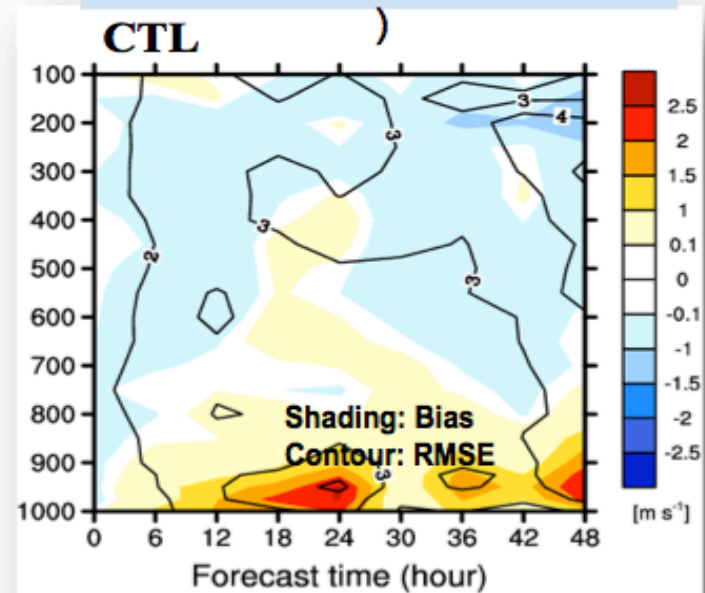


Contour : Zonal averaged zonal wind  
Shaded: Deviations from the RA2

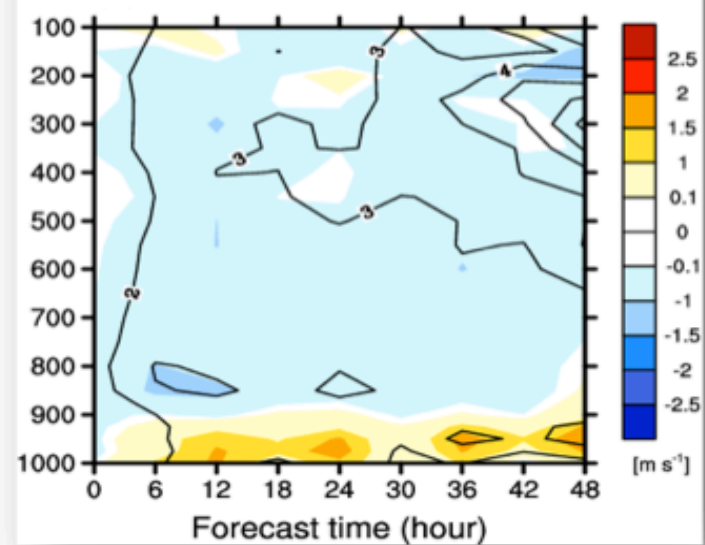


→ Improves the upper level jets  
(Kim and Hong, 2009)

Wind speed (3-5 Jan. 2010)



SOP



Flow blocking → Improves the low-level winds (Choi and Hong, 2015)

## 4.4 Non-orographic gravity wave drag (GWD)

### Conventional GWD parameterizations

- Simple assumptions on GW sources with either a **spatiotemporally uniform distribution** (e.g., Warner and McIntyre, 2001 in Unified Model; Scinocca, 2003 in IFS)

### Source-based GWD parameterizations

(e.g., Choi et al., 2018 in KIM  
Korean Integrated Model, Hong et al. 2018)

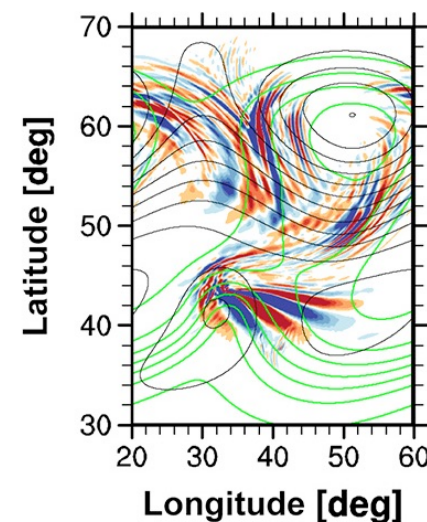
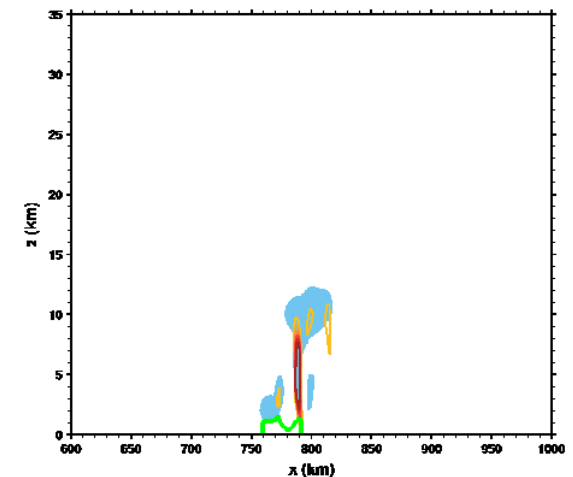
### *Convective GWD parameterizations*

- Chun and Baik (1998, 2002): **first to analytically formulate** the GW momentum flux spectrum for convective GWs
- Beres (2004), Song and Chun (2005): **non-stationary GW** parts together with spatiotemporal variations in the convective source
- Song and Chun (2008): **ray-based approach**
- Choi and Chun (2011): **moving speed of the convective source** and **wave-propagation direction**, were determined

### *Jet/Front GWD parameterizations*

- Not yet explored sufficiently
- Charron and Manzini (2002), Richter et al. (2010): introduced a **frontogenesis function** to diagnose the generation of frontal GWs
- de la Cámara and Lott (2015): based on the theoretical results of GW generation by **potential vorticity anomalies**

### Convective GWs



### Jet/front induced GWs

## 4.5 Improvement by source-based non-orographic GWD (Choi et al., 2018)

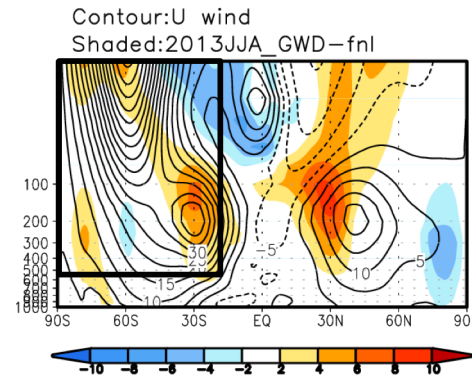
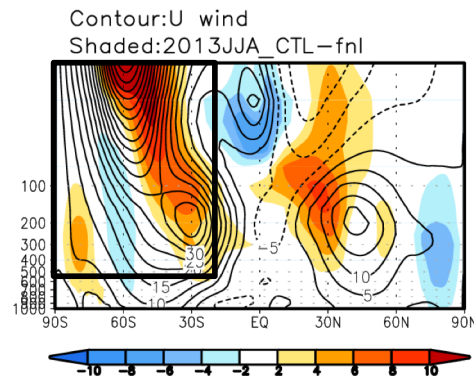
### Seasonal simulation

(2013JJA)

