

**Earth System Research Laboratory SCIENCE, SERVICE & STEWARDSHIP** 

New developments in RAP/HRRR physical parameterizations: MYNN-EDMF and mixing length revision

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# **NWP Development at NOAA-ESRL**



**RAP/HRRR** physics

# **Motivation/Outline**

# 1. RAP/HRRR bias that is clearly related to PBL scheme.

- High wind speed bias in PBL.
  - Rawinsonde, aircraft, tower data

#### 2. Mixing length revision.

- Update length scales & blending procedure.
- Make z-less.
- Scale-aware (Ito et al. 2015, *BLM* and Honnert et al. 2011, *JAS*) – work for 13, 3, and 0.75 km grid spacing.
- Cloud-specific mixing length.

# 3. Eddy Diffusivity-Mass Flux (EDMF)

- Improve representation of nonlocal mixing.
- Scale-aware (Honnert et al. 2011, JAS).
- Investigate momentum transport.





#### RAP/HRRR physics

### **Original MYNN Mixing Length Formulation**

The mixing length is designed such that the shortest length scale among,  $l_s$ ,  $l_t$ , and  $l_b$  will dominate:

$$\frac{1}{l_m} = \frac{1}{l_s} + \frac{1}{l_t} + \frac{1}{l_b}$$

where the surface layer length scale  $l_s$  is a function of the stability parameter( $\zeta = z/L$ ; L in the M-O length):

$$l_{s} = \begin{cases} kz(1 + \cos \zeta)^{-1} & \text{if } 0 \le \zeta \le 1\\ kz(1 - 100\zeta)^{0.2} & \text{if } \zeta < 0 \end{cases}$$

and the turbulent length scale  $l_t$  is:

$$l_{t} = \alpha_{1} \frac{\int_{z=0}^{PBLH} zq \, dz}{\int_{z=0}^{PBLH} q \, dz}$$

and the buoyancy length scale  $l_b$  is:

$$l_b = \alpha_2 \frac{q}{N} \left[ 1 + \alpha_3 \left( \frac{q_c}{l_t N} \right)^{1/2} \right]$$

where  $q_c$  is a turbulent velocity scale  $\sim O(w_*)$ 



#### **Problems associated with this Formulation**

$$\frac{1}{l_m} = \frac{1}{l_s} + \frac{1}{l_t} + \frac{1}{l_b}$$

- 1. "Harmonic" averaging:
  - a) The averaged mixing length is typically 20-40% smaller than the smallest length scale. This makes it very difficult to specify an exact mixing length needed in a given regime/part of atmosphere.
  - b) Never completely z-less if a z-dependent length scale can significantly impact the averaged mixing length.
- 2. Numerical noise can arise from buoyancy enhancement factor in  $l_b$ .  $l_b = \alpha_2 \frac{q}{N} \left[ 1 + \alpha_3 \left( \frac{q_c}{l_t N} \right)^{1/2} \right]$ 3. Not scale-aware.
- **Stable Conditions** 1500 1200 Height (m) 900 600 300 20 40 60 80 **Unstable Conditions** 1500 1200 Height (m) 900 600 300 20 40 60 80 mixing lengths (m)

### **New MYNN Mixing Length Revision**

All mixing length scales are defined to function specifically for their purpose. 1. The surface layer length scale  $l_s$  is defined the same way:

$$l_{s} = \begin{cases} kz(1 + \cos \zeta)^{-1} & \text{if } 0 \le \zeta \le 1 \\ kz(1 - \alpha_{4}\zeta)^{0.2} & \text{if } \zeta < 0 \end{cases}$$

The turbulent length scale  $l_t$  is also defined the same way:

$$l_{t} = \alpha_{1} \frac{\int_{z=0}^{PBLH} zq \, dz}{\int_{z=0}^{PBLH} q \, dz}$$

but now, the buoyancy enhancement factor in the buoyancy length scale  $l_b$  is removed:  $l_b = \alpha_2 \frac{q}{N}$  where  $q = \sqrt{(2 \times \text{TKE})}$  and N is the Brunt-Vaisala frequency.



Add a cloud-specific length scale  $l_c$  if clouds exist in grid cell, following Teixeira and Cheinet (2003, *BLM*):

$$l_c = \tau \sqrt{TKE}$$
 where  $\tau = 325$  seconds.

In the free atmosphere, the "BouLac" length scale is retained (Bougeault and Lacarrere 1989, MWR).

### New MYNN Mixing Length Revision (continued)

2. Define stable and unstable mixing lengths, using blending of *no more than two length scales*.

$$l_{stable} = (1 - w)l_s + wl_b$$

 $l_{unstable} = \frac{l_s}{1 + \frac{l_s}{1}}$ 

where  $w = z/h_s$ ,  $h_s$  is height of surface layer (= 0.2×PBLH).