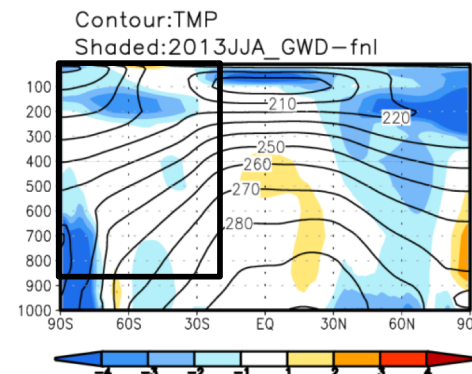
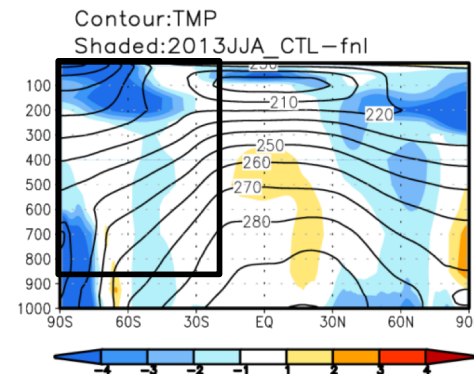
CTL

NEW

Zonal-mean  
zonal wind



Zonal-mean  
temperature



↑ Shading: Bias against the analysis

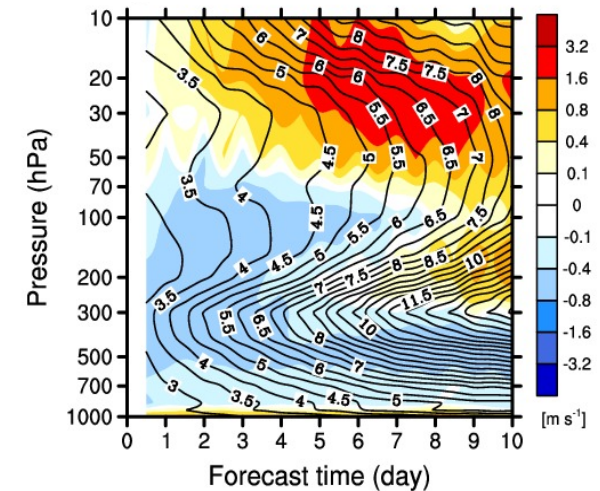
*This study is the first to investigate the effects of both the convective and frontal GWD schemes*

### Medium-range forecast

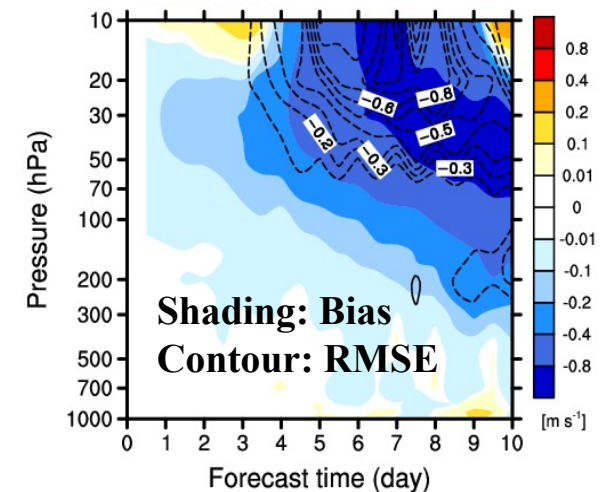
(January 2016)

Wind speed

CTL



NEW-CTL



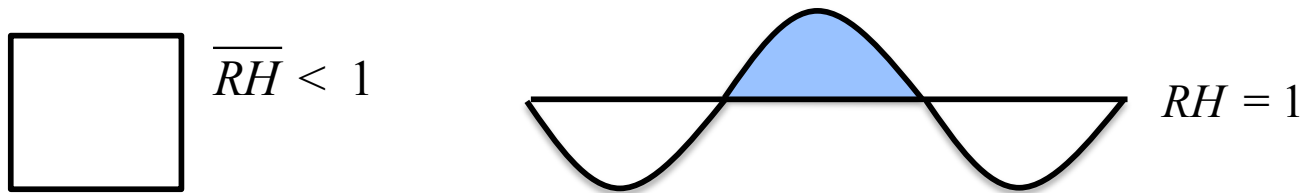
# Precipitation Processes

## Concept : precipitation algorithms (CPS and MPS)

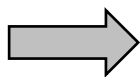
- In real atmosphere, dynamical motion  $\rightarrow RH > 1 \Rightarrow$  clouds form  $\rightarrow$  produces rain
- In modeled atmosphere, grid-mean  $RH < 1$

But generate clouds by sub-grid scale motion  $\rightarrow$  requires parameterized process  
(it is often related to the deep convection processes)

Deep convection : 2~10 km



$\Delta x \Rightarrow 0$  , more grid-resolvable precipitation



Thus, we need the cumulus parameterization scheme to account for releasing conditional instability due to subgrid scale motion

- Grid-resolvable (Microphysics scheme : MPS) : Supersaturation  $\rightarrow$  clouds
- Subgrid scale (Cumulus parameterization scheme : CPS) : CAPE removal  $\rightarrow$  clouds

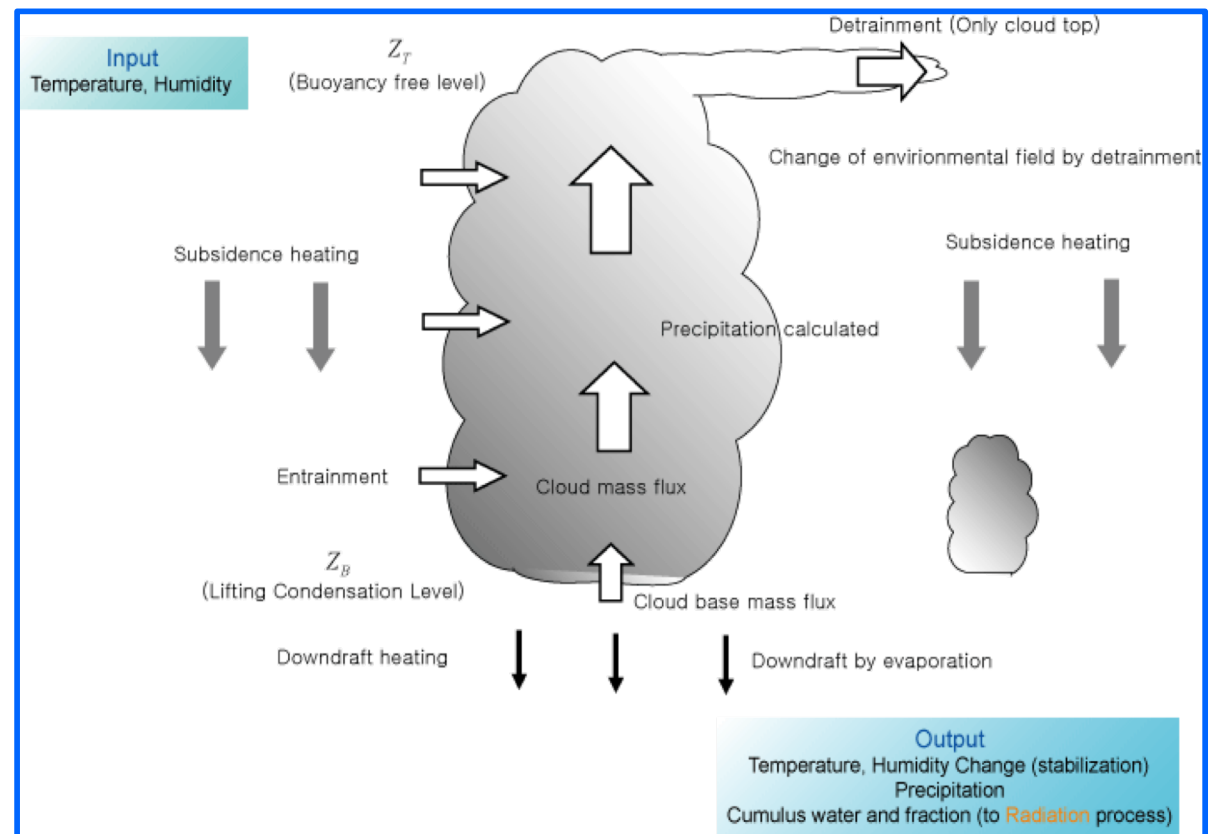


# 5. Cumulus parameterization scheme

## 5.1 Concept

- represents deep precipitating convection and feedback to large-scale
- must formulate the **collective effects of subgrid-scale clouds** in terms of the **prognostic variable in grid scale**

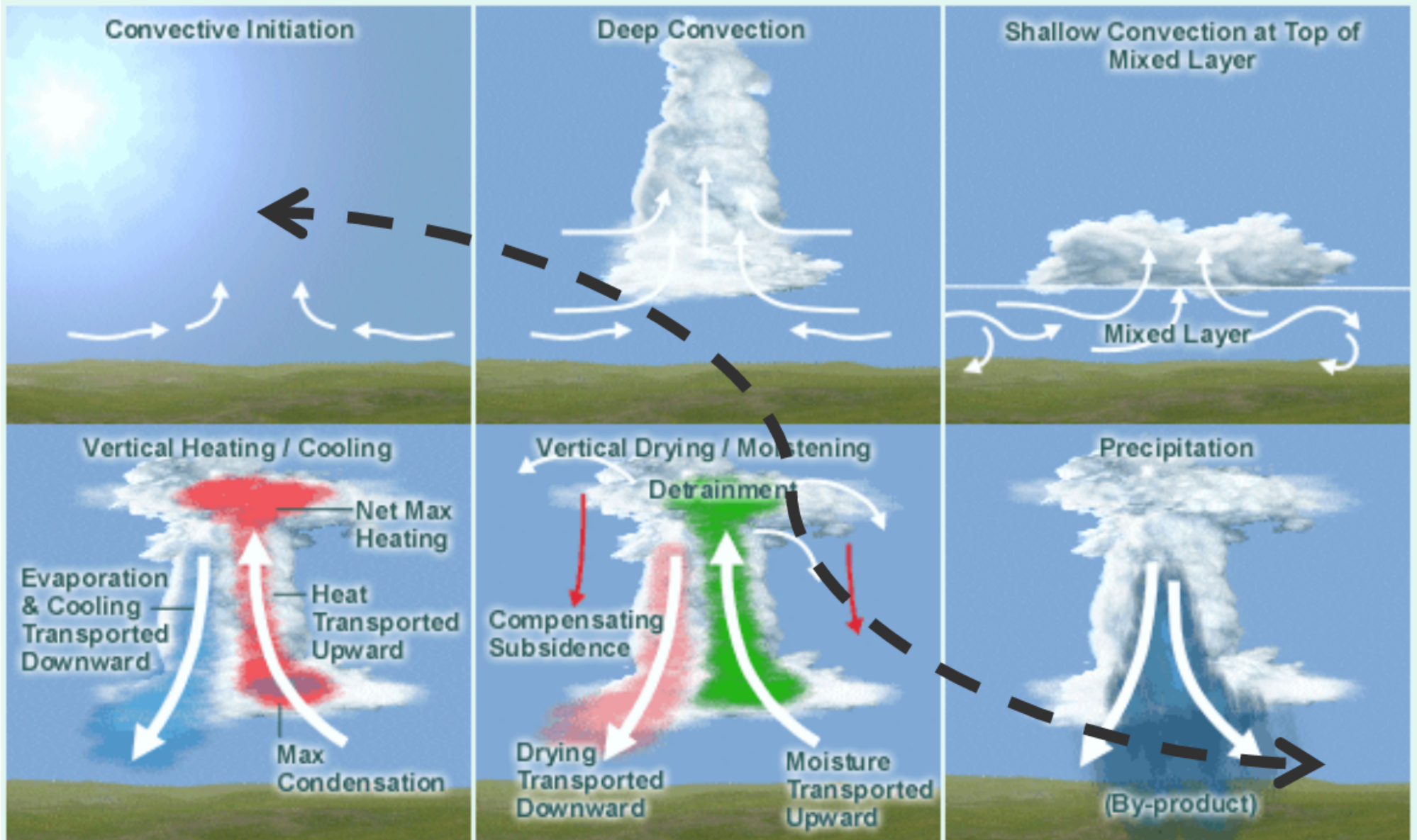
Parameterized convection  
Deep convection  
Subgrid-scale precipitation  
Implicit precipitation