3. Then the minimum is taken to get a mixing length for the PBL:

$$l = MIN(l_{stable}, l_{unstable})$$

4. Blend the PBL mixing length with the free-atmospheric mixing length:

$$l = (1 - w)l + wl_{BouLac}$$
 where  $w = \tanh\left[\frac{z - 1.3 \times PBLH}{0.15 \times PBLH}\right]$ 

5. Subsequent adjustment of mixing length parameters:  $\alpha_2$  (for  $l_b$ ),  $\alpha_4$  (for  $l_s$ ), "*cns*" (for  $l_s$ ), and possibly  $\alpha_1$  (for  $l_t$ ).

#### New MYNN Mixing Length Revision (continued)

6. Add scale-aware functionality, following Ito et al. (2015, *BLM, accepted*), using the similarity functions of Honnert et al. (2011, *JAS*).

 $l = l \times P_{TKE}$ 

Where  $P_{TKE}$  is a function of the model grid spacing  $\Delta x$  and boundary layer height, PBLH.



Adapted from Honnert et al. (2011, JAS). Nondimensional similarity function for TKE within the boundary layer

#### **Results: Alleviating Noise from GABLS3 SCM**



Tests performed by Wayne Angevine (NOAA/ESRL/CSD). 01 July 2006 over Cabauw, Ned. Fully interactive radiation and LSM, advection terms are prescribed. Credit Hugo Hartmann also for making us aware of a different, but somewhat related, instability in the unstable regime (improved upon but not fully resolved).

### **Case Study: Strong LLJ**



**RAP/HRRR** Physics

## **Case Study Validation (Towers)**

**RAP-ML** revision

12hr forecasts



**RAP-Control** 

Model:

**RAP/HRRR** Physics

 The RAP-Control is very high-biased in strong LLJ conditions.

- The revised mixing lengths reduces the wind speed bias by ~0.5 m s-1 near hubheights.
- Very little difference during the day.

Validated against 28 towers in southern Great Plains, only using data at heights > 70 m.

# **Mean Profile Comparisons**

Mean Profiles over the LLJ region (Kansas) between 06-08 UTC 11 June 2015.

- The revised mixing length reduces the wind speeds by ~2 m s<sup>-1</sup> below the LJJ max.
- The revised mixing length reduces the LLJ max by ~1 m s<sup>-1</sup> and elevates it ~150 m



# **Eddy Diffusivity/Viscosity**

 $K_{\phi} = S_{\phi} (2^* T K E)^{1/2} I_m$ 

- The revised eddy diffusivities/viscosities are much larger (by 25-50%) at night (much smaller difference during the day – not shown).
- The %-difference of mixing length is largest below the height of the LLJ max.
- Double-maxima in mixing length profile seems more appropriate for the shear layers above/below the LLJ.



# **Model Validation: Rawindsonde**

#### 12 hr fcsts compared to soundings across E-CONUS, 00 and 12 Z between 08-15 June 2015



## **Model Validation: Towers (>70m)**

#### 12 hr fcsts compared to 37 towers across Midwest, between 08-15 June 2015



# **MYNN-EDMF**

Improving the non-local transport in the MYNN PBL scheme by adding a mass-flux component.

#### **Questions:**

- 1) Can adding non-local transport to a local scheme that is designed to be more diffusive (in order to compensate) improve forecast skill?
- 2) What components of the mass-flux scheme are necessary to best fit in the RAP/HRRR framework (multi-parcel, stochastic entrainment/ detrainment rates, momentum transport, ensemble of closures, etc.)?

Plan to incorporate 3 different mass-flux schemes into MYNN and determine the best combination of features (not a bakeoff of individual mass-flux schemes):

- 1) Grell-Frietas-Olson scheme (NOAA-ERSL/GSD). Ensemble of closures, partially scale-aware.
- 2) TEMF (Wayne Angevine, NOAA-ESRL/CSD). Momentum transport, most tested, simplest.
- 3) StEM (Kay Suselj, Joao Teixeira, NASA-JPL). Multi-parcel, stochastic, with momentum transport, and partially scale-aware.

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### **MYNN-EDMF**

Preliminary SCM results from ARM case (21 June 2006), testing MYNN-EDMF with NASA-JPL's StEM scheme with momentum transport activated.



Credit Wayne Angevine (NOAA-ESRL/CSD) for running the SCM tests, Kay Suselj (NASA JPL) for providing mass-flux code, and Georgios Matheou (NASA JPL) for LES output.

### Summary

- Mixing length revision improves RAP/HRRR biases.
  - Reduces high nighttime wind speed bias in PBL.
  - Improves high daytime temperature bias in PBL.
  - Reduces noise found in idealized SCM case.
- New mixing length parameter estimation in progress.
  - WFIP1 & WFIP2 data will be used to determine final configuration.
- Promising SCM test results for MYNN-EDMF.
  - Future work will test various components of different mass-flux schemes in an attempt to develop a "best fit" mass-flux companion for MYNN.