Jul 25, 2010, Seoul

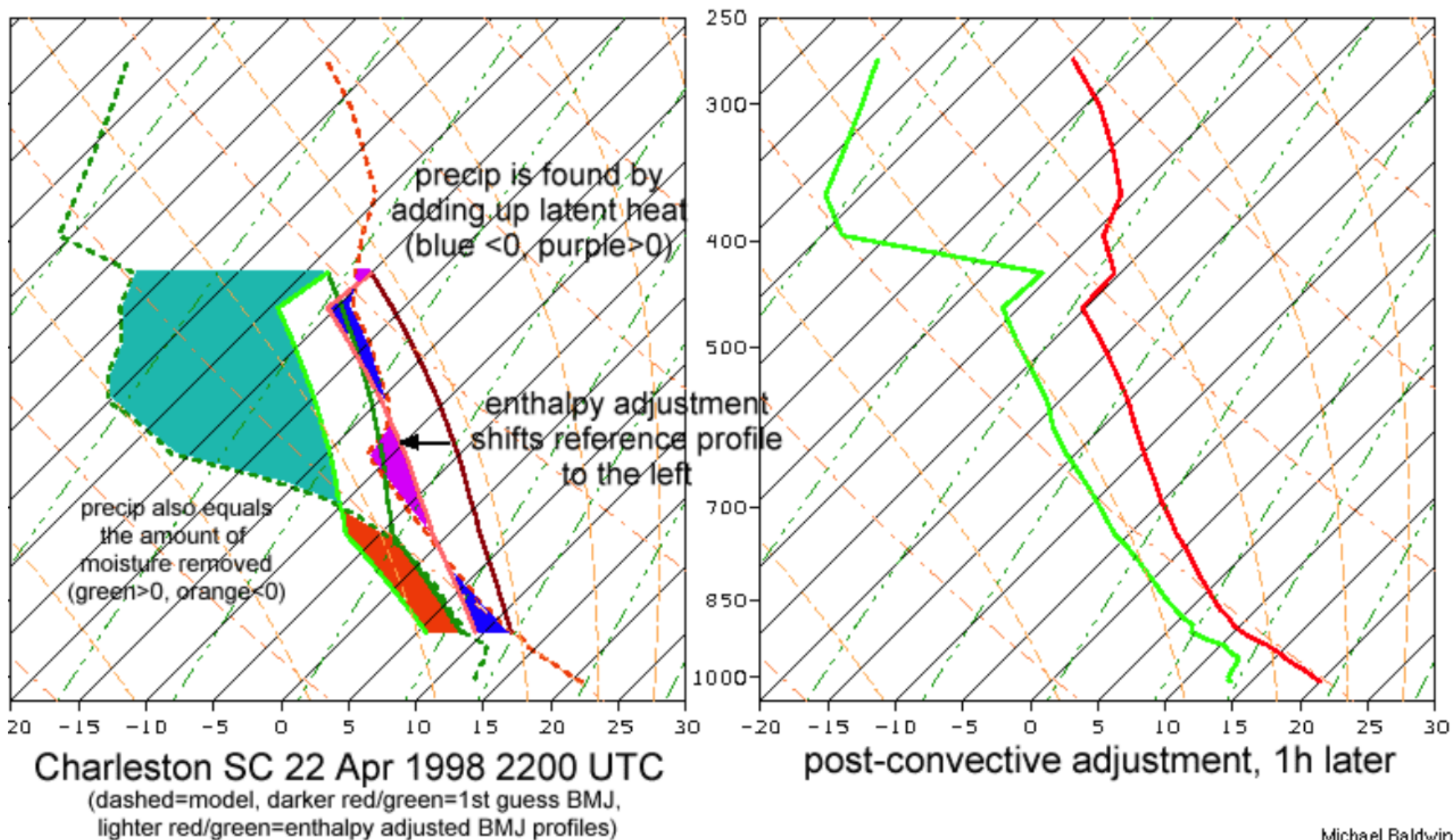
**b**

## Processes CP Schemes Need to Account For



The COMET Program

**CPS does not consider the detailed evolution of convection !**



Michael Baldwin

CPS consider changes in profile **before and after** convection



## 5.2 Kuo scheme (1965)

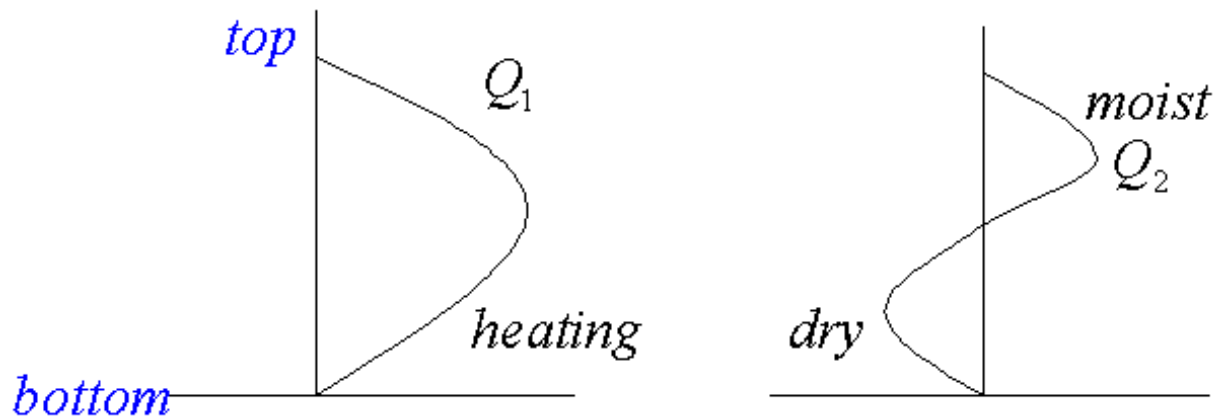
: Cloud is formed in proportional to **column-integrated moisture convergence**

$$M_t = -\frac{1}{g} \int_0^{P_s} \nabla \cdot (vq) dp + F_{g_s}$$

- Heating and moistening profiles (prescribed)

$$\frac{d\theta}{dt} = \frac{1}{\pi} [gL(1-b)M_t Q_1 + Q_r] \quad \frac{dq}{dt} = -g(1-b)M_t Q_2$$

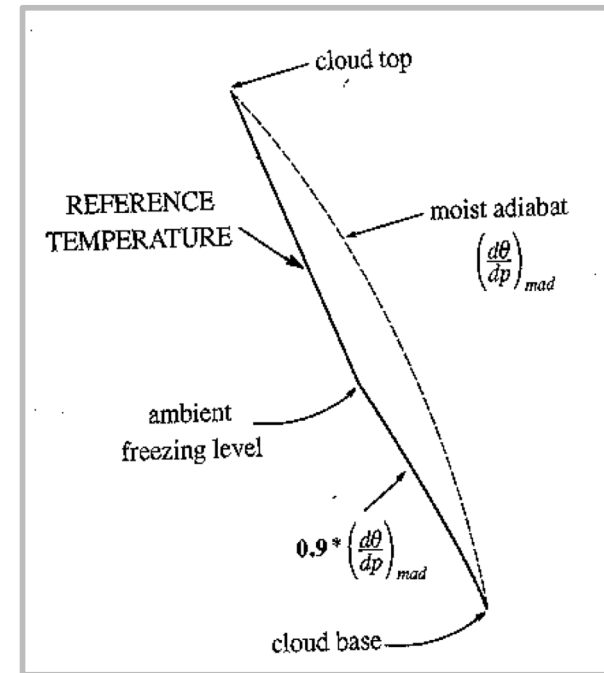
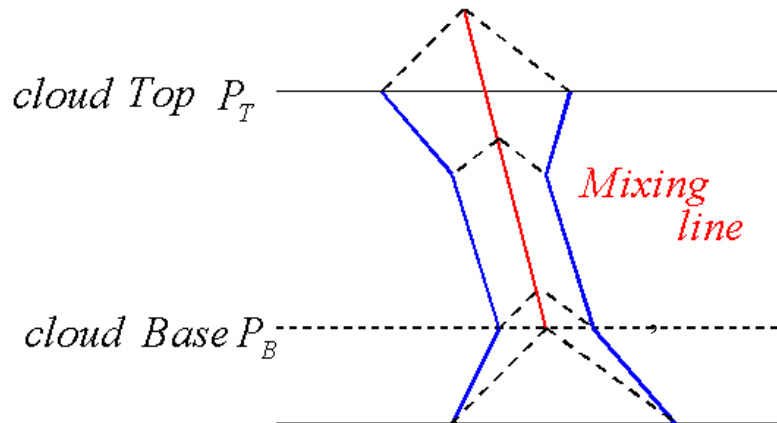
$$\int_0^{P_s} Q_1 dp = \int_0^{P_s} Q_2 dp = 1$$



- Kuo type scheme has been widely used before 1990

## 5.3 Betts-Miller scheme (1986)

: adjust **toward reference profiles** that are based on observational evidence of convective equilibrium



**Reference profile to be adjusted**

$$\theta'_R(P) = \bar{\theta}(P_B) + \beta M_\theta (P - P_B)$$

$$\frac{\partial q}{\partial p} = \beta \left( \frac{\partial q}{\partial p^*} \right)_M \quad \beta = \frac{\partial p^*}{\partial p}$$

$$M_\theta = 0.85 \left( \frac{\partial \theta^*}{\partial P^*} \right)_M$$

$$\text{Energy Constraints : } \int_{P_B}^{P_{T+1}} C_P (T_R - \bar{T}) dp = 0$$

**- Convective tendencies & Precipitation**

$$\left( \frac{\partial \bar{T}}{\partial t} \right)_{Cu} = \frac{T_R - \bar{T}}{\tau}$$

$$\left( \frac{\partial \bar{q}}{\partial t} \right)_{Cu} = \frac{q_R - \bar{q}}{\tau}$$

$$Precip = \int_{P_0}^{P_T} \left( \frac{q_R - \bar{q}}{\tau} \right) \frac{dP}{g}$$

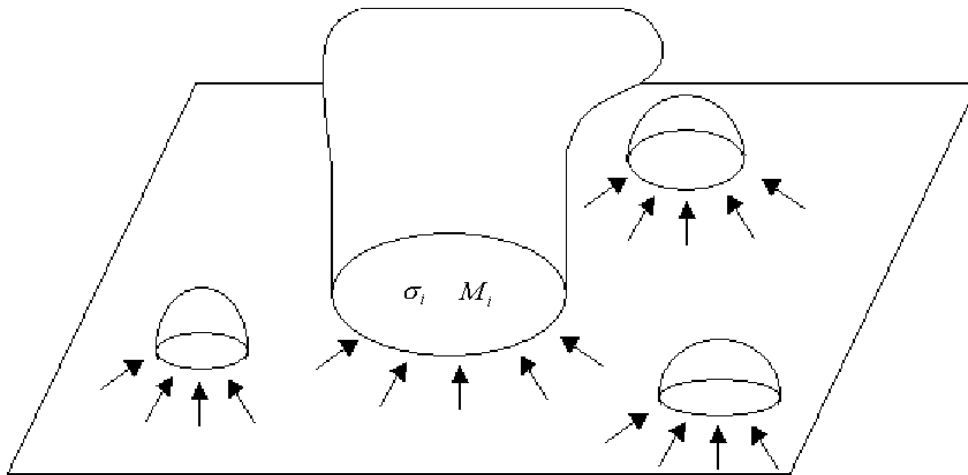
## 5.4 Mass-flux schemes : Arakawa-Schubert (1974)

### i) Concept

-- Mass flux approach, cloud ensemble, quasi-equilibrium

-- Theoretical frame work for CPS

- Area is large enough so that cloud ensemble can be a statistical entity
- Area is small enough so that cloud environment is approximately uniform horizontally



$M_i$  : vertical mass flux through ith cloud

$\sigma_i$  : fractional area covered by ith cloud

$M_c \equiv \sum_i M_i$  : total vertical mass flux

$\rho M = M_c + \tilde{M}_{\text{environment}}$   
: net mass flux/unit large-scale horizontal area

## ii) Quasi-equilibrium : cloud forcing $\sim$ large-scale adjustment

: CPS computes the warming (cooling) in the grid box due to adiabatic descent (ascent), rather than computing latent heat release in cloud models

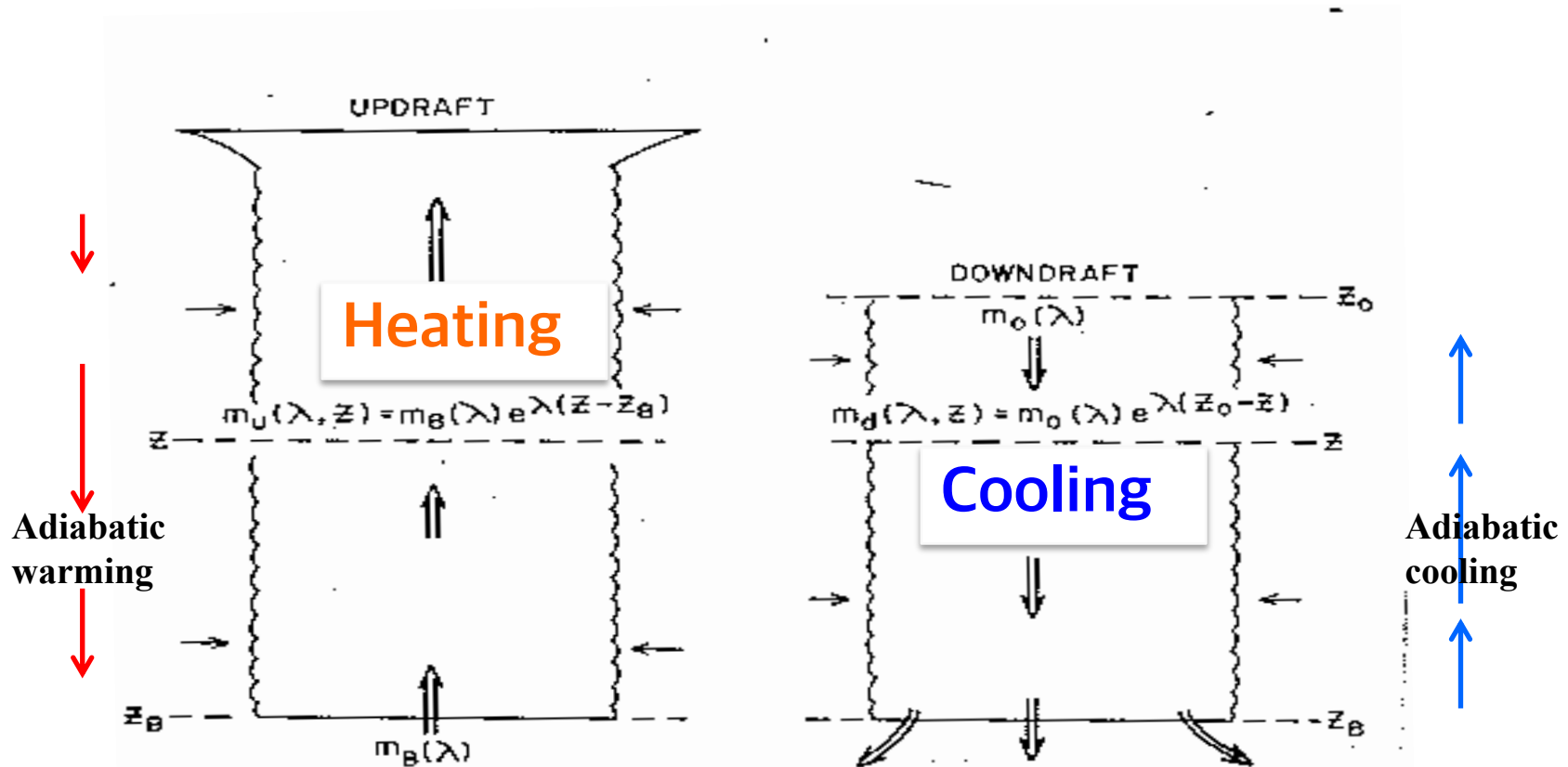


FIG. 4.10. Model for updraft and downdraft of cloud type  $\lambda$  (from Johnson 1976).

### iii) Energy budget equations

Large-scale flux across grid box



Exchange of S between environment and clouds



$$\frac{\partial}{\partial t} \rho (1 - \sigma_c) \tilde{s} = -\bar{\nabla} \cdot (\rho \tilde{V} \tilde{S}) - \frac{\partial}{\partial Z} (\tilde{M} \tilde{S}) - \sum_i \left( \frac{\partial M_i}{\partial Z} + \rho \frac{\partial \sigma_i}{\partial t} \right) S_{i_b} - LE + \tilde{Q}_R \quad : \text{Environment MSE}$$

$$\frac{\partial}{\partial t} \rho \sum \sigma_i S_i = -\frac{\partial}{\partial z} \left( \sum_i M_i S_i \right) + \sum_i \left( \frac{\partial M_i}{\partial z} + \rho \frac{\partial \sigma_i}{\partial t} \right) S_{i_b} + \sum_x (LC_i + Q_{Ri}) \quad : \text{in-cloud MSE}$$

$S_i : C_p T + gz$  (dry static energy) of  $i^{\text{th}}$  cloud

$S_{i_b} : C_p T + gz$  of the air entraining into or detraining from the  $i^{\text{th}}$  cloud

$C_i$  : condensation in the  $i^{\text{th}}$  cloud

$E$  : evaporation of liquid water in the environment

$Q_r$  : Radiative heating

- Entrainment :  $\frac{\partial M_i}{\partial z} + \rho \frac{\partial \sigma_i}{\partial t} > 0, \quad S_{i_b} = \tilde{S}$

- Detrainment :  $\frac{\partial M_i}{\partial z} + \rho \frac{\partial \sigma_i}{\partial t} < 0, \quad S_{i_b} = S_i$

#### iv) Approximation

- Assume  $\sigma_c \ll 1$  ,  $\bar{s} \approx \tilde{s}$  ( grid-mean, grid-resolvable = environment )

$$\frac{\partial}{\partial t} \rho \bar{s} = -\nabla \cdot (\rho \bar{v} \bar{s}) - \frac{\partial}{\partial z} (\rho \bar{w} \bar{s}) - \overline{\nabla \cdot (\rho \bar{v} s - \rho \bar{v} \bar{s})}$$

$$+ M_c \frac{\partial \bar{s}}{\partial z} - \sum_{dc} \left( \frac{\partial M_i}{\partial z} + \rho \frac{\partial \sigma_i}{\partial t} \right) (\delta_i - \bar{s}) - LE + \widetilde{\theta}_R$$

Detraining clouds      detrainment, entrainment

Adiabatic warming due to hypothetical subsidence between the clouds

$$\frac{\partial}{\partial t} \rho \bar{q} = -\nabla \cdot (\rho \bar{v} \bar{q}) - \frac{\partial}{\partial z} (\rho \bar{w} \bar{q}) - \overline{\nabla \cdot (\rho \bar{v} q - \rho \bar{v} \bar{q})}$$

$$+ M_c \frac{\partial \bar{q}}{\partial z} - \sum_{dc} \left( \frac{\partial M_i}{\partial z} + \rho \frac{\partial \sigma_i}{\partial t} \right) (q_i - \bar{q}) - E$$

- Spectral cloud ensemble :

$$M_c(z) = \int_0^{\lambda_{\max}} \underbrace{m(z, \lambda) d\lambda}_{\text{mass flux of between } \lambda \text{ and } d\lambda + \lambda} \quad \text{Sub-ensemble}$$

$$= \int_0^{\lambda_{\max}} \underbrace{m_B(\lambda)}_{\text{Mass flux at cloud base}} \eta(z, \lambda) d\lambda$$

$$\eta(z, \lambda) \equiv \frac{m(z, \lambda)}{m_B(\lambda)} \quad ; \quad \text{normalized subensemble mass flux}$$

## v) Closure

$$\frac{\partial m(z, \lambda)}{\partial z} = \mu(z, \lambda) \eta(z, \lambda)$$

$$\eta(z, \lambda) = e^{\lambda(z-z_B)} \quad ; \text{ mass flux profile}$$

- Cloud work function : measure of buoyancy

$$A(\lambda) = \int_{z_B}^{z_D(\lambda)} \eta(z, \lambda) g \frac{T_c(z, \lambda) - \bar{T}(z)}{\bar{T}} dz$$

- Q-G equilibrium

$$\frac{dA(\lambda)}{dt} = \underbrace{\frac{dA(\lambda)}{dt} \Big|_{LS}}_{\text{Large-scale forcing}} + \underbrace{\frac{dA(\lambda)}{dt} \Big|_C}_{\text{Adjustment}} \simeq 0$$

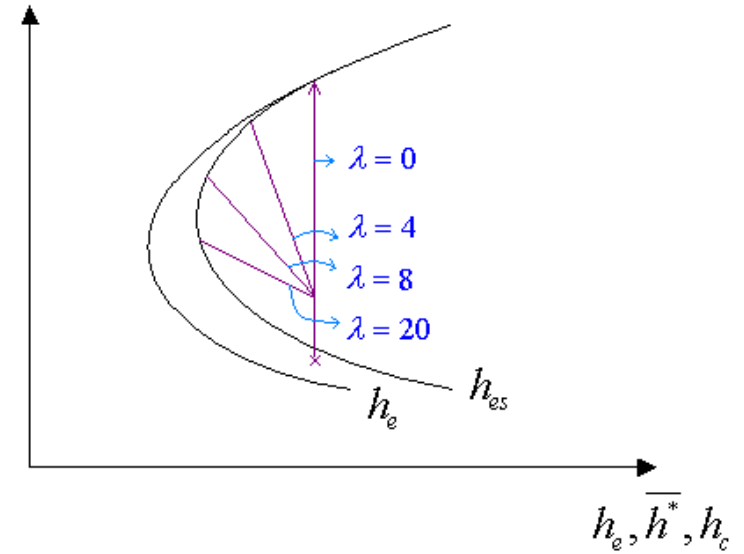
Large-scale forcing  
>0 : destabilized

Adjustment  
<0 : stabilization

Kernel : Cloud scheme kinetic energy

$$K_{ij} = \frac{A'_i - A_i}{(m_B \Delta t)} \quad \sum_j^{i_{\max}} K_{ij} (m_B \Delta t)_j + F_i = 0 \Rightarrow m_B$$

—————> compute  $\frac{\partial \bar{s}}{\partial t}, \frac{\partial \bar{q}}{\partial t}$  with  $\eta, m_B$



## 5.5 Other schemes

### \* Arakawa-Schubert type mass flux schemes

Grell scheme (1993) : Updraft/downdraft couplet without lateral mixing to find the deepest cloud

Simplified AS (SAS, Han and Pan 2011): revised cloud physics from the Grell

Relaxed AS (RAS, Moorthi and Suarez 1992): linearized profile function

### \* Other mass flux schemes : Low-level control convective schemes (Stensrud 2007)

Kain and Fritsch (2004) : CAPE based sophisticated convective plume model

Emanuel (1991) : Stochastic mixing cloud model

Tiedtke (1989) : Large-scale moisture convergence (KUO) based mass flux (ECMWF IFS model)

Gregory-Rowntree (1990): Parcel buoyancy based turbulence in cloud model (UK model)



# 6. Shallow Convection

## 6.1 Concept

more vigorous **vertical mixing of q and T above the mixed layer top**. With the enhanced vertical eddy transport between LCL and inversion level, this process does not allow the excess moisture trapped near the surface in synoptically inactive regions (**non-precipitating convection**).

➔ Cooling and moistening above LCL and heating and warming below.



## 6.2 Classification

- Moist adjustment type : Betts and Miller (1993), Lock et al. (2000), Tiedtke (1983)
- Mass flux type : Kain (2004), Park and Bretherton (2009), Han and Pan (2011)

### Tiedtke (1983)

$$\frac{\partial T}{\partial t} = \frac{1}{\rho} \frac{\partial}{\partial z} \left( \rho K \left[ \frac{\partial T}{\partial z} + \Gamma \right] \right)$$

$$\frac{\partial q}{\partial t} = \frac{1}{\rho} \frac{\partial}{\partial z} \left( \rho K \frac{\partial q}{\partial z} \right)$$

### Han and Pan (2011)

$$\frac{1}{\eta} \frac{\partial \eta}{\partial z} = \varepsilon - \delta$$

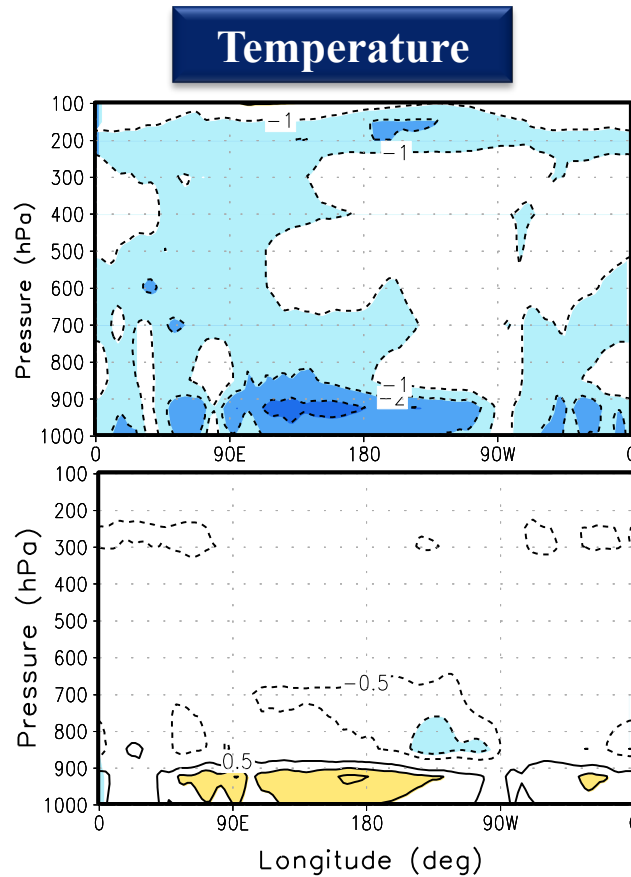
$$\frac{\partial(\eta s)}{\partial z} = (\varepsilon \bar{s} - \delta s) \eta$$

$$\frac{\partial[\eta(q_v + q_l)]}{\partial z} = \eta[\varepsilon \bar{q}_v - \delta(q_v + q_l) - r]$$

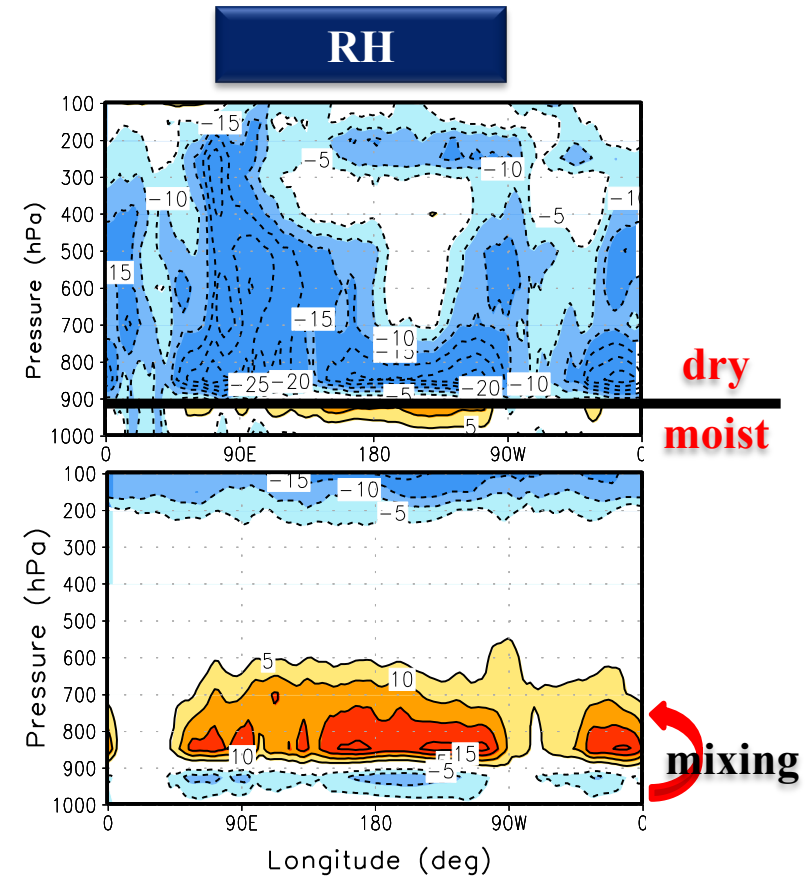
## 6.3 Impact of the shallow convection scheme (SCV)

: JJA 1996 simulation in a GCM

NO SCV – RA2



SCV – NO



\* SCV plays crucial role over the oceans.

# 7. Microphysics scheme

large-scale precipitation  
grid-resolvable scale precipitation  
explicit moisture scheme  
cloud scheme  
non-convective precipitation scheme

## 7.1 Concept

- Remove supersaturation after deep and shallow convection, and feedback to large-scale

## 7.2 Classification : according to the complexity in microphysics

- Diagnostic** : condensation, evaporation of falling precipitation
- Bulk microphysics** : hydrometeors with size distribution in inverse-exponential function
  - Single moment : predict mixing ratios of hydrometeors
  - Double moment : + number concentrations
  - Triple moment : + reflectivity
- Bin microphysics** : divides the particle distribution into a number of finite size or mass categories.

## 7.3 Precipitate size distributions

Marshall and Palmer(1948) : exponential law

Heymsfield and Platt (1984) : Power law

$$N_R(D_R) = aD_R^b$$

The rain and snow particles are assumed to follow the size distribution derived by Marshall and Palmer(1948), and Gunn and Marshall(1958), respectively. The size distributions for both rain and snow are formulated according to an inverse-exponential distribution and its formula for rain can be expressed by

$$N_R(D_R) = N_{0R} \exp(-\lambda_R D_R)$$

for rain, where  $N_{0R}$  is the intercept parameter of the rain distributions. Slope parameter is

$$\lambda_R = \left( \frac{\pi \rho_w N_{0R}}{\rho q_R} \right)^{1/4}$$

Due to the size distribution in exponential manner (integration of precip for whole size results in a constant), we can apply the bulk property microphysics terms.

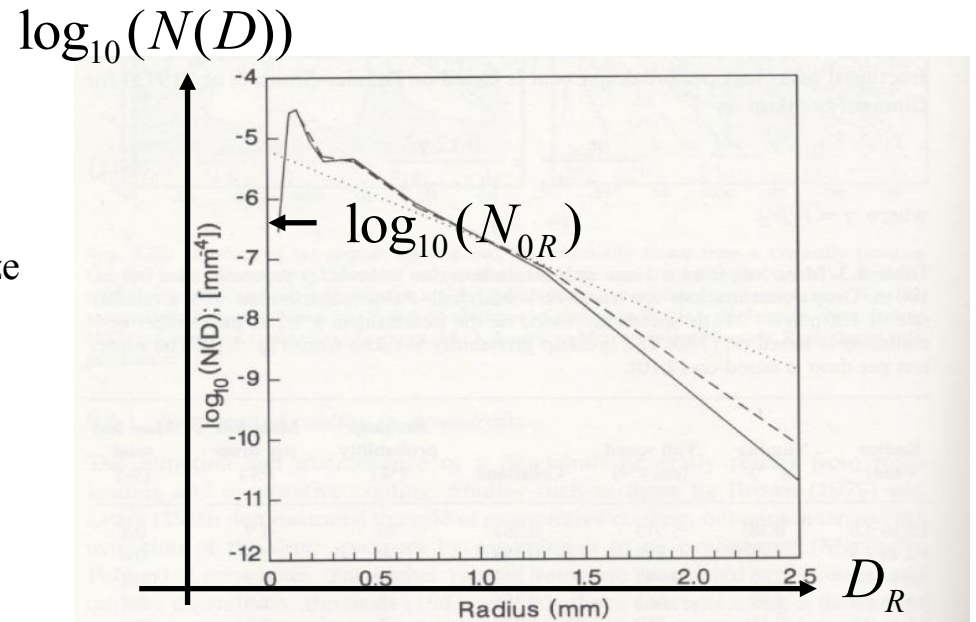


FIG. 9.21. Evolution of a drop spectrum in a subsaturated rainshaft including the effects of coalescence and breakup. The spectrum at the top of the rainshaft is Marshall–Palmer with a rainfall rate of  $110 \text{ mm h}^{-1}$ . (.....), 0 km; (---), 1 km; (—), 2 km;  $R = 110 \text{ mm h}^{-1}$ ;  $t = 24 \text{ min}$ . From Tzivion et al. (1989). *Journal of Atmospheric Sciences*, **46**, 21. American Meteorological Society. Reproduced with permission

$$\begin{aligned} \Gamma(x) &= \int_0^{\infty} t^{x-1} \exp(-t) dt \\ &= \int_0^{\infty} D_R^{4-1} \exp(-\lambda_R D_R) dD_R \\ &= \Gamma(4) / \lambda_R^4 \end{aligned}$$

## 7.4 Bulk Method : 1-Moment versus 2-Moment

Mixing Ratio

(1-moment/ 2-moment scheme)

$$\left( \int \frac{dM(D_R)}{dt} dN_{DR} \right) / \rho = \frac{dq}{dt} (\text{kg kg}^{-1} \text{s}^{-1})$$

Number concentration

(2-moment scheme)

$$\left( \int \frac{d \text{Prob}(D_R)}{dt} dN_{DR} \right) = \frac{dN}{dt} (\text{m}^{-3} \text{s}^{-1})$$



Single moment scheme

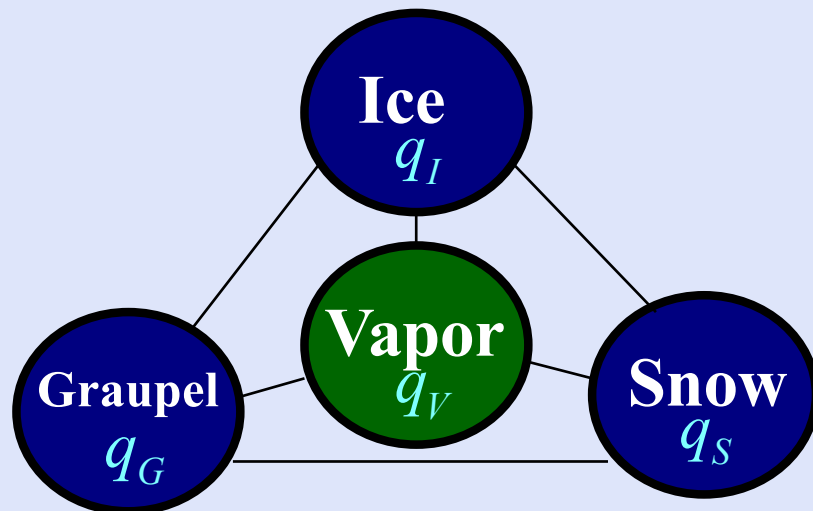
$$dN_{DR} = N_{0R} \exp(-\lambda_R D_R) dD_R$$

Double moment scheme

$$dN_{DR} = N_R \lambda_R^2 (N_R) D_R \exp(-\lambda_R D_R) dD_R$$

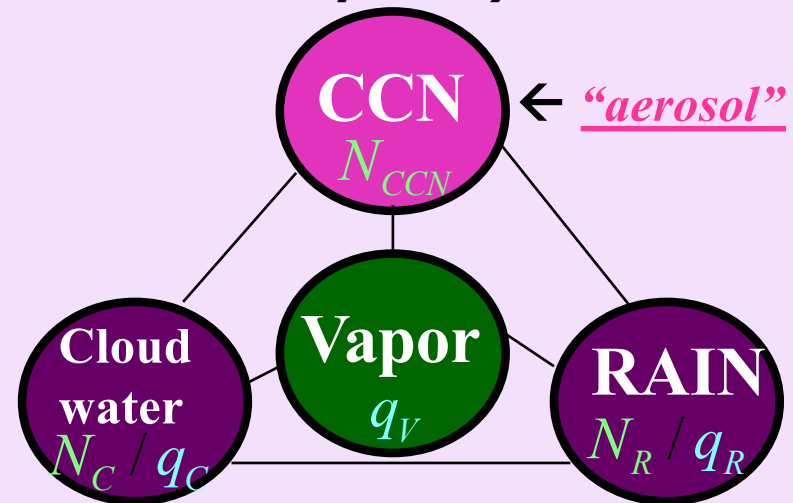
## 7.5 Bulk Method : 1-Moment (WSM) versus 2-Moment (WDM)

**Cold rain processes :**  
(Hong et al. 2004; Hong and Lim 2006)



$q$  for 4 hydrometeors will be predicted  
(Single Moment)

**Warm rain processes :**  
(Khairoutdinov and Kogan 2000;  
Cohardt and Pinty 2000)



$N, q$  for 2 hydrometeors will be predicted  
(Double Moment)

**N**: Cloud water, Rain, CCN

**q**: Cloud water, Rain, Ice,  
Snow, Graupel, Vapor

**WDM6**

(Lim and Hong, 2010)

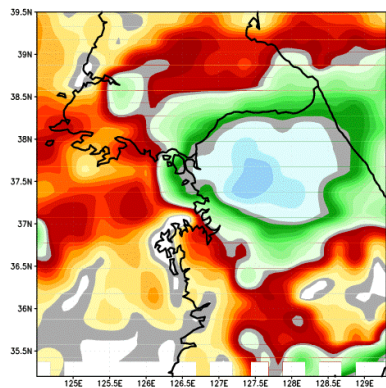


# Resolution Dependency

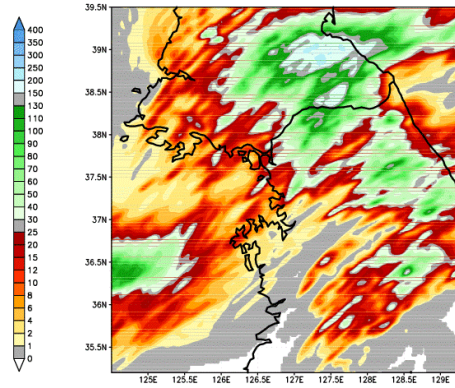
**Cut-off horizontal grid length:** current status in research communities

- PBL :  $\sim 50$  m (Mirocha, 2008 WRF workshop )
- GWDO :  $\sim 3$  or 1 km (hydrostatic approximation)
- GWDC:  $\sim 3$  or 1 km (go with CP)
- Cumulus parameterization :  $\sim 3$  or 1 km (cloud resolving scale)

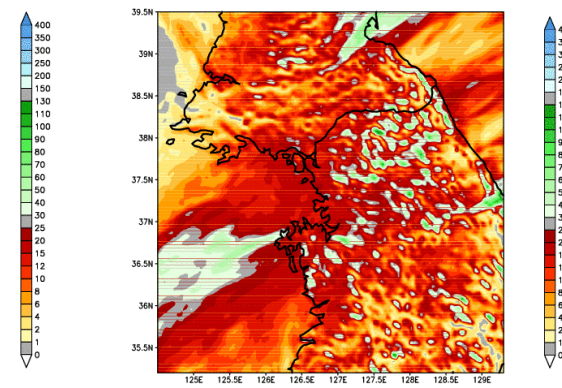
**CPS gray-zone issue : A heavy rainfall simulated by WRF at 3 km**



**OBS**



**No CPS**



**With CPS**

**Gray-zone** : partly resolved and partly parameterized (Hong and Dudhia 2012)

- CPS (1 km $\sim$ 10 km) : Gerard, Grell and Freitas, Arakawa and Ming, Pan and Han, Kwon and Hong
- PBL (100 m $\sim$ 1 km) : Honnert, Boutle, Shin and Hong
- Other processes such as shallow convection may also consider gray-zone

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# Thanks for your attention !

Modeling is to understand what is  
happening in nature !